

1 **Brief Communication: Modified KdV equation for Rossby-Khantadze waves in a**
2 **sheared zonal flow of the ionospheric E-layer**

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15
16 **Abstract**

17
18 The system of nonlinear equations for Rossby-Khantadze waves in a weakly ionized
19 conductive ionospheric plasma with sheared zonal flow is given. Use of multiple-scale analysis
20 allows reduction of obtained set of equations to one-dimensional so called nonlinear modified
21 KdV (MKdV) equation describing the propagation of solitary Rossby-Khantadze waves.
22

23 **Keywords:** Rossby-Khantadze waves; nonlinearity, sheared zonal flow
24

25 **1. Introduction**
26

27 Different satellite and ground-based investigations indicate presence of zonal flows in
28 various atmospheric regions around the Earth (Pedlosky, 1987). The presence of sheared flow
29 along the meridians with inhomogeneous velocity, is connected with the ultra-low-frequency
30 perturbations in E and F regions of the ionosphere (Satoh, 2004; Champeaux et al. 2008; Shukla
31 et al., 2003; Onishchenko et al. 2004; Kaladze T.D., 2007; Kaladze et al., 2008). Effect of
32 sheared flow can be seen in the linear and nonlinear properties of the waves, and conditions
33 suitable for that are available in Earth's ionosphere. This gives rise to a variety of linear and
34 nonlinear phenomena like zonal flows. Rossby waves may exist and this may occur in the upper
35 atmosphere and in the oceans; and these play a significant role in the global atmospheric
36 circulation.

37 The reason for the existence of zonal flows is the non-uniform warming of the Earth's
38 atmospheric regions by the sun. In recent decades, several nonlinear phenomena related to the
39 excitation of sheared (zonal) flows were investigated. For instance, Benkadda et al., (2011)
40 investigated the excitation of zonal flows in fusion plasmas by taking drift waves into account.
41 Using Hasegawa Wakatani model, Champeaux et al. (2008) investigated the excitation of zonal
42 flows in drift wave turbulence. Lately, zonal flow generations for short wavelength
43 electromagnetic Rossby waves have also been studied by taking Reynolds stresses into account
44 (Shukla et al., 2003; Onishchenko et al. 2004).

45 The production of zonal flow in E-ionospheric layer by Rossby waves was investigated by
46 Kaladze et al. (2007). The authors considered the effects of the zonal flows on nonlinear
47 structures in Rossby waves and it was shown that such structures split into various segments
48 based on the collection of zonal flow energy (Kaladze et al., 2008). A different concept was
49 considered by Benkadda et al. (2011) where they emphasized on the interaction of high
50 frequency drift waves with those having lower frequencies.

51 More recently, it was seen, that propagation of coupled Rossby-Khantadze (RK) waves
52 could be self-organized into dipolar (solitary) vortices (Kaladze, 2014). The spatial
53 inhomogeneity of Coriolis parameter and ambient magnetic field along the meridians causes
54 the propagation of such coupled RK waves. The numerical simulation of the same problem
55 was carried out by Kaladze, et al. (1999) where the possibility of modulation mechanism in
56 ionospheric plasmas was discussed, in turn leading to the formation and transport of dust flows
57 and particles respectively. Numerical solutions of RK waves with sheared zonal flow in weakly
58 ionized ionospheric E-plasma region was investigated as well (Futatani et al., 2015). In his
59 work he has pointed out the splitting of vortices, where the energy is transported by sheared
60 flow into multiple pieces. The equatorial Rossby waves were discussed under similar sheared
61 flows in the beginning of the millennia (Qiang et al., 2001) and the existence of solitons was
62 confirmed by the observations of *Freja and Viking satellites* (Qiang et al., 2001; Bostrom,
63 1992; Lindqvist et al., 1994; Dovner et al. 1994, YunLong Shi et al. 2018). The generation of
64 shear flow by Rossby-Khantadze waves in E-ionospheric region has also been discussed
65 (Kaladze, et al. 2014). Earlier, Kaladze et al. (2009) had earlier investigated the properties of
66 magnetized solitary Rossby waves with the interaction of zonal flows (sheared) and developed
67 the mKdV equation. By considering the β -plane approximation, similar work was done by Jian
68 et al., (2009).

69 In the present paper, we have considered the effect of magnetic field for the MKdV
70 equation for RK waves. Such case has not been reported so far and thus provides novelty to
71 this work. Here we have investigated solitary Rossby-Khantadze (RK) waves by incorporating
72 sheared zonal flows in a partially ionized conducting plasma, found in the ionospheric E-
73 region. In Sec. 2, by using the multiple scale analysis and perturbation approach from a system
74 of nonlinear two-dimensional equations we derive one-dimensional MKdV equation
75 describing solitary Rossby-Khantadze waves' dynamics along with zonal (shear) flows. In Sec.
76 3, includes the discussion of the results.

77 78 79 **2. Mathematical Preliminaries**

80
81 We consider weakly ionized E-ionospheric region comprising of electrons, ions and
82 neutrals particles, where the ionospheric plasma is enclosed in a geomagnetic field $\mathbf{B}_0 =$
83 $(0, B_{0y}, B_{0z})$. Similarly, the angular velocity of the earth contains no x-component, $\mathbf{\Omega} =$
84 $(0, \Omega_{0y}, \Omega_{0z})$. In this layer therefore, two-dimensional consideration of the wave motion
85 provides complete information about its propagation in terms of stream function $\psi(x, y, t)$, $\mathbf{v} =$
86 $(u, v, 0)$, with $u = -\frac{\partial\psi}{\partial y}$ and $v = \frac{\partial\psi}{\partial x}$.

87
88 The nonlinear behavior of the sheared Rossby-Khantadze waves can be described by
89 the following system of 2D equations, where the first equation is obtained from the z-
90 component of the curl of vorticity with $\zeta_z = \Delta\psi$ and the second one is obtained from the z-
91 component of magnetic field, through Faraday's law. Here we obtained the following system
92 of Eqs. (1) under the assumption that electron and ion flows due to the small concentration
93 number (compared to the neutral particles) gives the contribution only in the inductive
94 current (Kaladze, et al. 2013). The quantity $\zeta_z = \mathbf{e}_z \cdot \nabla \times \mathbf{v}$ is the z-component of the
95 vorticity. It should also be noted that the ambient magnetic field and Coriolis parameter are
96 spatially inhomogeneous, $f = 2\Omega_{0z}$ (Kaladze, et al., 2014):

$$97 \quad \begin{cases} \frac{\partial\Delta\psi}{\partial t} + \beta \frac{\partial\psi}{\partial x} + J(\psi, \Delta\psi) - \frac{1}{\mu_0\rho} \beta_B \frac{\partial h}{\partial x} = 0, \\ \frac{\partial h}{\partial t} + J(\psi, h) + \beta_B \frac{\partial\psi}{\partial x} + c_B \frac{\partial h}{\partial x} = 0, \end{cases} \quad (1)$$

99 In Eq. (1), h represents the z-component of perturbed magnetic field and $\beta = \frac{\partial f}{\partial y} = \frac{2\partial\Omega_{0z}}{\partial y}$
100 describes the latitudinal inhomogeneity present in the vertical component of angular velocity.
101 Also the parameter $c_B = \beta_B/en\mu_0$ with $\beta_B = \frac{\partial B_{0z}}{\partial y}$, describes the latitudinal inhomogeneity in
102 the background magnetic field, n is the number density of the charged particles, μ_0 is the
103 magnetic permeability and $J(a, b) = \frac{\partial a}{\partial x} \frac{\partial b}{\partial y} - \frac{\partial a}{\partial y} \frac{\partial b}{\partial x}$ is the Jacobian (responsible for the vector
104 nonlinearity) and $\Delta = \partial_x^2 + \partial_y^2$.

105
106 The boundary conditions that are fulfilled in this system are given as,

$$107 \quad \psi(0) = \psi(1) = 0, \quad (2)$$

108 which represents the flow's edges, specifically along the south and north direction [1, 2].
109
110
111

112 2.1 Perturbation and weakly nonlinear approach

113
114 The background stream function is considered as:

$$115 \quad \Psi(y) = - \int [U(y) - c_0] dy. \quad (3)$$

116
117 Here $U(y)$ describes the basic background flow with c_0 as a constant eigenvalue. The whole
118 stream function ψ is considered as the sum of background (zonal flow) stream function $\Psi(y)$
119 and a disturbed stream ψ' function, along with a normalized small parameter $\varepsilon \ll 1$. Which
120 forms a weakly nonlinear system, that is the subject of this study. While the perturbed magnetic
121 field is also characterized by a small a parameter ε . Therefore the stream function and the
122 magnetic perturbations takes the form,

$$123 \quad \begin{aligned} \psi &= \Psi(y) + \varepsilon\psi' = - \int [U(y) - c_0] dy + \varepsilon\psi', \\ h &= \varepsilon h' \end{aligned} \quad (4)$$

124
125 Using Eq (4) into (1) gives
126
127

$$128 \quad \begin{cases} \frac{\partial \Delta \psi'}{\partial t} + (U(y) - c_0) \frac{\partial \Delta \psi'}{\partial x} + (\beta - U'') \frac{\partial \psi'}{\partial x} + \frac{\beta_B}{\mu_0 \rho} \frac{\partial h'}{\partial x} + \varepsilon J(\psi', \Delta \psi') = 0, \\ \frac{\partial h'}{\partial t} + \varepsilon J(\psi', h') + (U(y) - c_0) \frac{\partial h'}{\partial x} + \beta_B \frac{\partial \psi'}{\partial x} + c_B \frac{\partial h'}{\partial x} = 0. \end{cases} \quad (5)$$

129
130 where $U'' = \frac{d^2 U}{dy^2}$.

131
132 By using the multiple scale analysis, we obtain the asymptotic solution where we can
133 take the spatial and temporal parameters as $X = \varepsilon x$ and time $T = \varepsilon^3 t$ respectively. Further
134 by eliminating h' from 5(b) into 5(a) we get the single equation for ψ'
135

$$136 \quad \mathcal{L}_0(\psi) + \varepsilon^2 \mathcal{L}_1(\psi) + \varepsilon J\left(\psi, \frac{\partial^2 \psi}{\partial y^2}\right) + \varepsilon^3 J\left(\psi, \frac{\partial^2 \psi}{\partial X^2}\right) + \varepsilon^4 \frac{\partial^3 \psi}{\partial T \partial X^2} = 0. \quad (6)$$

137
138 Throughout Eq. (6) the prime on the perturbed stream function is dropped.
139

Here we introduce the following linear differential operators:

$$\mathcal{L}_0 = \left[(U - c_0) \frac{\partial^2}{\partial y^2} + p(y) + \frac{\alpha(y)}{U - c_0 + c_B} \right] \frac{\partial}{\partial x}, \quad \mathcal{L}_1 = \frac{\partial}{\partial T} \frac{\partial^2}{\partial y^2} + (U - c_0) \frac{\partial^3}{\partial X^3}, \quad (7)$$

where $\alpha(y) = \frac{\beta_B^2}{\mu_0 \rho}$ and $p(y) = \beta - U''$. Here the parameter α involves the spatial inhomogeneous background magnetic field which was not considered before in Kaladze et al. [16].

Furthermore, we denote the disturbed stream function ψ (the asymptotic expansion) as:

$$\psi = \psi_0 + \varepsilon \psi_1 + \varepsilon^2 \psi_2 + \dots \quad (8)$$

By using Eq. (8) into Eq. (6), from the lowest order $O(\varepsilon^0)$, we get the following equation with conditions of boundary:

$$\mathcal{L}_0[\psi_0] = 0, \quad \text{with} \quad \psi_0 = 0 \text{ for } y = 0, 1. \quad (9)$$

The above equation (9) is a linear differential equation. By performing a separation of variables method for $\psi_0 = A(X, T) \Phi_0(y)$ into this form and substitute it into Eq. (7) we get the following equation with conditions of boundary:

$$\left(\frac{d^2}{dy^2} + \frac{p(y)}{U - c_0} + \frac{\alpha(y)}{(U - c_0)(U - c_0 + c_B)} \right) \Phi_0 = 0, \quad \text{with} \quad \Phi_0(0) = \Phi_0(1) = 0. \quad (10)$$

Here we consider $U - c_0 \neq 0$ and $U - c_0 + c_B \neq 0$. This is an eigenvalue problem for eigen value c_0 . By specifying $p(y)$ and $\alpha(y)$, $\Phi_0(y)$ can be found. Since $p(y)$ and $\alpha(y)$ have dependence on the variable y , it is not easy to solve this eigen value problem analytically. From the lowest order $O(\varepsilon^0)$, we see that the problem is time independent, but cannot be analytically solved as we have not substituted any definite dependence on y for the parameters $p(y)$ and $\alpha(y)$. Thus, in order to gain more information about the amplitude of these waves, we go to the next order i.e. $O(\varepsilon^1)$ from Eqs. (7) and (8), we obtain

$$\mathcal{L}_0[\psi_1] = -\mathcal{J} \left(\psi_0, \frac{\partial^2 \psi_0}{\partial y^2} \right) \equiv F_1 = A \frac{\partial A}{\partial X} \left(\frac{p(y)}{U - c_0} + \frac{\alpha}{(U - c_0)(U - c_0 + c_B)} \right)_y \Phi_0^2, \quad (11)$$

Furthermore, we carry out a separation of variables in the following manner $\psi_1 = \frac{1}{2} A^2(X, T) \Phi_1(y)$ for non-singular neutral solutions into (11)

$$\left(\frac{d^2}{dy^2} + \frac{p(y)}{U - c_0} + \frac{\alpha}{(U - c_0)(U - c_0 + c_B)} \right) \Phi_1 = \left(\frac{p(y)}{U - c_0} + \frac{\alpha}{(U - c_0)(U - c_0 + c_B)} \right)_y \frac{\Phi_0^2}{(U - c_0)}, \quad (12)$$

For the given boundary conditions $\Phi_1(0) = \Phi_1(1) = 0$. To get amplitude we solve Eqs.(7) and (8) in the next order i.e. $O(\varepsilon^2)$ which gives

$$\mathcal{L}_0[\psi_2] = -\mathcal{L}_1[\psi_0] - \mathcal{J} \left(\psi_0, \frac{\partial^2 \psi_1}{\partial y^2} \right) - \mathcal{J} \left(\psi_1, \frac{\partial^2 \psi_0}{\partial y^2} \right) \equiv F_2, \quad (13)$$

with $\psi_2(0) = \psi_2(1) = 0$.

185
186 Here it is pointed out that the dispersion effect competes with weakly nonlinear effect.
187

188 Furthermore, we take $\psi_2 = B(X, T)\Phi_2(y)$ and multiply Eq. (13) by ψ_0 and integrate over y ,
189 which yields

$$190 \int_0^1 dy \frac{F_2}{U-c_0} \Phi_0 = 0. \quad (14)$$

192 By substituting F_2 and using $\psi_1 = \frac{1}{2} A^2(X, T) \Phi_1(y)$ into above Eq. (14) results in the
193 modified KdV (MKdV) equation

$$194 \frac{\partial A}{\partial T} + N A^2 \frac{\partial A}{\partial X} + D \frac{\partial^3 A}{\partial X^3} = 0. \quad (15)$$

196 Here

$$197 N = \frac{I_2}{I_0}, \quad D = -\frac{I_1}{I_0}, \quad (16)$$

$$200 \left\{ \begin{array}{l} I_0 = \int_0^1 dy \Phi_0^2(y) \left[\frac{p(y)}{(U-c_0)^2} + \frac{\alpha}{(U-c_0)^2(U-c_0+c_B)} \right], \\ I_1 = \int_0^1 dy \Phi_0^2(y) \\ I_2 = \int_0^1 dy \frac{\Phi_0^2(y)}{U-c_0} \left\{ \begin{array}{l} \frac{3}{2} \left(\frac{p(y)}{U-c_0} + \frac{\alpha}{(U-c_0)(U-c_0+c_B)} \right)_y \Phi_1(y) \\ -\frac{1}{2} \Phi_0^2(y) \left[\left(\frac{p(y)}{U-c_0} + \frac{\alpha}{(U-c_0)(U-c_0+c_B)} \right)_y \frac{1}{U-c_0} \right]_y \end{array} \right\} \end{array} \right\}, \quad (17)$$

202 Kaladze et al. [16] and Jian et al. [17] also obtained the same MKdV equation for Rossby waves
203 as equation (15) and pointed out that the background flow shear is a necessary condition for
204 the existence of solitary waves, whereas in this work, we get the MKdV for the Rossby-
205 Khantadze waves where the coefficients have been modified by inclusion of inhomogeneity in
206 geomagnetic field. Moreover, the effect of shear basic flow on the spatial structure, propagation
207 velocity and wave width of solitary Rossby waves have been studied.

208 Here the function $\beta(y)$, $\alpha(y)$ and $U(y)$ are related to the coefficients N and D.
209

210 The solution of modified KdV equation (15) is,
211

$$212 A(x, t) = \pm \sqrt{\frac{6c}{N}} \operatorname{sech} \left(\sqrt{\frac{c}{D}} (x - ct) \right). \quad (18)$$

213 Where the coefficients D, and N are defined above.
214

215 3. Discussion

216 In the given paper, we studied the nonlinear dynamics of large-scale electromagnetic
217 Rossby-Khantadze waves with zonal flows in E-ionospheric plasma. Both the latitudinal
218 inhomogeneities in angular velocity of the earth's rotation and the geomagnetic field are taken
219 into account. The inhomogeneity of the magnetic field with latitude is responsible for coupled
220 Rossby-Khantadze waves. Such coupling results in an appearance of dispersion of Khantadze
221 waves. To derive the nonlinear modified KdV we used the multiple scale analysis technique.
222
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224

225 From the lowest order of $O(\epsilon^0)$, we get an eigen value problem with constant eigen value c_0
226 along with the boundary conditions. The parameters $p(y)$ and $\alpha(y)$ have dependence on the
227 variable y , making it not possible to solve this eigen value problem analytically. From the next
228 order $O(\epsilon^1)$, by using separation of variables techniques and after doing some mathematical
229 manipulations we arrive at the modified KdV equation of one dimension. The derived quantity
230 $\sqrt{\frac{6c}{N}}$ describes the amplitude of solitary waves. The obtained coefficients N and D depend on
231 spatial inhomogeneous Coriolis force $\alpha(y)$ and background magnetic field $\beta(y)$, respectively.
232 The parameter $\sqrt{\frac{6c}{N}}$ obtained above describes the amplitude of obtained such Rossby-
233 Khantadze waves.

234 **AUTHOR DECLARATIONS:**

235 **Conflict of Interest**

236 The authors have no conflicts to disclose.

237 **Data Availability**

238 The data that support the findings of this study are available within the article.

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