



Electron Holes in a Regularized Kappa Background

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Abstract. The pseudopotential method is used to derive electron hole structures in a suprathermal plasma having a regularized κ probability distribution function background. The regularized character allows the exploration of small κ values beyond the standard suprathermal case, for which $\kappa > 3/2$ is a necessary condition. We have found the nonlinear dispersion relation yielding the amplitude of the electrostatic potential in terms of the remaining parameters, in particular the drift velocity, the wavenumber and the spectral index. Periodic, solitary wave, drifting and non-drifting solutions have been identified. In the linear limit, the dispersion relation yields generalized Langmuir and electron acoustic plasma modes. Standard electron hole structures are regained in the $\kappa \gg 1$ limit.

1 Introduction

The phenomenon of so-called electron holes in a plasma has received growing attention in the recent past specially due to recent spacecraft observation of such structures, see, e.g., (Steinvall, 2019a, b). In particular a recent study resolved the phase space density deficit of trapped electrons and proved that the solitary waves with bipolar profiles observed in space plasma are electron holes (Mozer, 2018). For the application to space plasmas a quantitative treatment of electron holes should take into account the presence of a suprathermal, i.e. non-Maxwellian background plasma. This was already pointed out in (Schamel, 2015, 2023) and carried out in (Haas, 2021; Aravindakshan, 2018, 2020; Jenab, 2021). In (Haas, 2021) the Maxwellian description of the trapped (hole) and untrapped (background) electron populations was substituted by one with a so-called standard kappa distribution (SKD).

The SKD is a simple generalization of a Maxwellian that was originally introduced by (Olbert, 1968) to describe non-Maxwellian power-law distributions of suprathermal plasma species, which are frequently observed in space. Since then the SKD has been applied successfully to numerous space plasma and laboratory scenarios. Along with these successes also various limitations of the SKD were identified: it exhibits diverging velocity moments, a positive lower limit of allowed kappa parameter values ($\kappa > 3/2$), and a non-extensive entropy (for a recent overview see (Lazar, 2021)). In addition, two types of SKDs were identified, namely the original one introduced by Olbert (Olbert, 1968) with a prescribed reference speed and a modified one that can be traced to Matsumoto (Matsumoto, 1972) with a temperature equal to that of the associated



Maxwellian, and it was demonstrated (Lazar , 2016) that care has to be taken in selecting one of those for a given physical
25 system. The kappa distribution was proposed in (Vasyliunas , 1968); extensive discussion on the different kappa distributions
can be found in (Pierrard , 2010; Hau , 2007).

All of these complications in employing the SKD can be avoided when one uses the *regularized kappa distribution* (RKD)
introduced non-relativistically in (Scherer , 2017) and for the relativistic case in (HanThanh , 2022). The RKD exhibits an
exponential cut-off of the power at high velocities. Such cut-off is a result of the fact that any acceleration process can only
30 occur on a finite spatial scale and a finite time scale. Consequently, such power law cannot extend to infinity (as in the case
of the standard kappa distribution) but must cut-off. The main purpose of the present work is to adopt a regularized version of
the SKD and to analyze the consequences. The RKD particularly removes all divergences in the theory and moves the lower
limit for the kappa parameter to zero (Scherer , 2019). Both features have consequences for correspondingly described physical
systems: in (Yoon , 2014) it was demonstrated that an ‘infrared catastrophe’ is avoided when using the RKD instead of the SKD
35 and in (Liu , 2020) it was shown that extending the range of kappa values to zero broadens the possible properties of solitary
ion acoustic waves in a plasma with RKD electrons. Here the reference value of κ is adopted according to Eq. (2) for the SKD.

Since also the first generalization of the analytical treatment of electron holes in an equilibrium plasma to a suprathermal
plasma was achieved by employing the SKD (Haas , 2021), the same constraints remain: not all moments of the velocity
distribution functions exist and kappa has to be greater than 3/2, thereby potentially preventing the study of a physically
40 interesting regime because harder velocity distributions are observed, see, e.g. (Gloeckler , 2012; Pierrard , 2022) and were
associated with observations of various solitary waves (Vasko , 2017). Therefore, the present work revisits the quantitative
treatment of electron holes in a suprathermal plasma, where the electron velocity distribution is described with the RKD.

The structure of the paper is as follows: in section II the one-dimensional RKD is defined, in section III various dimensionless
variables are introduced, in section IV the method of the pseudopotential is applied and in section V special solutions of the
45 resulting Poisson equation are derived. After an analysis of the corresponding dispersion relation in section VI for homogeneous
trapped electrons distributions, the final section VII contains the conclusions of the study.

2 One-dimensional regularized κ distribution

The starting point (Scherer , 2019; Liu , 2020) is the three-dimensional isotropic regularized *kappa* distribution (RKD),

$$f_3(\mathbf{u}) = \frac{n_0}{(\pi\kappa\theta^2)^{3/2} U\left(\frac{3}{2}, \frac{3}{2} - \kappa, \alpha^2\kappa\right)} \left(1 + \frac{u^2}{\kappa\theta^2}\right)^{-\kappa-1} \exp\left(-\frac{\alpha^2 u^2}{\theta^2}\right), \quad (1)$$

50 where n_0 is the equilibrium electrons number density, $\kappa > 0$ is the spectral index, θ is a reference speed, U is a Kummer
function of the second kind (or Tricomi function) described in (Scherer , 2019; Liu , 2020), \mathbf{u} is the velocity vector with
 $u = |\mathbf{u}|$ and $\alpha \geq 0$ is the cutoff parameter.

In the non-regularized limit $\alpha \rightarrow 0$ one regains the SKD

$$f_3(\mathbf{u}) = \frac{n_0 \Gamma(\kappa + 1)}{(\pi\kappa\theta^2)^{3/2} \Gamma\left(\kappa - \frac{1}{2}\right)} \left(1 + \frac{u^2}{\kappa\theta^2}\right)^{-\kappa-1}, \quad \alpha \rightarrow 0, \quad (2)$$



55 where Γ is the gamma function, which is positive defined provided $\kappa > 1/2$. For the RKD this constraint is not imposed on $\kappa > 0$.

For the treatment of electrostatic structures it is convenient to define the one-dimensional RKD. For this purpose we use cylindrical coordinates in velocity space and write $u^2 = v^2 + w^2$, where v is the component of the velocity parallel to the electric field and \mathbf{w} contains only the perpendicular velocity components, with $w = |\mathbf{w}|$. In the isotropic case the one-dimensional RKD
 60 is

$$f(v) = 2\pi \int_0^\infty dw w f_3(\mathbf{u})$$

$$= \frac{n_0 (\alpha^2 \kappa)^\kappa e^{\alpha^2 \kappa}}{(\pi \kappa \theta^2)^{1/2} U\left(\frac{3}{2}, \frac{3}{2} - \kappa, \alpha^2 \kappa\right)} \Gamma\left[-\kappa, \alpha^2 \kappa \left(1 + \frac{v^2}{\kappa \theta^2}\right)\right], \quad (3)$$

where here Γ is the incomplete gamma function of the indicated arguments. In other words, $f(v)$ comes from the three-dimensional version after integration over the two perpendicular velocity components.

65 In the non-regularized limit $\alpha \rightarrow 0$ one regains the standard one-dimensional κ distribution (Summers, 1991; Podesta, 2005)

$$f(v) = \frac{n_0 \Gamma(\kappa)}{(\pi \kappa \theta^2)^{1/2} \Gamma\left(\kappa - \frac{1}{2}\right)} \left(1 + \frac{v^2}{\kappa \theta^2}\right)^{-\kappa}, \quad \alpha \rightarrow 0, \quad (4)$$

which is positive definite provided $\kappa > 3/2$.

In the treatment of electrostatic structures, to satisfy Vlasov's equation the distribution function is a function of the constants
 70 of motion. In the one-dimensional, time-independent case, the available constants of motion are given by

$$\epsilon = \frac{mv^2}{2} - e\phi, \quad \sigma = \text{sgn}(v), \quad (5)$$

where $\phi = \phi(x)$ is the scalar potential, where m is the electron mass and $-e$ is the electron charge. The sign of the velocity $\sigma = v/|v|$ is an additional constant of motion just in the case of untrapped particles. The energy variable ϵ can be used to distinguish untrapped ($\epsilon > 0$) and trapped ($\epsilon < 0$) electrons.

75 In analogy with (Schamel, 1972, 2015, 2023) (where the background is not in the RKD form), presently one starts from Eq. (3) making for the untrapped part the replacement $v \rightarrow \sigma \sqrt{2\epsilon/m} + v_0$, where v_0 is a drift velocity, defining the distributions of untrapped and trapped electrons according to

$$f = f(\epsilon, \sigma) = \frac{A n_0}{\theta} \left(1 + \frac{k_0^2 \Psi}{2}\right) \left[H(\epsilon) \Gamma\left(-\kappa, \alpha^2 \kappa \left(1 + \frac{1}{\kappa \theta^2} (\sigma \sqrt{2\epsilon/m} + v_0)^2\right)\right) \right. \\ \left. + H(-\epsilon) \Gamma\left(-\kappa, \alpha^2 \kappa \left(1 + \frac{v_0^2}{\kappa \theta^2}\right)\right) \left(1 - \frac{\beta \epsilon}{m \theta^2}\right) \right], \quad (6)$$

$$80 \quad A = \frac{(\alpha^2 \kappa)^\kappa e^{\alpha^2 \kappa}}{(\pi \kappa)^{1/2} U\left(\frac{3}{2}, \frac{3}{2} - \kappa, \alpha^2 \kappa\right)}, \quad (7)$$

where $H(\epsilon)$ is the Heaviside function. The quantities k_0 and Ψ are dimensionless variables proportional respectively to the wavenumber of periodic oscillations and to the electrostatic field amplitude, as will be qualified in the following. In addition,



β is a dimensionless quantity associated to the inverse temperature of the trapped electrons distribution. Unlike singular distributions as in (Schamel, 2015, 2023; Haas, 2021; Schamel, 2018) here the velocity shifted hole distribution is assumed continuous at the separatrix ($\epsilon = 0$) and an analytic function of the energy for both trapped and untrapped electrons. These choices have been made in order to focus on the role of the cutoff parameter α instead of further aspects.

In the non-regularized case, using

$$(\alpha^2 \kappa)^\kappa \Gamma(-\kappa, \alpha^2 \kappa s) \rightarrow \frac{s^{-\kappa}}{\kappa}, \quad \alpha \rightarrow 0, \quad \kappa > 0, \quad (8)$$

for a generic argument s , and

$$90 \quad U\left(\frac{3}{2}, \frac{3}{2} - \kappa, \alpha^2 \kappa\right) \rightarrow \frac{\Gamma(\kappa - 1/2)}{\Gamma(\kappa + 1)}, \quad \alpha \rightarrow 0, \quad \kappa > 1/2, \quad (9)$$

from Eq. (6) one obtains

$$f = \frac{n_0(1 + k_0^2 \Psi/2)}{(\pi \kappa \theta^2)^{1/2}} \frac{\Gamma(\kappa)}{\Gamma(\kappa - 1/2)} \left[H(\epsilon) \left(1 + \frac{1}{\kappa \theta^2} (\sigma \sqrt{\frac{2\epsilon}{m}} + v_0)^2 \right)^{-\kappa} + H(-\epsilon) \left(1 + \frac{v_0^2}{\kappa \theta^2} \right)^{-\kappa} \left(1 - \frac{\beta \epsilon}{m \theta^2} \right) \right], \quad (10)$$

which is the κ version of Schamel's distribution that is given in its original form, e.g., in Eq. (4) in (Schamel, 1986) and illustrated in Fig. 1. A slight difference in comparison to the original formulation (Schamel, 1986, 2012) is that here the trapped electrons are described by a linear function of the energy instead of a Maxwellian function.

Finally, the Poisson equation

$$\frac{\partial^2 \phi}{\partial x^2} = \frac{e}{\epsilon_0} (n - n_0), \quad n = n(\phi) = \int_{-\infty}^{\infty} dv f(\epsilon, \sigma) \quad (11)$$

is needed, where ϵ_0 is the vacuum permittivity. An uniform ionic background n_0 has been assumed.

100 3 Dimensionless variables

To avoid the use of a large number of parameters, it is convenient to adopt dimensionless variables. For the RKD, it comes the question on which will be the reference speed defining the velocity rescaling. It would be tempting to consider the use of a thermal speed v_T defined in terms of the averaged squared velocity, but it is a cumbersome expression containing Kummer functions,

$$105 \quad v_T^2 = \frac{\langle u^2 \rangle}{3} = \frac{1}{3} \frac{\int d^3 u u^2 f_3(\mathbf{u})}{\int d^3 u f_3(\mathbf{u})} = \frac{\kappa \theta^2}{2} \frac{U\left(\frac{5}{2}, \frac{5}{2} - \kappa, \alpha^2 \kappa\right)}{U\left(\frac{3}{2}, \frac{3}{2} - \kappa, \alpha^2 \kappa\right)}, \quad (12)$$

the factor $1/3$ introduced to comply with the one-dimensional geometry. Therefore, for the sake of simplicity, instead of the thermal speed it is indicated to consider θ as the reference speed. In this way, the rescaled variables are

$$\tilde{x} = \frac{x}{\lambda}, \quad \tilde{v} = \frac{v}{\theta}, \quad \tilde{v}_0 = \frac{v_0}{\theta}, \quad \tilde{\phi} = \frac{e\phi}{m\theta^2}, \quad \tilde{n} = \frac{n}{n_0}, \quad \tilde{f} = \frac{f}{n_0/\theta}, \quad \tilde{\epsilon} = \frac{\epsilon}{m\theta^2}, \quad (13)$$

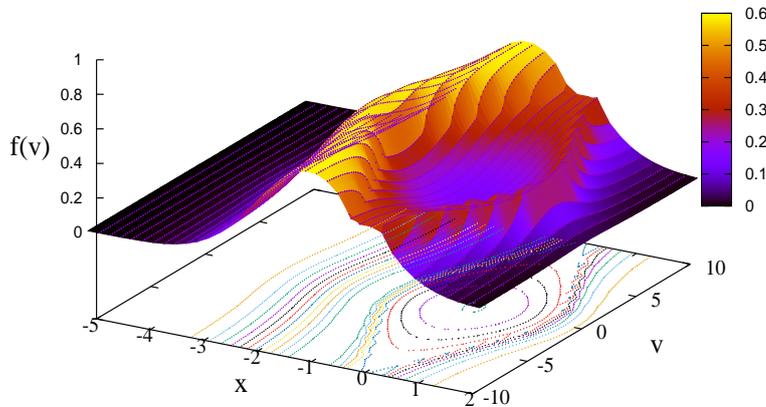


Figure 1. An illustration of the Schamel distribution (in arbitrary units and for the sech-potential in Eq.(13) in (Schamel , 1986)) for the values $\beta = -0.9, k_0 = 1.0, \psi = 1.0, \kappa = 0.5, \gamma = 0.1$ of the parameters used for the notation in (Haas , 2021).

where $\lambda = [\epsilon_0 m \theta^2 / (n_0 e^2)]^{1/2}$ is a modified Debye length.

110 As discussed in (Lazar , 2016) in the non-regularized context, our standard choice of θ as a κ -independent parameter better fits a scenario with enhanced tail in velocity space. Alternatively one could choose v_T from Eq. (12) to be κ -independent, which would be adequate for an enhanced core.

In dimensionless variables omitting for simplicity the tildes, the one-dimensional hole RKD from Eq. (6) is

$$\begin{aligned}
 f(\epsilon, \sigma) = A & \left(1 + \frac{k_0^2 \Psi}{2} \right) \left[H(\epsilon) \Gamma \left(-\kappa, \alpha^2 \kappa \left(1 + \frac{1}{\kappa} (\sigma \sqrt{2\epsilon} + v_0)^2 \right) \right) \right. \\
 115 & \left. + H(-\epsilon) \Gamma \left(-\kappa, \alpha^2 \kappa \left(1 + \frac{v_0^2}{\kappa} \right) \right) (1 - \beta \epsilon) \right], \tag{14}
 \end{aligned}$$

while Poisson's equation (11) is

$$\frac{\partial^2 \phi}{\partial x^2} = n - 1, \quad n = n(\phi) = \int_{-\infty}^{\infty} dv f(\epsilon, \sigma), \tag{15}$$

where $\epsilon = v^2/2 - \phi$ and $\sigma = \text{sgn}(v)$. In the remaining, the purpose is to evaluate the number density in Eq. (15) in terms of ϕ and to characterize the possible solutions of the Poisson's equation, specially regarding the behavior according to the

120 parameters κ, α .



4 Pseudopotential method

From Eqs. (14) and (15) one has

$$\begin{aligned} \frac{n}{A} &= \left(1 + \frac{k_0^2 \Psi}{2}\right) \left[\int_{-\infty}^{-\sqrt{2\phi}} dv \Gamma\left(-\kappa, \alpha^2 \kappa \left(1 + \frac{1}{\kappa}(\sqrt{2\epsilon} - v_0)^2\right)\right) + \right. \\ &+ \int_{\sqrt{2\phi}}^{\infty} dv \Gamma\left(-\kappa, \alpha^2 \kappa \left(1 + \frac{1}{\kappa}(\sqrt{2\epsilon} + v_0)^2\right)\right) + \\ 125 \quad &+ \left. \Gamma\left(-\kappa, \alpha^2 \kappa \left(1 + \frac{v_0^2}{\kappa}\right)\right) \int_{-\sqrt{2\phi}}^{\sqrt{2\phi}} dv (1 - \beta\epsilon) \right], \end{aligned} \quad (16)$$

assuming $0 \leq \phi \leq \Psi$, where Ψ denotes the peak-to-peak amplitude of the electrostatic potential, so that at $\phi = \Psi$ one has $d\phi/dx = 0$.

The integrals in Eq. (16) for the contribution of untrapped particles can be evaluated only in the weakly nonlinear limit. Expanding the integrands in a formal power series on $\sqrt{\phi}$ the result is

$$130 \quad n = 1 + \frac{k_0^2 \Psi}{2} + a\phi + b\phi\sqrt{\phi} + \mathcal{O}(\phi^2), \quad (17)$$

keeping the term proportional to Ψ as it has the same order of magnitude of ϕ where

$$\begin{aligned} a &= \frac{2}{\kappa U\left(\frac{3}{2}, \frac{3}{2} - \kappa, \alpha^2 \kappa\right)} \left[U\left(\frac{1}{2}, \frac{1}{2} - \kappa, \alpha^2 \kappa\right) + \right. \\ &+ \left. \frac{v_0}{\sqrt{\pi \kappa}} P \int_{-\infty}^{\infty} \frac{ds}{s - v_0} e^{-\alpha^2 s^2} \left(1 + \frac{s^2}{\kappa}\right)^{-\kappa-1} \right] \end{aligned} \quad (18)$$

where P stands for the principal value, and

$$\begin{aligned} 135 \quad b &= \frac{4\sqrt{2}}{3} \beta A \Gamma\left(-\kappa, \alpha^2 \kappa \left(1 + \frac{v_0^2}{\kappa}\right)\right) + \\ &+ \frac{8\sqrt{2}e^{-\alpha^2 v_0^2} [v_0^2 + 2\alpha^2 v_0^4 + \kappa(-1 + 2(1 + \alpha^2)v_0^2)]}{3\kappa^2 \sqrt{\pi \kappa} (1 + v_0^2/\kappa)^{\kappa+2} U\left(\frac{3}{2}, \frac{3}{2} - \kappa, \alpha^2 \kappa\right)}. \end{aligned} \quad (19)$$

It is possible to proceed in the same way to determine the average velocity $\langle v \rangle$ from

$$n\langle v \rangle = \int_{-\infty}^{\infty} dv v f(\epsilon, \sigma) \quad (20)$$

yielding

$$140 \quad \langle v \rangle = -v_0(1 - a\phi) + \mathcal{O}(\phi^{3/2}) \quad (21)$$

giving a more precise meaning of $-v_0$ which is the global drift velocity only in the limit of zero field amplitude. In addition notice the trapped electrons do not contribute to the average velocity, which comes from the untrapped part only, as found from the detail of the procedure similar to Eq. (16).



Poisson's equation (15) can be rewritten in terms of the pseudopotential $V = V(\phi)$,

$$145 \quad \frac{d^2\phi}{dx^2} = n - 1 = -\frac{\partial V}{\partial\phi}, \quad (22)$$

where

$$-V = \frac{k_0^2 \Psi \phi}{2} + \frac{a\phi^2}{2} + \frac{2b\phi^2\sqrt{\phi}}{5} + \mathcal{O}(\phi^3), \quad (23)$$

The case where the solutions are either periodic or solitary waves requires

1. $V(\phi) < 0$ in the interval $0 < \phi < \Psi$;
- 150 2. $V(\Psi) = 0$,

the latter implying

$$k_0^2 + a + \frac{4b\sqrt{\Psi}}{5} = 0, \quad (24)$$

which allows rewriting Eq. (23) as

$$-V = \frac{k_0^2\phi}{2}(\Psi - \phi) + \frac{2b\phi^2}{5}(\sqrt{\phi} - \sqrt{\Psi}), \quad (25)$$

155 up to $\mathcal{O}(\phi^3)$.

Equation (24) is the nonlinear dispersion relation (NDR) of the problem, providing a relation between phase velocity v_0 , wavenumber k_0 and amplitude proportional to Ψ , taking into account the expressions (18) and (19) for a, b . On the other hand, Eq. (22) can be integrated yielding

$$\frac{1}{2} \left(\frac{d\phi}{dx} \right)^2 + V(\phi) = 0, \quad (26)$$

160 where the integration constant was set to zero due to property (I) and since at the potential maximum $\phi = \Psi$ the electric field is zero. Following the usage from (Schamel, 2015, 2023; Haas, 2021; Schamel, 1972, 2018, 1986, 2012), the proposed *Ansatz* has tailored Ψ so that it is the root of $V(\phi)$ in Eq. (25). Otherwise, an irrelevant additive constant would be incorporated in the pseudopotential. The same applies to Eqs. (27) and (31) below.

5 Special solutions

165 5.1 Periodic solutions

As discussed in (Schamel, 2015, 2023; Haas, 2021; Schamel, 1972, 2018, 1986, 2012), the expansion of the number density in powers of $\sqrt{\phi}$ starting from an *Ansatz* such as in Eq. (14) can give periodic or localized solutions, according to specific conditions to be identified. For the sake of reference, we collect some of the known analytic solutions, remembering that of



course now the coefficients are adapted to the RKD equilibrium. For localized solutions as a by-product one has decaying
 170 boundary conditions.

The quadrature of Eq. (26) yields closed form solutions in special cases. In the linear limit, for a small amplitude so that $\sqrt{\Psi} \ll k_0^2/b$, neglecting the nonlinearity term $\sim b$, one has

$$V = \frac{k_0^2 \phi}{2} (\phi - \Psi). \quad (27)$$

Then from Eq. (26) immediately one has

$$175 \quad \phi = \frac{\Psi}{2} [1 + \cos(k_0(x - x_0))]. \quad (28)$$

Hence it is verified that k_0 indeed corresponds to the wavenumber of linear oscillations with $0 \leq \phi \leq \Psi$ in this case.

Assuming $k_0 \neq 0$, more insight is provided by the further rescaling

$$\bar{\phi} = \frac{\phi}{\Psi}, \quad \bar{x} = k_0 x, \quad \bar{V} = \bar{V}(\bar{\phi}) = \frac{V}{k_0^2 \Psi^2}, \quad \bar{b} = \frac{2b\sqrt{\Psi}}{5k_0^2} \quad (29)$$

reduces Eq. (26) to

$$180 \quad \frac{1}{2} \left(\frac{d\bar{\phi}}{d\bar{x}} \right)^2 + \bar{V}(\bar{\phi}) = 0, \quad (30)$$

where

$$\begin{aligned} -\bar{V}(\bar{\phi}) &= \frac{\bar{\phi}}{2} (1 - \bar{\phi}) + \bar{b} \bar{\phi}^2 \left(\sqrt{\bar{\phi}} - 1 \right) \\ &= \frac{\bar{\phi}}{2} \left(1 - \sqrt{\bar{\phi}} \right) \left(1 + \sqrt{\bar{\phi}} - 2\bar{b}\bar{\phi} \right), \end{aligned} \quad (31)$$

containing only one free parameter \bar{b} . The condition (II) for periodic or localized solutions amounts to $\bar{V}(\bar{\phi}) < 0$ within the
 185 interval $0 < \bar{\phi} < 1$. In view of the factorization in Eq. (31) it is easy to demonstrate the condition is always satisfied for
 $\bar{b} < 1$. The existence of periodic solutions such that $0 \leq \bar{\phi} \leq 1$ for $\bar{b} < 1$ comes from the shape of the rescaled pseudopotential
 shown in Figs. 2 and 3. The case $\bar{b} > 1$ also has periodic solutions, but with a smaller amplitude as apparent from Fig. 4. The
 physically meaningful solutions always occur for $\bar{V} < 0$ within the interval $0 < \bar{\phi} < 1$. Notice that with the further rescaling
 (29) the amplitude of oscillation is set to unity, as shown in the referred figures. The required weakly nonlinear analysis always
 190 supposes $\tilde{\phi} \sim \Psi \ll 1$ or, according to Eq. (13), $e\phi/(m\theta^2) \ll 1$, where ϕ is the physical scalar potential.

The exact quadrature of Eq. (30) with all terms has been fully discussed in (Schamel, 2012, 2000), where the pseudopotential
 is formally the same as in Eq. (31) after rescaling. It is given in terms of Jacobi elliptic functions showing a periodic behavior
 and higher order Fourier harmonics. The present work extends these results for the case of a background RKD, with the adapted
 coefficients.

195 It is apparent that the control parameter \bar{b} depending on several variables such as the effective trapped particles inverse
 temperature β determines the qualitative aspects of the oscillatory solutions. Figures 5 and 6 and 7 show in a different style
 how a smaller (and possibly negative) $\bar{b} < 1$ corresponds to a larger wavenumber, which is exactly k_0 only in the linear case.

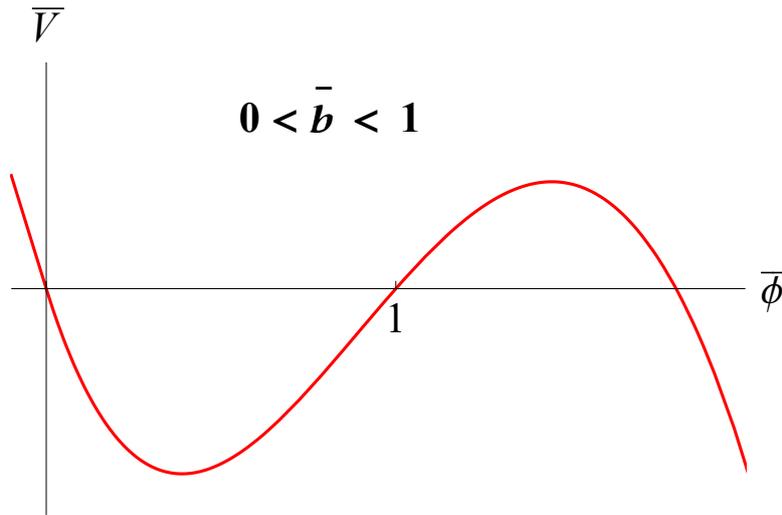


Figure 2. Rescaled pseudopotential from Eq. (31) for $0 < \bar{b} < 1$.

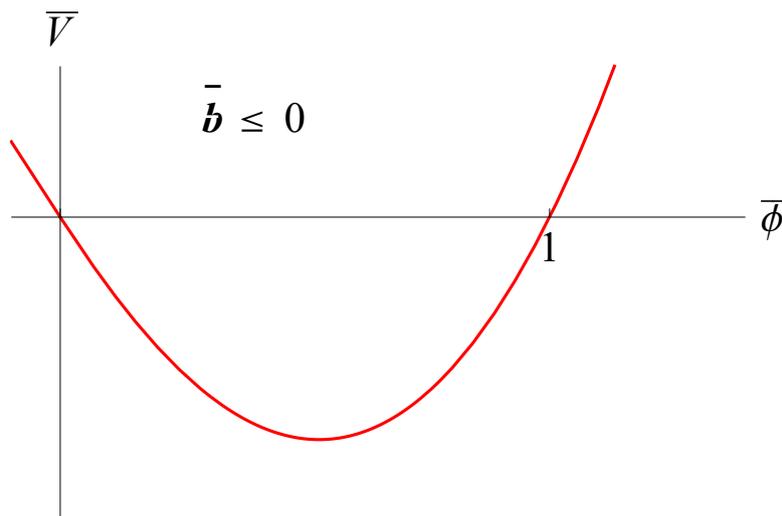


Figure 3. Rescaled pseudopotential from Eq. (31) for $\bar{b} \leq 0$.

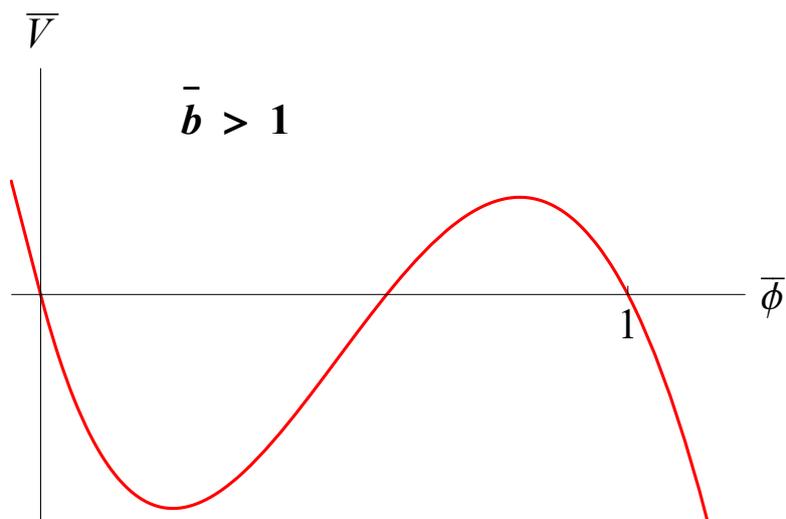


Figure 4. Rescaled pseudopotential from Eq. (31) for $\bar{b} > 1$. Periodic solutions exist in a smaller interval $0 \leq \bar{\phi} < 1$.

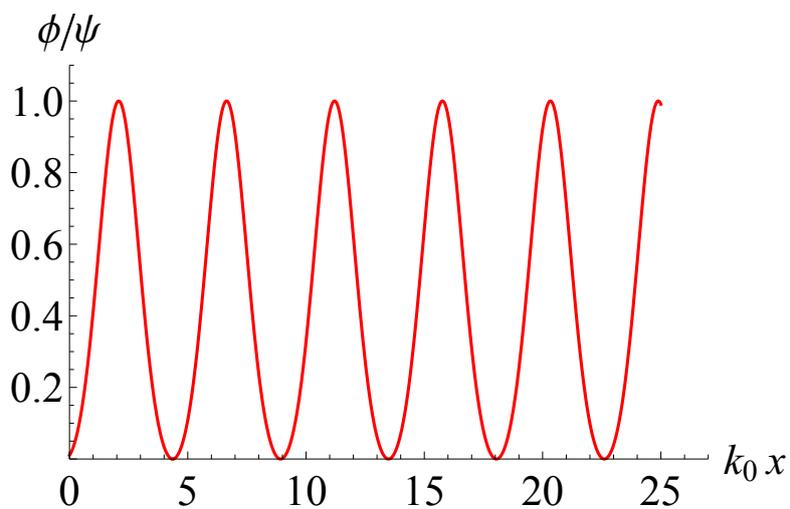


Figure 5. Numerical solution of Eq. (30) with $\bar{b} = -2, \bar{\phi}(0) = 10^{-3}$.

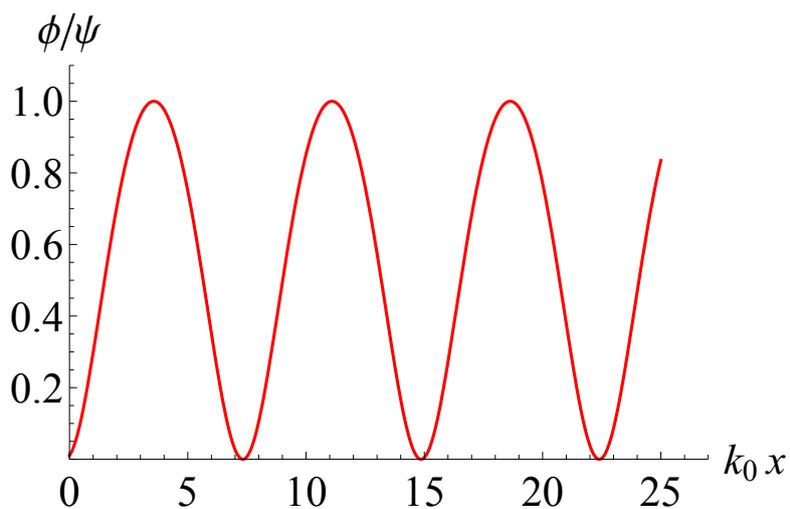


Figure 6. Numerical solution of Eq. (30) with $\bar{b} = 0.5, \bar{\phi}(0) = 10^{-3}$.

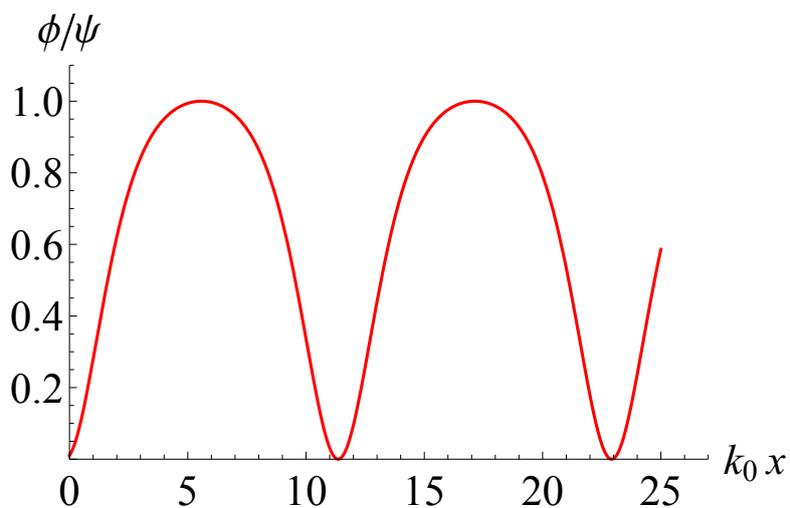


Figure 7. Numerical solution of Eq. (30) with $\bar{b} = 0.9, \bar{\phi}(0) = 10^{-3}$.

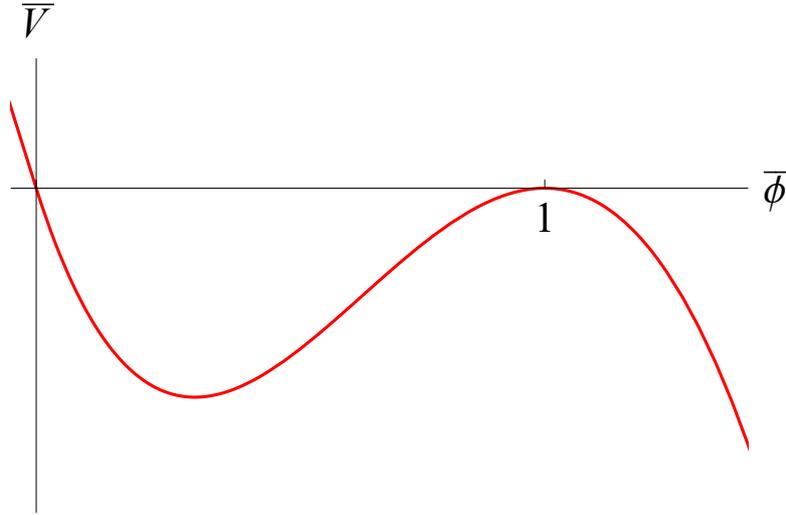


Figure 8. Rescaled pseudopotential from Eq. (31) for $\bar{b} = 1$.

5.2 Localized solution with $\bar{b} = 1, k_0 \neq 0$

The limit case $\bar{b} = 1$ with $k_0 \neq 0$ is special since then $d\bar{V}/d\bar{\phi} = 0$ at $\bar{\phi} = 1$, as shown in Fig. 8, yielding a localized, non-periodic
 200 solution. Moreover this case is amenable to the simple quadrature

$$\bar{\phi} = \frac{1}{4} \left[1 - 3 \tanh^2 \left(\frac{\sqrt{3}}{4} (\bar{x} - \bar{x}_0) \right) \right]^2, \quad (32)$$

see Fig. 9. The corresponding rescaled electric field is shown in Fig. 10. The total electrostatic energy is finite since the integral
 $(1/2) \int_{-\infty}^{\infty} d\bar{x} (d\bar{\phi}/d\bar{x})^2 = 6\sqrt{3}/35$ converges.

5.3 Solitary waves with $k_0 = 0$.

205 On the other hand if $k_0 = 0$ one has

$$V = \frac{2b\phi^2}{5} (\sqrt{\Psi} - \sqrt{\phi}), \quad (33)$$

yielding the solitary pulse

$$\phi = \Psi \operatorname{sech}^4 \left[\left(\frac{-b\sqrt{\Psi}}{20} \right)^{1/2} (x - x_0) \right], \quad (34)$$

which is well defined everywhere provided $b < 0$, which can be attainable e.g. for sufficiently small β, v_0^2 .

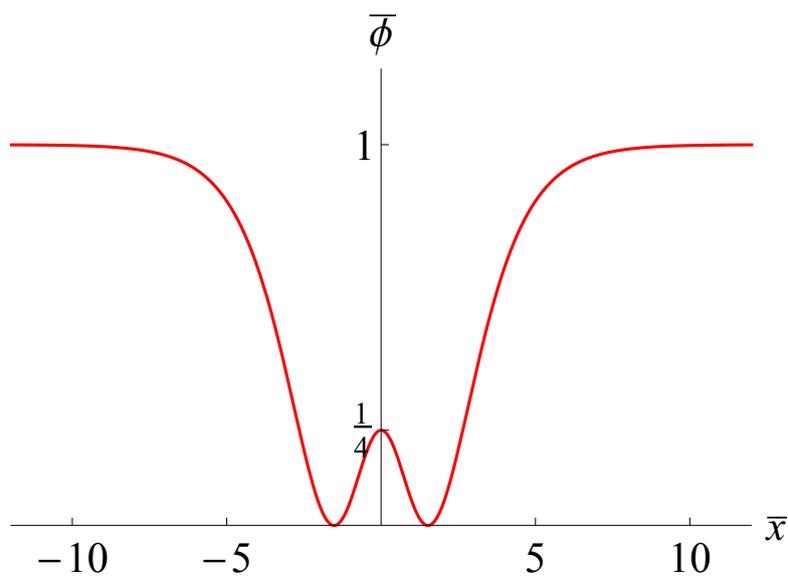


Figure 9. Rescaled electrostatic potential from Eq. (32) for $\bar{x}_0 = 0$.

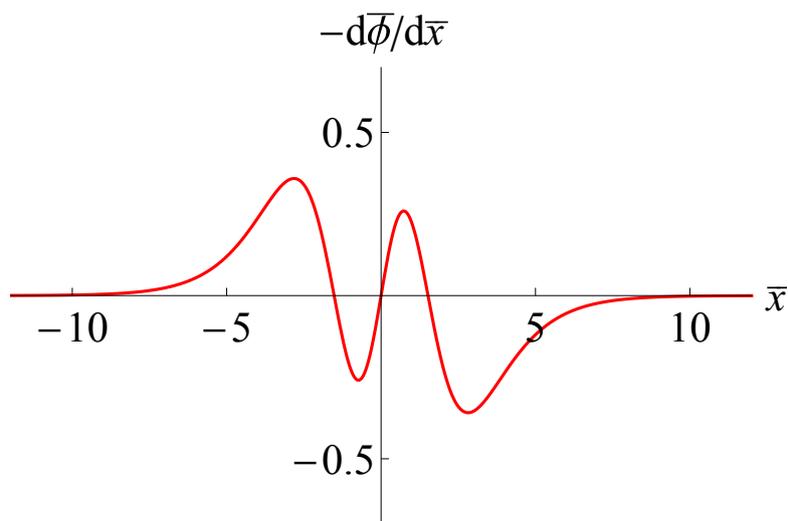


Figure 10. Rescaled electric field $-\frac{d\bar{\phi}}{d\bar{x}}$ where $\bar{\phi}$ is given in Eq. (32) for $\bar{x}_0 = 0$.



210 6 Dispersion relation

The NDR (24) provides several behaviors according to the values in parameter space. For the sake of simplicity it will be considered the case where the trapped particle distribution is homogeneous in phase space, which amounts to the dimensionless quantity $\beta = 0$ in Eq. (10). This is an increasingly better approximation for small enough amplitude so that $e\Psi \ll m\theta^2$, yielding a relatively smaller trapped area in phase space. Clearly this limit situation does not correspond to “holes”, since in
 215 this case the trapped particles are not in a depression in phase space as shown e.g. in Fig. 1. However, the analytic simplicity motivates the approach. Furthermore subcases can be identified: drifting, non-drifting; oscillating, non-oscillating, as follows. Our main purpose is to provide an investigation showing a regular behavior for small κ values, as long as $\alpha > 0$.

6.1 Non-drifting, non-oscillating

If the trapped distribution is homogeneous and non-drifting with respect to the fixed ionic background ($v_0 = 0$), one has from
 220 Eq. (24)

$$k_0^2 + \frac{2U\left(\frac{1}{2}, \frac{1}{2} - \kappa, \alpha^2\kappa\right)}{\kappa U\left(\frac{3}{2}, \frac{3}{2} - \kappa, \alpha^2\kappa\right)} - \frac{32\sqrt{2}\Psi}{15\kappa\sqrt{\pi\kappa}U\left(\frac{3}{2}, \frac{3}{2} - \kappa, \alpha^2\kappa\right)} = 0. \quad (35)$$

Furthermore in the non-oscillating case $k_0 = 0$ one can solve Eq. (35) as

$$\Psi = \frac{\pi}{2}\kappa \left[\frac{15}{16}U\left(\frac{1}{2}, \frac{1}{2} - \kappa, \alpha^2\kappa\right) \right]^2, \quad (36)$$

which is the amplitude of the solitary wave in terms of the remaining parameters κ, α only. Figure 11 shows the resulting
 225 amplitude. The regular behavior as $\kappa \rightarrow 0$ is apparent. A larger α implies a smaller solitary wave amplitude. In the non-regularized limit $\alpha \rightarrow 0$ it is possible to show that from Eq. (36) one has $\Psi \rightarrow 1.38$ as $\kappa \rightarrow \infty$, which is beyond the weakly nonlinear assumption. From Fig. 11 one also has that the $\alpha = 0$ case only admits small amplitude holes for $\kappa \ll 1$, which is in contradiction with the constraint $\kappa > 3/2$ for the non-regularized equilibrium. It is interesting to note that the weakly nonlinear condition $\Psi \ll 1$ is much better fulfilled for sufficiently high α . Hence, such hole structures (with $\beta = 0$, non-drifting and
 230 non-oscillating) are more reliable in a RKD background. Note, however, that high α values limit the extent of the power laws.

6.2 Non-drifting, oscillating

Allowing with $k_0 \neq 0$ for oscillating solutions one also has a regular behavior of the amplitude as $\kappa \ll 1$. In this limit, assuming $\alpha > 0$, it can be shown that Eq. (35) reduces to

$$k_0^2 + \frac{\sqrt{\pi}\alpha}{\sqrt{\kappa}} - \frac{32\alpha\sqrt{\Psi}}{15\sqrt{2\pi\kappa}} = 0, \quad \kappa \ll 1, \quad \alpha > 0 \quad (37)$$

235 yielding a vanishingly small amplitude as $\kappa \rightarrow 0$. Figure 12 shows Ψ from Eq. (35) as a function of κ , for $\alpha = 1.5$ and different k_0 values. It is found that a larger k_0 yields a larger amplitude.

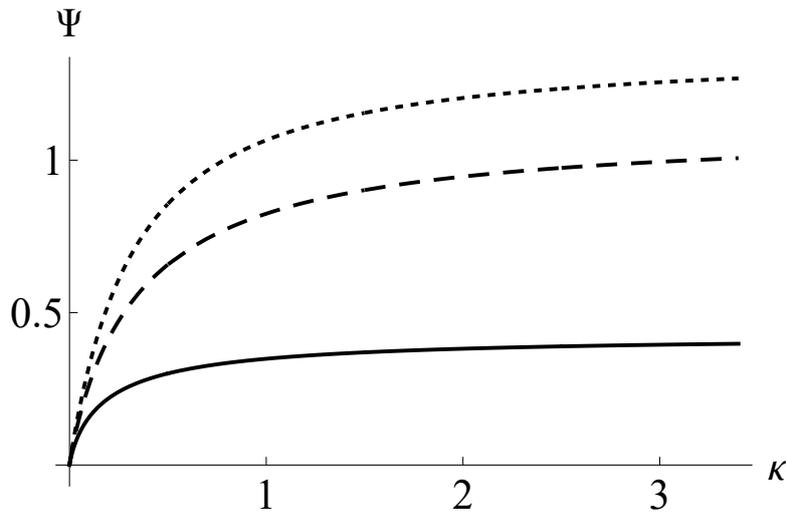


Figure 11. Solitary wave amplitude in the homogeneous trapped distribution, non-drifting and non-oscillating case as a function of κ and different α 's, from Eq. (36). Upper, dotted line: $\alpha = 0.0$; mid, dashed: $\alpha = 0.5$; Lower, solid: $\alpha = 1.5$.

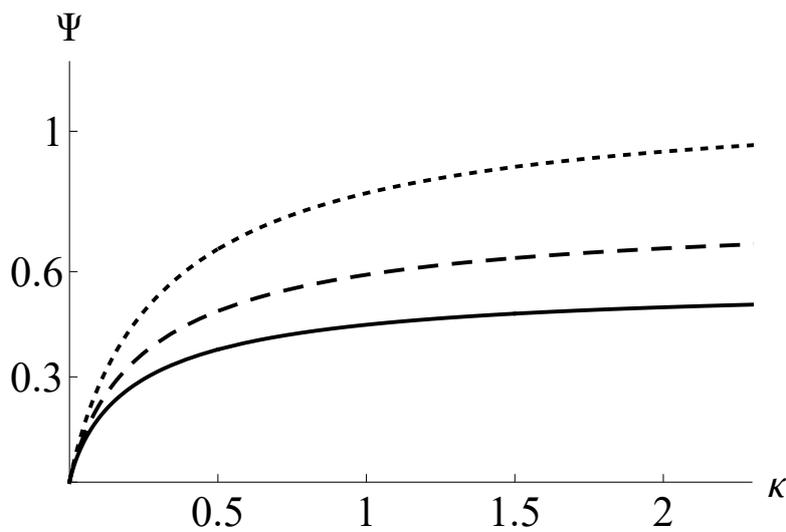


Figure 12. Wave amplitude in the homogeneous trapped distribution, non-drifting and oscillating case as a function of κ and different wavenumbers, for $\alpha = 1.5$, from Eq. (35). Lower, solid: $k_0 = 1.0$; mid, dashed: $k_0 = 1.5$; upper, dotted line: $k_0 = 2.0$.



6.3 Dispersion relation with $v_0 \neq 0$

Allowing for drifting structures so that $v_0 \neq 0$, for simplicity disregarding the nonlinear term $\sim b\sqrt{\Psi}$ and still with homogeneous trapped electrons distribution ($\beta = 0$), one has from Eq. (24),

$$\begin{aligned}
 240 \quad k_0^2 + \frac{2}{\kappa U \left(\frac{3}{2}, \frac{3}{2} - \kappa, \alpha^2 \kappa \right)} \left[U \left(\frac{1}{2}, \frac{1}{2} - \kappa, \alpha^2 \kappa \right) + \right. \\
 \left. + \frac{v_0}{\sqrt{\pi} \kappa} P \int_{-\infty}^{\infty} \frac{ds}{s - v_0} e^{-\alpha^2 s^2} \left(1 + \frac{s^2}{\kappa} \right)^{-\kappa - 1} \right] = 0. \quad (38)
 \end{aligned}$$

Setting $v_0 = \omega_0/k_0$, Eq. (38) produces similar thumb curves as for holes in a Maxwellian background (Schamel, 1986), now adapted for the RKD. Figure 13 show results for different small κ values, in all cases with $\alpha = 0.1$. As usual, one has a high frequency (Langmuir) mode together with a slow electron-acoustic mode (Fried, 1961) now adapted to the RKD background, where both modes coalesce in a certain point according to the parameters. As seen, the behavior is regular even for small κ values. At the extremal k value where both modes coalesce, apparently the group velocity is infinite. As discussed in (Schamel, 2013; Valentini, 2012), at this point taking into account the nonlinear trapping the phase velocity of the hole should replace the diverging linear group velocity.

7 Conclusions

250 For the first time, electron holes have been discussed in a RKD background. Unlike (Haas, 2021), for simplicity this time the background distribution function has no singular features. It was verified that the regularization of the SKD avoids all divergent properties of the solutions for a vanishingly small spectral index κ . In terms of the hole distribution function for both trapped and untrapped electrons, the number density has been evaluated yielding the pseudopotential in the weakly nonlinear limit. As a consequence, the most prominent solutions of the resulting Poisson equation have been found. Drifting, non-drifting, oscillating and non-oscillating solutions have been discussed. The linear dispersion relation has been also analyzed, yielding a κ -dependent plasma modes diagram showing a high frequency Langmuir mode and a low frequency electron acoustic mode (Fig. 13). Unlike for a SKD background, all findings remain regular even for very small κ values. The results are relevant especially for plasmas having a suprathermal equilibrium with small spectral κ index, for which the SKD is not appropriate, as frequently happens in space plasmas (Gloeckler, 2012; Pierrard, 2022).

260 *Data availability.* The data that support the findings of this study are available from the corresponding author upon reasonable request.

Author contributions. The authors contributed equally to this work, regarding original idea, basic theory and applications.

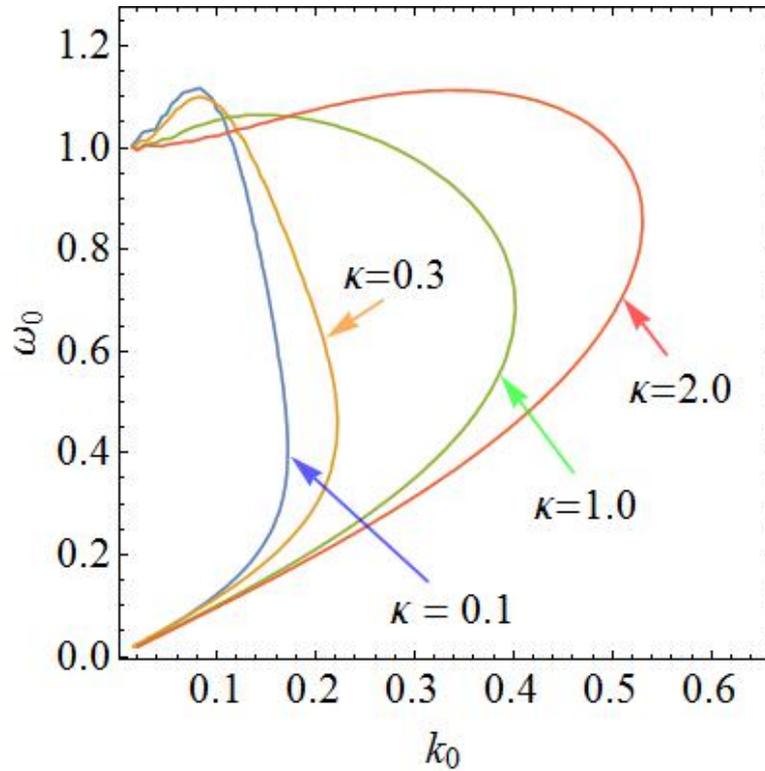


Figure 13. Dispersion relation (38) with $v_0 = \omega_0/k_0$ for $\alpha = 0.1$ and $\kappa = 0.1, 0.3, 1.0, 2.0$, as indicated.

Competing interests. The authors report no conflict of interest.

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