

Climate Bifurcations in a Schwarzschild Equation Model of the Arctic Atmosphere

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Abstract. A column model of the Arctic atmosphere is developed including the nonlinear positive feedback responses of surface albedo and water vapour to temperature. The atmosphere is treated as a gray gas and the flux of longwave radiation is governed by the two-stream Schwarzschild equations. Water vapour concentration is determined by the Clausius-Clapeyron equation. Representative carbon pathways (RCPs) are used to model carbon dioxide concentrations into the future. The resulting nine-dimensional two-point boundary value problem is solved under various RCPs and the solutions analyzed. The model predicts that under the highest carbon pathway, the Arctic climate will undergo an irreversible bifurcation to a warm steady state, which would correspond to annually ice-free conditions. Under the lowest carbon pathway, corresponding to very aggressive carbon emission reductions, the model exhibits only a mild increase in Arctic temperatures. Under the two intermediate carbon pathways, temperatures increase more substantially, and the system enters a region of bistability where external perturbations could possibly cause an irreversible switch to a warm, ice-free state.

1 Introduction

Climate change is causing rapid temperature increases in the polar regions. A fundamental question is whether these temperature increases are reversible. If humanity fails to prevent a substantial warming of the planet in the next few decades, which is appearing to be more and more likely, will it be possible in the future to reverse our effects on climate enough to restore lower temperatures? Or will we have passed a tipping point beyond which return to the present state is impossible? We address this question in particular for the Arctic, where the observed climate change is the most dramatic.

The Earth's climate is an extremely complex system. Modelling efforts range from simple models attempting to isolate the most pertinent features, to very complicated numerical models trying to capture as many details as possible. The model presented here is close to the simple end of this spectrum, although not as simple as some, in that it is a nine-dimensional nonlinear two-point boundary value problem. The advantage of relatively simple models is that they allow more direct analysis of cause and effect, which is often obscured in highly complicated models.

The term “tipping point” is used by different researchers in various ways; see Russill (2015) and Lenton et al. (2008) for some definitions and discussion of the term. In all cases however, tipping points are associated with large qualitative changes in a system due to relatively small changes in the parameters, or “forcings” that drive the system. In the present paper,

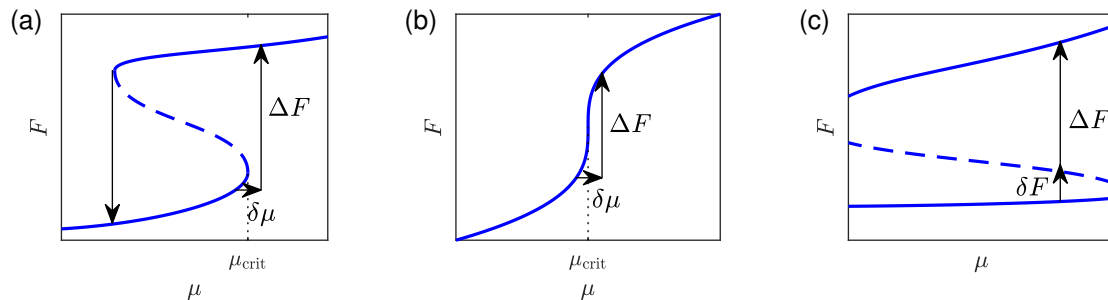


Figure 1. (a) Hysteresis arising from two saddle-node bifurcations, (b) cusp bifurcation, (c) bistable system. Solid curves indicate stable states, and dashed curves are unstable states. In (a) and (b) a small change $\delta\mu$ in the control parameter near the bifurcation value μ_{crit} causes a large change ΔF in the system. In (a) the return bifurcation happens at a different value of μ . In (c) a small perturbation δF in the system causes a large change ΔF .

25 tipping points arise as a result of saddle-node and cusp bifurcations in the mathematical model. The mathematical theory of bifurcations is well-developed (Kuznetsov, 2004) and employed here. Figure 1 illustrates the typical behavior associated with these bifurcations. In Fig. 1(a) there are two saddle-node bifurcations resulting in a parameter interval of bistability, that is, two stable states coexist for an interval of μ values. If the system is on the lower stable state, then, as μ increases through a critical value μ_{crit} , there is an abrupt jump to the upper stable state. In contrast, as μ decreases, the jump back to the lower state does not occur until a much smaller critical value of μ . This phenomenon is called hysteresis. If present in the Earth's climate system, it implies that once the upward jump occurs, it may be very difficult to achieve the reverse jump back to the original climate state. Fig. 1(b) illustrates the situation of a cusp bifurcation, where the two saddle-node bifurcations in (a) have coalesced, for example, as another parameter of the system is varied. In this situation a small change in μ will also cause a large change in the system F , although it will be smooth and reversible. In Fig. 1(c), even though the saddle-node bifurcations may be outside the parameter interval of interest, abrupt large transitions in the system can result from a small noise- or perturbation-induced change to the system even when the parameter value remains constant. It is the presence of saddle-node bifurcations in a mathematical model, even if not occurring precisely at the system's current parameter value, that is the root cause of all of the behaviors shown in Fig. 1.

For a tipping point to be present, the underlying mathematical model will be characterized by nonlinearity, generally in the form of a positive feedback that accelerates change once change has begun. For the Arctic, one of the primary positive feedbacks is the surface albedo. When the Arctic Ocean is frozen, the surface reflects a significant portion of the insolation back into space, but open water absorbs much more heat from the sun. Timing of the melt in the spring has significant impact (Zheng et al., 2021). An earlier melt means considerably more heat is absorbed by open water, raising the water temperature and delaying freeze up in the autumn. The freeze up date for the Beaufort, Chukchi, Laptev, and Kara Seas, for example, has been getting later by 6–11 days per decade since 1979 (Stroeve et al., 2014). September sea ice extent has been decreasing at an accelerating rate. The linear trend from 1979 to 2001 is -7% per decade, but including data up to 2013, the linear trend is -14%

per decade (Stroeve et al., 2014). Thus the observational evidence indicates that the processes behind this phenomenon are not linear at all, but nonlinear.

Past studies on GCMs have given mixed results regarding the presence of multiple stable states for ice conditions in the Arctic. Some indicate that there appears to be a continuous transition from perennial ice cover to annually ice-free that is reversible, (Schröder and Connolley, 2007; Tietsche et al., 2011; Armour et al., 2011). Other studies have shown evidence for nonlinear behavior in sea ice loss, especially in the transition from seasonally ice-free to annually ice-free, (Winton, 2006, 2008; Ridley et al., 2008). On the other hand, smaller conceptual models generally show bistability and abrupt transitions in sea ice cover, (Thorndike, 1992; Müller-Stoffels and Wackerbauer, 2011; Eisenman and Wettlaufer, 2009; Björk and Söderkvist, 2002; Abbot et al., 2011; Merryfield et al., 2008; Flato and Brown, 1996). The most common result from all these models seems to be that sea ice will likely transition from perennial to seasonally ice-free in a continuous, reversible manner, but significant warming beyond that point will likely cause an abrupt change to annually ice-free (Bathiany et al., 2016). See the introduction in Eisenman (2012). The model we present here is an annually averaged model with no seasonal component. It is not a model of sea ice in particular but rather a column model of the atmosphere that incorporates a nonlinear albedo response to surface temperature. Bistability in our model with both warm and cold solutions, corresponds to annually averaged ice-covered, or ice-free situations.

The Arctic climate model presented here is motivated by three observations. First is the observation that the climate changes taking place on the Earth today are most dramatic in the high Arctic. Therefore, it is prudent to put a special focus on understanding Arctic climate change. Second, irreversible change is inevitably the result of nonlinear geophysical processes. So, while this model is kept very simple, it does include significant nonlinear phenomena that can lead to tipping points. Third, the 3D spherical shell of atmosphere of the Earth is rotationally symmetric about the polar axis, if annually and zonally averaged. Due to the rotation of the Earth, Hadley, Ferrel and Polar Cells form in the global circulation. If perfect rotational symmetry is assumed, the polar axis becomes flow-invariant, and this remains approximately true for the real Earth. Thus, a 1D model restricted to the polar axis, can be expected to give useful information about climate in a neighbourhood of the pole. The study of a rotationally symmetric spherical shell model by Lewis and Langford (2008) gives support to this hypothesis. A vertical column of atmosphere at other points on Earth would have a horizontal component of velocity, invalidating the type of analysis used here. Globally averaged climate models do reduce to one (vertical) dimension, but they give little information specific to the Arctic.

The present model builds on the simple energy balance slab models of Dortmans et al. (2019), which was applied to paleo-climate transitions, and Kypke et al. (2020), which was applied to anthropogenic climate change. The primary improvement of the present model is a more physically accurate description of the atmosphere. Instead of using a slab to represent a uniform atmosphere with absorption properties similar to the real atmosphere, here we use the Schwarzschild two-stream equations to model absorption in the atmosphere explicitly as a function of altitude (Pierrehumbert, 2010, pg. 191).

A bifurcation analysis is performed on the model, tracking the steady-state solutions as carbon dioxide levels increase. The question of reversibility is a question of whether the current cold state simply warms but persists. The disappearance of this cold state through a saddle-node bifurcation would result in an abrupt change in climate that may be practically irreversible.

The simpler model of Kypke et al. (2020) showed this behaviour under certain CO₂ representative concentration pathway scenarios. We seek here to determine if the present, more accurate model, also displays this behaviour.

85 Section 2 and Appendix A provide a detailed derivation of the model. The model parameter values and calibration of some of them to empirical data is presented in Appendix B. Although much of the detail is relegated to the appendices, the authors feel this detail constitutes an essential part of the contribution of the manuscript, providing clarity, justification of choices, and the information necessary for replication. Hence we consider them essential reading. Section 3 presents the results, and the conclusions are in Section 4.

2 Model

90 The model is developed from first principles and has the following features.

- The atmosphere is a one-dimensional column at the North Pole with physical properties that vary with altitude, from the surface to the tropopause.
- The incoming solar radiation is annually averaged and undergoes reflection and absorption in the atmosphere as well as at the Earth's surface.
- 95 – The surface albedo is a nonlinear function of the surface temperature.
- A well-mixed surface boundary layer is included.
- The Earth emits longwave radiation as a black body.
- The atmosphere is considered to be a gray gas.
- The Schwarzschild two-stream equations govern the absorption and emission of both upward and downward directed longwave radiation in the atmosphere.
- 100 – The atmospheric absorption of longwave radiation is due to three factors: water vapour, CO₂ concentration, and clouds.
- Water vapour concentration is governed by the nonlinear Clausius-Clapeyron equation.
- Transfer of latent and sensible heat from the surface to the atmosphere is modelled.
- Both ocean and atmospheric meridional heat transport to the Arctic are dictated by empirical values.
- 105 – In the Arctic, there is a slow downward movement of air in the column corresponding to the polar circulation cell near the pole (Lewis and Langford, 2008; Langford and Lewis, 2009; Lutgens and Tarbuck, 2019). This is achieved via mass transport of air into the column in its upper portion and out of the column near the bottom.
- The radiation absorption coefficients are calibrated by fitting the model to global average data.
- The functional forms of the mass transport and atmospheric heat transport are used to calibrate the model to an empirical Arctic temperature profile.
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The annually and zonally averaged Earth’s atmosphere is rotationally symmetric around the polar axis, which is invariant under the flow. Therefore, if one considers a column of the atmosphere near the North Pole, it is reasonably approximated by a one-dimensional model with altitude-varying quantities. This approximation becomes exact in the limit as the diameter of the column shrinks to zero. Alternatively, one can view the model as a meridional and zonal average over a cylinder centred at the North Pole. Further, although the Arctic Ocean is not zonally symmetric, in the above view, the contribution of ocean heat transport can be reasonably captured as a scalar quantity. Thus our model is more precisely a model of the North Pole rather than the Arctic. Nonetheless, we do use some empirical data for the region north of 70° to calibrate the model for two reasons: 1) data further north is not readily available, and 2) the data we use is not likely to alter too much if it were measured closer to the pole. The values for atmospheric heat transport and ocean heat transport are two that may change significantly as one moves north from 70° , and we therefore analyze the behaviour of the model over a wide range for these parameter values.

The model domain is a vertical cylinder of cross sectional area A (m^2) and circumference C_b (m). The atmosphere is assumed to be uniform in the cross-sectional direction so that the model’s dependent variables can be interpreted as cross-sectional averages that vary with the one-dimensional vertical coordinate $z \in [0, z_T]$, where z_T (m) is the height of the tropopause. This domain is divided into a surface boundary layer with height z_B (m), $z_B \ll z_T$, and the troposphere proper, $z \in [z_B, z_T]$. The model consists of a set of initial value problems (IVPs), with a spatial independent variable z on $[0, z_B]$, that can be solved analytically, and a two-point boundary value problem (BVP) on $[z_B, z_T]$ that depends on the solutions to the IVPs. The model has equations governing the vertical wind speed, w (m s^{-1}), the air density, ρ (kg m^{-3}), the upward and downward longwave radiation, I_+ and I_- (W m^{-2}), the downward shortwave radiation, I_S (W m^{-2}), the latent and sensible heat transport, F_C , (W m^{-2}), and the temperature T (K). Any reflection of shortwave radiation from either the surface or the atmosphere is ignored and is simply considered as leaving the system. The model is depicted in Fig. 2 and is derived and explained in detail in the following subsections. Many of the details of the model including its non-dimensionalization, the vanishing conduction limit, and modelling choices used for various functional forms, are in Appendix A. Calibration of the model parameters to empirical data is described in detail in Appendix B and the reader is referred to Tables B1 and B2 of that appendix for the values of the parameters.

2.1 Mass, Momentum, and Energy Balance

The model equations in the troposphere are developed from the fundamental transport theorem in one spatial dimension:

$$\partial_t f + \partial_z \chi = S, \tag{1}$$

where f is the density of some “property”, χ is the flux of that property, and S is a source/sink term. The time derivative term will be taken as zero since only the steady-state solution is considered. The properties subject to this equation are mass, momentum, and energy. To model the Arctic, the cylinder is centred at the north pole, and, since the atmospheric polar cell has slow downward movement near the pole, it is assumed that $w < 0$.

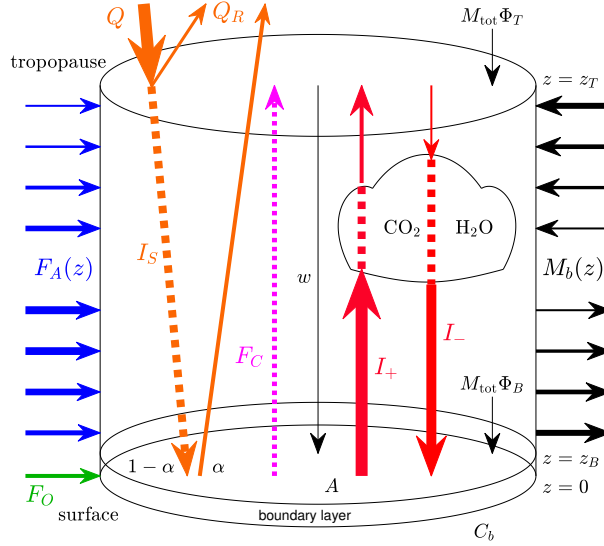


Figure 2. Schematic illustration of the model. Symbols as described in the text.

2.1.1 Mass

If the property f in (1) is the mass density, ρ (kg m^{-3}), then the flux is $\chi = \rho w$. There is mass flux across the vertical boundary of the cylinder, $M_b(z)$ ($\text{kg m}^{-2} \text{s}^{-1}$), which is assumed to be immediately spread out evenly across the layer, hence the mass flux across the vertical boundary in the model is really a mass source term in the interior giving

$$S = \frac{\text{mass entering into cylinder layer of width } \Delta z}{\text{volume of layer}} = \frac{M_b(z)C_b\Delta z}{A\Delta z} = \frac{C_b}{A}M_b(z).$$

Thus, at steady-state the mass balance equation is

$$\frac{d(\rho w)}{dz} = \frac{C_b}{A}M_b(z).$$

The mass flux through the vertical boundary into the column is written as

$$M_b(z) = \frac{A}{C_b(z_T - z_B)}M_{\text{tot}}\phi\left(\frac{z - z_B}{z_T - z_B}\right),$$

where M_{tot} ($\text{kg m}^{-2} \text{s}^{-1}$) is a nonnegative constant and $\phi(x) : [0, 1] \mapsto \mathbb{R}$ is a dimensionless function that represents the portion of inward mass flux across the vertical boundary of the column at the given altitude. Positive ϕ indicates inward flow. The ratio of the cross sectional area of the column to the area of its side, $A/(C_b(z_T - z_B))$, in the definition of $M_b(z)$ is included as a useful simplifying convenience. The (positive inward) mass fluxes across the bottom and top boundaries of the column are given by $M_{\text{tot}}\Phi_B$ and $M_{\text{tot}}\Phi_T$, respectively, where Φ_B and Φ_T are dimensionless constants in the interval $[-1, 1]$. These quantities must satisfy

$$\int_0^1 \phi(x) dx + \Phi_B + \Phi_T = 0, \quad \text{and} \quad \int_0^1 \max(\phi(x), 0) dx + \max(\Phi_B, 0) + \max(\Phi_T, 0) = 1. \quad (2)$$

The first of these conditions dictates that there is no net mass entering the system, and the second is a normalization condition so that M_{tot} represents both the total mass entering the column per unit cross sectional area, and the magnitude of the total mass leaving the column per unit cross sectional area. In order to obtain downward vertical flow throughout the cylinder, it shall be assumed that $\Phi_B < 0$, $\Phi_T > 0$, $\phi \leq 0$ in the lower part of the cylinder, and $\phi \geq 0$ in the upper part. With these definitions, the mass balance equation becomes

$$\frac{d(\rho w)}{dz} = \frac{M_{\text{tot}}}{z_T - z_B} \phi \left(\frac{z - z_B}{z_T - z_B} \right). \quad (3)$$

2.1.2 Momentum

Now take the property f in (1) to be the momentum density, ρw . The vertical flux is $\chi = \rho w^2$, and the source term S has two components, one due to contact forces (stress) (van Groesen and Molenaar, 2017, pg. 56) and one due to internal body forces (gravity):

$$S = -\frac{dP}{dz} - \rho g,$$

where P (N m^{-2}) is the pressure and g (m s^{-2}) is the gravitational acceleration. It is assumed that mass entering the cylinder from the vertical boundary has no vertical momentum. Thus the momentum balance equation at steady-state is

$$\frac{d(\rho w^2)}{dz} = -\frac{dP}{dz} - \rho g. \quad (4)$$

(In the case of no flow ($w = 0$) the above would read $\frac{dP}{dz} = -\rho g$, which is the hydrostatic equation.)

2.1.3 Energy

Finally consider the case where the property in (1) is the total energy density given by

$$e = \frac{1}{2}\rho w^2 + \rho g z + c_v \rho T,$$

which corresponds to the sum of kinetic energy, gravitational potential energy, and internal heat energy densities. Here c_v ($\text{J kg}^{-1} \text{K}^{-1}$) is the specific heat capacity of the air. The flux has two components, one due to advection and one due to conduction:

$$\chi = ew - k \frac{dT}{dz},$$

where k ($\text{W m}^{-1} \text{K}^{-1}$) is the thermal conductivity. The source/sink S has eight terms, one due to work done by contact forces, two due to mass entering or leaving across the vertical boundary (one of these accounts for gravitational potential energy and the other internal heat energy; there is no addition to kinetic energy since the mass appearing has no velocity), three terms due to radiation (shortwave downward, and longwave upward and downward), one due to latent and sensible heat transport, and one due to atmospheric heat transport. It is important to distinguish the difference between the mass transport across the boundary and the atmospheric heat transport. It is assumed that the mass moving across the vertical boundary is at the

same temperature as the mass inside at each altitude. Since mass transfer across the vertical boundary is inward in the upper portion, where the temperature is cooler, and outward in the bottom portion, mass transport results in a net transport of heat out through the vertical boundary, but this will be small since the mass flux, $M_b(z)$, is small. The reason $M_b(z)$ is small is that it generates the average slow movement of air downward near the north pole (about 1 mm/s), due to the circulation of the polar cell. This slow averaged circulation of air does not account for the atmospheric heat transport. The main transport of heat in the atmosphere is via turbulent mixing captured in our model by a source term, $F_A(z)$ (W m^{-3}), whose functional form is discussed in Section A3. Thus S is given by

$$S = -\frac{d(Pw)}{dz} + \frac{M_{\text{tot}}}{z_T - z_B} \phi \left(\frac{z - z_B}{z_T - z_B} \right) gz + \frac{M_{\text{tot}}}{z_T - z_B} \phi \left(\frac{z - z_B}{z_T - z_B} \right) c_v T - \frac{dI_+}{dz} + \frac{dI_-}{dz} + \frac{dI_S}{dz} - \frac{dF_C}{dz} + F_A(z).$$

The governing equations for the longwave radiation intensities are the two-stream Schwarzschild equations and for the short-wave radiation a standard absorption equation:

$$\frac{dI_+}{dz} = -\kappa (I_+ - \sigma T^4), \quad (5)$$

$$\frac{dI_-}{dz} = \kappa (I_- - \sigma T^4). \quad (6)$$

$$\frac{dI_S}{dz} = k_S \rho I_S, \quad (7)$$

where k_S ($\text{m}^2 \text{kg}^{-1}$) is the shortwave absorption coefficient, σ ($\text{W m}^{-2} \text{K}^{-4}$) is the Stefan-Boltzmann constant, and κ (m^{-1}) is the long wave absorption coefficient with terms corresponding to absorption by clouds, carbon dioxide and water vapour:

$$\kappa(\rho, T) = k_{Cl} + k_C \frac{M_{\text{CO}_2}}{M_A} \left(\frac{\mu}{10^6} \right) \rho + k_W \delta \left(\frac{z - z_B}{z_T - z_B} \right) P^{\text{sat}}(T). \quad (8)$$

Here k_{Cl} (m^{-1}), k_C ($\text{m}^2 \text{kg}^{-1}$), and k_W ($\text{s}^2 \text{kg}^{-1}$) are absorption coefficients that will be calibrated, μ (ppm) is the CO_2 concentration expressed as the ratio of moles of CO_2 to moles of dry air, M_{CO_2} and M_A (kg mol^{-1}) are the molar masses of CO_2 and dry air, respectively, $\delta \left(\frac{z - z_B}{z_T - z_B} \right)$ is the relative humidity at altitude z , and $P^{\text{sat}}(T)$ (N m^{-1}) is the saturated water vapour partial pressure at temperature T . The dependence of this last quantity on T is given by the Clausius-Clapeyron equation

$$P^{\text{sat}}(T) = P^{\text{sat}}(T_R) \exp \left(\frac{L_v}{R_W T_R} \frac{T - T_R}{T} \right), \quad (9)$$

where $P^{\text{sat}}(T_R)$ is the pressure at a reference temperature T_R (which we take to be 273.15 K), L_v ($\text{m}^2 \text{s}^{-2}$) is the latent heat of vapourization for water, and $R_W = R/M_W$ ($\text{J K}^{-1} \text{kg}^{-1}$) is the gas constant for water, R ($\text{J K}^{-1} \text{mol}^{-1}$) is the universal gas constant, and M_W (kg mol^{-1}) is the molar mass of water. The corresponding density is $\rho_W^{\text{sat}}(T) = P^{\text{sat}}(T)/(R_W T)$ by the ideal gas law. The vertical heat transport (latent and sensible heat) is assumed to be governed by a simple exponential decay

$$\frac{dF_C}{dz} = -bF_C, \quad (10)$$

where b (m^{-1}) is a suitable decay constant. Substituting these expressions for the total energy density, its flux and sources into (1), and combining the expressions from (5)–(10), the energy balance equation at steady-state for $z \in [z_B, z_T]$ is given by

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$$\frac{d}{dz} \left(\frac{1}{2} \rho w^3 + \rho g z w + c_v \rho T w \right) - k \frac{d^2 T}{dz^2} = - \frac{d(Pw)}{dz} + \frac{M_{\text{tot}}}{z_T - z_B} \phi \left(\frac{z - z_B}{z_T - z_B} \right) (gz + c_v T) + \kappa(\rho, T) (I_+ + I_- - 2\sigma T^4) + k_S \rho I_S + b F_C + F_A(z). \quad (11)$$

The nonlinear effects of both water vapour and carbon dioxide concentration on longwave radiation absorption in the atmosphere are contained within the factor $\kappa(\rho, T)$, defined by (8). In earlier work (Dortmans et al., 2019; Kypke et al., 2020) these two effects were studied separately, before combining them. It was shown there that, if the atmosphere becomes warmer, then the concentration of water vapour increases due to the Clausius-Clapeyron relation, and this accelerates the greenhouse warming of the atmosphere, well beyond that due to carbon dioxide alone. This is an important positive feedback in the model.

In order to complete the system a constitutive relation between the density ρ and the pressure P is needed for which we use the ideal gas law,

$$225 \quad P = R_A \rho T, \quad (12)$$

where $R_A = R/M_A$ ($\text{J kg}^{-1} \text{K}^{-1}$) is the gas constant for air.

The mass, momentum, and energy balance equations, (3), (4), and (11), along with the Schwarzschild equations (5)–(6), and the equations governing shortwave absorption (7) and sensible and latent heat transport (10), are the differential equations for the BVP for $z \in [z_B, z_T]$, with dependent variables w , ρ , I_+ , I_- , I_S , F_C , T , and $\frac{dT}{dz}$. Equations (8), (9), and (12) define certain quantities in these differential equations in terms of these dependent variables. The forms of the functions $F_A(z)$ and $\phi(z)$ are prescribed; the process of choosing these functions is described in detail in Appendix A3.2 and A3.3, respectively. Before discussing the boundary conditions for the BVP it is necessary to consider the surface boundary layer.

230

2.2 Surface Boundary Layer

The model includes a boundary layer extending from $z = 0$ to $z = z_B$. It is assumed that this layer is well-mixed so that temperature $T_B = T(z_B)$, density $\rho_B = \rho(z_B)$, and relative humidity $\delta_B = \delta(z_B)$ in this layer are constant. The temperature of the surface, T_S , can in general be larger or smaller than T_B .

The primary reason for including a boundary layer is a numerical one. As shown in Appendix A2, the model is numerically stiff due to the thermal conductivity of air being very small. To remove the stiffness, a limit to vanishing conduction is taken, and this results in an algebraic expression for the temperature gradient that includes the vertical wind speed as a factor in the denominator. As the vertical wind speed must be zero at the Earth's surface, there is a singularity in the temperature gradient there. The introduction of the surface boundary layer avoids this singularity.

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The total mass crossing from the atmosphere into the boundary layer per unit time is $M_{\text{tot}} \Phi_B A$. This quantity is negative, since $\Phi_B < 0$, indicating flow out of the atmosphere and into the boundary layer. This mass exits through the vertical boundary

of the layer with an assumed constant mass flux K , at each z , and conservation of mass dictates

$$245 \quad \int_0^{z_B} K C_b dz = M_{\text{tot}} \Phi_B A \quad \implies \quad K = \frac{M_{\text{tot}} \Phi_B A}{C_b z_B}.$$

($K < 0$ indicates the flux is outward.) This exiting mass carries gravitational potential energy. The change of potential energy in a slab of height Δz at height z in the boundary layer is $C_b \Delta z K g z$ so that the total change in potential energy over the boundary layer is

$$\Delta_{PE} = \int_0^{z_B} C_b K g z dz = \frac{1}{2} M_{\text{tot}} \Phi_B A g z_B. \quad (13)$$

250 Consistent with the modelling assumption that mass flux across the vertical boundary conveys no momentum or kinetic energy to the system, the loss of mass out of the vertical boundary of the boundary layer also has no effect on the momentum or kinetic energy. Further, since the temperature in the boundary layer, T_B , is assumed to be equal to the temperature of the atmosphere at $z = z_B$, it follows that there is also no net energy change in the boundary layer due to advection of internal energy — the internal energy entering via advection at the top of the layer is equal to the internal energy leaving the layer through the vertical
255 boundary.

Consider now the energy balance at the Earth's surface. There is energy transport from the surface to the boundary layer in the form of sensible and latent heat, which is modelled, as per Pierrehumbert (2010, pgs. 396–398), as

$$F_{C0}(\rho_B, T_B, T_S) = F_C(0) = F_{\text{sensible}} + F_{\text{latent}} = c_v C_D U \rho_B (T_S - T_B) + \frac{L_v}{R_W T_B} C_D U (P^{\text{sat}}(T_S) - \delta(0) P^{\text{sat}}(T_B)), \quad (14)$$

where C_D is a dimensionless drag coefficient, U (m s^{-1}) is the horizontal wind speed, and $P^{\text{sat}}(T)$ is given by (9). Along
260 with this there is energy input to the surface from the sun, $I_S(0)$, some of which is reflected by the surface albedo, longwave radiation both inward, $I_-(0)$, and outward, $I_+(0)$, and ocean heat transport, F_O (W m^{-2}). Therefore the energy balance at the surface is

$$F_O - I_+(0) + I_-(0) + I_S(0)(1 - \alpha(T_S)) - F_C(0) = 0, \quad (15)$$

where

$$265 \quad \alpha(T_S) = \frac{1}{2} \left[(\alpha_w + \alpha_c) + (\alpha_w - \alpha_c) \tanh \left(\frac{T_S - T_R}{T_R \omega} \right) \right] \quad (16)$$

is the surface albedo, here modelled as a sigmoid function increasing from α_c at cold temperatures to α_w at warm temperatures, with the midway point being at the reference temperature T_R (freezing point) and with a steepness of transition determined by the dimensionless constant ω .

Now consider the energy balance for the combined surface and boundary layer (one could alternatively consider just the
270 boundary layer without the surface, but the chosen formulation results in a slightly smaller equation). Input energy to this combined surface and boundary layer includes ocean heat transport, and short and longwave radiation entering at z_B . Output

energy includes upward longwave radiation at z_B , the shortwave radiation reflected from the surface, and the latent and sensible heat F_C at z_B . Further, there are kinetic and gravitational potential energy fluxes and heat conduction in/out of the layer through its top at z_B , and there is gravitational potential energy loss through the vertical boundary given by (13). Therefore the energy
275 density balance for the combined surface and boundary layer is

$$F_O + I_S(z_B) + I_-(z_B) - I_+(z_B) - I_S(0)\alpha(T_S) - F_C(z_B) - \frac{1}{2}\rho_B w(z_B)^3 - \rho_B g z_B w(z_B) + k \frac{dT}{dz}(z_B) + \frac{1}{2}gM_{\text{tot}}\Phi_B z_B = 0. \quad (17)$$

Since temperature, pressure, and relative humidity are constant in the boundary layer, the radiation equations may be solved
280 analytically inside the layer in order to relate the radiation terms at $z = 0$ with those at $z = z_B$. The simple ODE for F_C is also easily solved in the boundary layer. The initial (spatial independent variable $z = 0$) condition for the upward longwave radiation, $I_+(0)$, is that it is equal to the black body radiation of the surface, σT_S^4 . The initial condition for that latent heat, $F_C(0)$, is given by (14). Initial conditions for I_- and I_S are not necessary since only a relation between the values of these functions at 0 in terms of their value at z_B is required. The IVPs for I_+ , and F_C , and the ODEs for I_- and I_S in the boundary
285 layer are:

$$\begin{aligned} \frac{dI_+}{dz} &= -\kappa(\rho_B, T_B)(I_+ - \sigma T_B^4), & I_+(0) &= \sigma T_S^4, \\ \frac{dF_C}{dz} &= -bF_C, & F_C(0) &= F_{C0}(\rho_B, T_B, T_S), \\ \frac{dI_-}{dz} &= \kappa(\rho_B, T_B)(I_- - \sigma T_B^4), \\ \frac{dI_S}{dz} &= k_S \rho_B I_S, \end{aligned}$$

290 and their solutions, via standard means, give

$$I_+(z_B) = (\sigma T_S^4 - \sigma T_B^4)e^{-\kappa(\rho_B, T_B)z_B} + \sigma T_B^4, \quad (18)$$

$$F_C(z_B) = F_{C0}(\rho_B, T_B, T_S)e^{-bz_B}, \quad (19)$$

$$I_-(0) = (I_-(z_B) - \sigma T_B^4)e^{-\kappa(\rho_B, T_B)z_B} + \sigma T_B^4, \quad (20)$$

$$I_S(0) = I_S(z_B)e^{-k_S \rho_B z_B}. \quad (21)$$

295 Equations (18)–(19) provide two boundary conditions for the BVP on the troposphere. The energy balance equations (15), (17) along with equations (16), (20), and (21) provide two further boundary conditions.

2.3 Boundary Conditions for the BVP

There are eight unknown dependent variables: w , ρ , I_+ , I_- , I_S , F_C , T , and $\frac{dT}{dz}$; in addition, the surface temperature, T_S , is an unknown constant (independent of z) that is determined through (14), while the pressure, P , can be written in terms of the
300 others via (12). The boundary conditions for the system on the interval $[z_B, z_T]$ are wind speed at z_B given by the requirement

that the advected mass $Aw(z_B)\rho_B$ equals the mass flux $M_{\text{tot}}\Phi_B A$, pressure at the surface equal to the standard pressure, P_0 , upward longwave radiation at z_B given by (18), vertical heat transport at z_B given by (19), the energy balance equations (15) and (17) with expressions from (20) and (21) substituted in, no downward longwave radiation at z_T , shortwave radiation at z_T equal to the insolation, Q , less what is reflected by the clouds, Q_R , and a local critical point for T at z_T , which, respectively, correspond to the following equations:

$$\rho(z_B)w(z_B) = M_{\text{tot}}\Phi_B, \quad (22)$$

$$R_A\rho(z_B)T(z_B) = P_0, \quad (23)$$

$$I_+(z_B) = (\sigma T_S^4 - \sigma T(z_B)^4) e^{-\kappa(\rho(z_B), T(z_B))z_B} + \sigma T(z_B)^4, \quad (24)$$

$$F_C(z_B) = F_{C0}(\rho(z_B), T(z_B), T_S) e^{-bz_B}, \quad (25)$$

$$0 = F_O - \sigma T_S^4 + (I_-(z_B) - \sigma T(z_B)^4) e^{-\kappa(\rho(z_B), T(z_B))z_B} + \sigma T(z_B)^4 + I_S(z_B) e^{-k_S\rho(z_B)z_B} (1 - \alpha(T_S)) - F_{C0}(\rho(z_B), T(z_B), T_S), \quad (26)$$

$$0 = F_O - I_+(z_B) + I_+(z_B) + I_S(z_B) - I_S(z_B) e^{-k_S\rho(z_B)z_B} \alpha(T_S) - F_C(z_B) + k \frac{dT}{dz}(z_B) - \frac{1}{2} \rho(z_B)w(z_B)^3 - \rho(z_B)gz_B w(z_B) + \frac{1}{2} g M_{\text{tot}} \Phi_B z_B, \quad (27)$$

$$I_-(z_T) = 0, \quad (28)$$

$$I_S(z_T) = Q - Q_R, \quad (29)$$

$$\frac{dT}{dz}(z_T) = 0, \quad (30)$$

where F_{C0} is given by (14). The last three terms of Eqn. (27) may be simplified using Eqn. (22) so that they read

$$-\frac{M_{\text{tot}}\Phi_B w(z_B)^2}{2} - \frac{M_{\text{tot}}\Phi_B g z_B}{2}.$$

There are nine boundary conditions, but there are only eight dependent variables for which we have differential equations. The discrepancy is explained by the presence of T_S , which is an additional scalar unknown. The nine boundary conditions determine eight conditions for the differential equations as well as the value for T_S . One way of treating this is simply to extend the system of differential equations to include the equation

$$\frac{dT_S}{dz} = 0. \quad (31)$$

In addition, as described in Appendix A, to avoid numerical stiffness we take the limit as the heat conduction of air, k , tends to zero. This effectively reduces the size of the model by one dimension.

The model is nondimensionalized and put in standard form as detailed in Appendix A. The parameter values and their calibration to empirical data are provided in Appendix B.

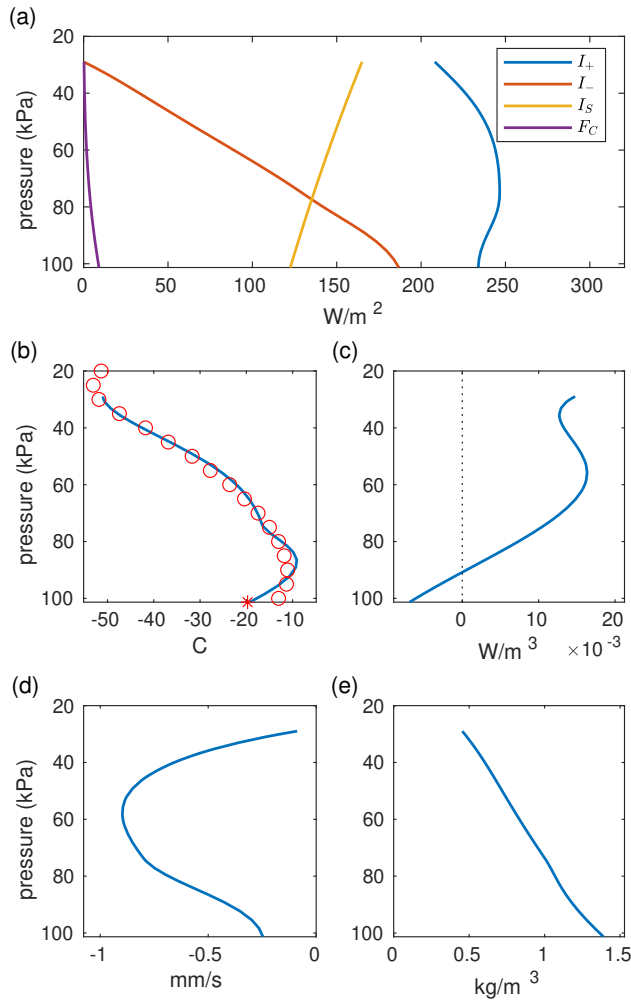


Figure 3. Results of the fully calibrated Arctic model at $\mu = 390$ ppm. The vertical axis in all plots is the pressure. (a) Energy transport via radiative terms, and latent and sensible heat. (b) Temperature of the atmosphere. The red circles are the data from Cronin and Jansen (2016). The red asterisk is the surface temperature, T_S . (c) Atmospheric heat transport, F_A . (d) Vertical wind speed. (e) Density.

3 Results

For the Arctic parameter values given in Appendix B and for a given CO_2 concentration, μ , the model can be solved numerically. We used MATLAB's builtin BVP solver "bvp5c" to solve individual instances of the (nondimensionalized) model, and AUTO for continuation calculations. The results of the model for $\mu = 390$ ppm are shown in Figure 3. The altitude dependence of the mass flux ϕ and the atmospheric meridional heat transport, F_A , were calibrated to an empirical Arctic temperature profile from Cronin and Jansen (2016) as detailed in Appendix B. From this figure we see that the model fits the temperature profile very well, with some discrepancy near the surface. The overall atmospheric heat transport, Fig. 3(c), indicates that, for

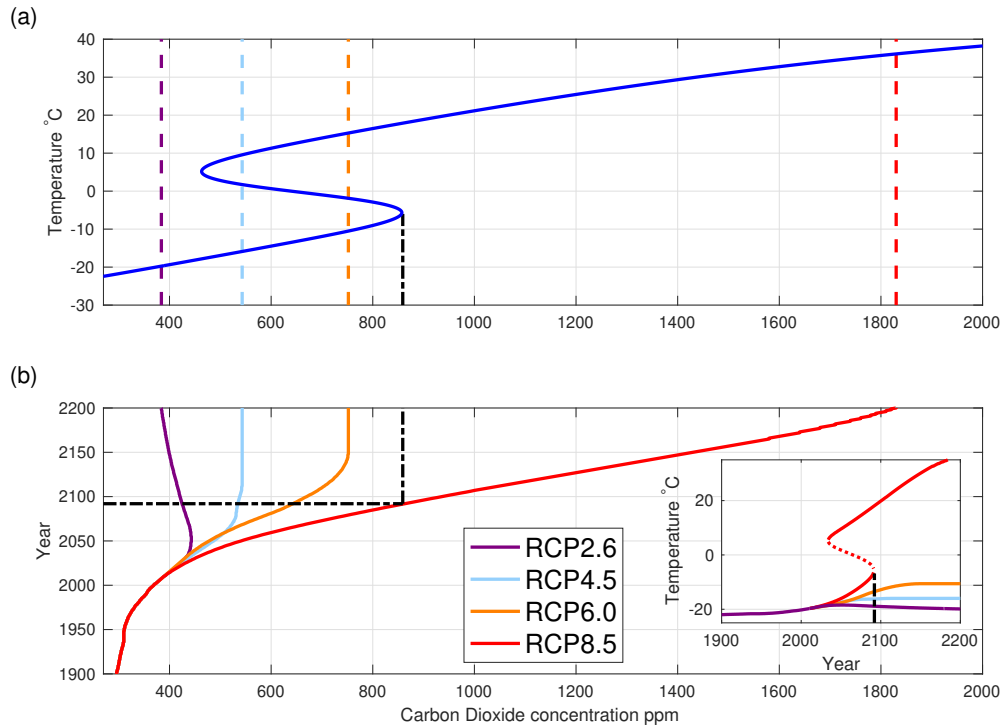


Figure 4. (a) Model surface temperature as a function of CO₂ concentration, μ . (b) CO₂ concentration levels for RCPs 2.6, 4.5, 6.0, and 8.5 (left to right). The dashed lines in panel (a) correspond to the CO₂ concentration levels for the four RCPs in the year 2200. The dashed-dotted line extends from the bifurcation point in panel (a) to RCP 8.5 in the lower panel indicating the bifurcation occurs in approximately the year 2100 for this scenario. The inset shows the predicted surface temperature as a function of the year for the four RCPs.

335 our model the upper half of the troposphere receives the most input of heat, while the bottom fifth actually has a net outward
 heat transport. This removal of heat near the bottom is likely the cause of the discrepancy between our model values and the
 Cronin and Jansen data. The negative values of F_A near the surface are due to our modelling choice for F_A . It is possible that
 alternative modelling formulations for F_A could yield a better fit to the Cronin and Jansen data, however, the forms that we
 tried (including the ones reported in Appendix B and a few others) did not improve upon the one used here. Nonetheless, given
 340 the relative simplicity of our model, we feel the fact that it can fit the data as well as it does is remarkable.

For $\mu = 390$, the surface temperature from the solved model is -19.7°C . Starting from this solution we numerically continued
 the solution to other values of μ , resulting in the S-shaped curve in the top panel of Fig. 4. Between approximately $\mu = 464$
 and $\mu = 859$ there are three solutions. The lower and upper solutions are stable cold and warm solutions, while the middle
 branch is unstable. Currently the Arctic is on the cold branch of this curve. The model predicts that as CO₂ levels rise, the
 345 equilibrium surface temperature in the Arctic will increase gradually at first. However, when μ exceeds 859 ppm, the model
 displays a saddle-node bifurcation where the stable cold solution is annihilated together with the unstable solution. At this

point the climate would rapidly approach the warm stable solution where surface temperatures are significantly higher. The temperatures on the warm branch may seem unreasonably high. Although our model includes specific details of some of the natural phenomena governing the Arctic climate behavior, and is calibrated with real data, it is primarily a qualitative model rather than quantitative. It is the fact that the model predicts the qualitative feature of a saddle-node bifurcation that is the important result, not the precise temperature of the model's warm solution. Part of the reason the model has relatively high temperatures on the warm branch is because it has constant values for F_O and F_A^{tot} . It is likely that should the Arctic's annually averaged temperature rise dramatically, both F_O and F_A^{tot} would be affected in a downward direction, which would reduce the warm equilibrium temperatures to some degree. Another important emergent feature of the model is that there is bistability for μ in the range [464, 859]. Thus, even though the Arctic may in the future be on the cold branch in this range, it is possible that a strong enough climate disturbance could push the climate out of the basin of attraction of the cold solution and into the basin of attraction of the warm solution. The necessary strength of such a disturbance decreases as one moves closer to the upper end of this range. Our model does not incorporate the seasonal variation of solar input to the Arctic, but rather uses an annually averaged value. Thus the equilibria in our model are annual averages. The seasonal variation of insolation would effectively result in oscillations around the annual average. These oscillations themselves may make a significant contribution to the "disturbance" needed to push the system out of the basin of attraction of the cold branch.

The International Panel on Climate Change (IPCC) has published various CO₂ emission scenarios for the future based on possible levels of global action to suppress such emissions (Intergovernmental Panel on Climate Change, 2013, Box TS.6), (van Vuuren et al., 2011). These scenarios are called Representative Concentration Pathways (RCPs) and are numbered based on the radiative forcing in the year 2100 due to anthropogenic emissions compared to the year 1750. The original four published RCPs are RCP 2.6, RCP 4.5, RCP 6.0, and RCP 8.5, representing strong mitigation (2.6) through "ignore the problem" (8.5) responses by world governments. Each RCP indicates likely CO₂ concentration levels in the atmosphere out to the year 2100. From 2100 to 2200, the scenarios assume a "constant composition commitment," which essentially freezes emission levels and eventually leads to a constant CO₂ level in the atmosphere for all but RCP 8.5. These RCPs are plotted in the lower panel of Fig. 4. The CO₂ levels for the four different pathways at the year 2200 are continued as dashed lines into the upper panel. It is clear from the figure that RCP 8.5 leads to CO₂ concentrations that far surpass the saddle-node bifurcation, whereas the other three RCPs do not. This result is in agreement with our simpler model (Kypke et al., 2020). Both RCP 6.0 and RCP 4.5 end at levels within the bistable range and indeed all the RCPs except RCP 2.6 are within that range from about the year 2050 onward. The black dashed-dotted line extending from the bifurcation in the upper panel to RCP 8.5 in the lower panel illustrates that RCP 8.5 reaches the bifurcation near the year 2092.

The curve of equilibria in the upper panel of Fig. 4 displays hysteresis: CO₂ levels rising past 859 ppm will cause a jump from the cold equilibrium state to the warm state, but a return to the cold state will not happen until CO₂ levels are brought below 464 ppm, where the saddle-node bifurcation of the warm equilibrium is located (left bend of the S-curve). If CO₂ levels follow a trajectory similar to RCP 8.5, Arctic climate may change drastically in less than 100 years, but a return to the current cold state may be essentially impossible for thousands of years afterward, assuming humankind can develop and implement the required technology to reduce atmospheric carbon levels sufficiently.

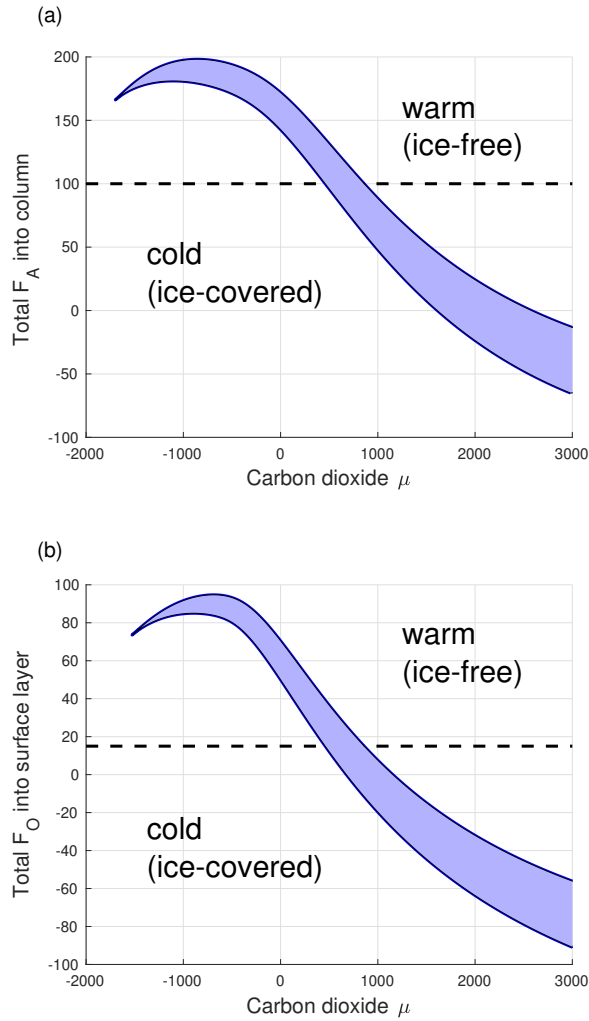


Figure 5. (a) Bifurcation diagram showing location of saddle-node bifurcations as F_A^{tot} and μ are varied. The dashed line indicates the value of F_A^{tot} used in the model. The shaded area between the curves is the region of bistability. Abrupt transition from a cold state to a warm state occurs on the right curve, while abrupt transition from a warm to a cold state occurs at the left curve. (b) Same except F_O is the varying parameter on the vertical axis rather than F_A^{tot} .

Total atmospheric heat transport, F_A^{tot} , and ocean heat transport, F_O , are inputs to the model that are empirically based and are not altered by the state of the model itself. We investigated the persistence of the saddle-node bifurcation and hysteresis in the presence of the uncertainty of these two values. Figure 5 shows two two-dimensional bifurcation diagrams plotting the
 385 locations of the saddle-node bifurcations in the model (the left and right bends of the S-curve in Fig. 4). The top panel of

Fig. 5 shows the bifurcations as both F_A^{tot} and μ vary. The bottom panel shows a similar diagram where F_O replaces F_A^{tot} . The right curve in each of the panels is a curve of saddle-node bifurcations corresponding to the right bend in the S-curve in Fig. 4; they are transitions from a cold state to a warm state. Similarly, the left curve in each of the panels is a curve of saddle-node bifurcations that correspond to the left bend in the S-curve, and are the transitions from a warm state to a cold state. The shaded area between the two curves is the region of bistability. The curves meet in a cusp bifurcation, but it is important to note that this only occurs when CO_2 levels are mathematically negative. Thus the two saddle-node bifurcations and the hysteresis phenomenon will be present for all physically possible CO_2 levels, and for all reasonable values of F_A^{tot} and F_O . These bifurcation curves also make it evident that a transition from a cold to a warm state occurs if either F_A^{tot} or F_O are increased, even if CO_2 levels are constant. If F_A^{tot} and F_O increase along with carbon dioxide levels in the near future, which seems reasonably likely, then the saddle-node bifurcation will occur at lower levels of CO_2 . For example, a ten percent increase of F_A^{tot} to 110 W m^{-2} from 100 would cause the saddle-node bifurcation to occur at about $\mu = 754 \text{ ppm}$, which would mean that both RCP 8.5 and RCP 6.0 would pass through the saddle-node bifurcation.

4 Conclusions

Although the model presented here is clearly a simplification of the climate, made possible by the near invariance of the vertical flow on the polar axis, we believe it captures some of the most important aspects relevant for Arctic climate change. The model predicts that if humanity keeps carbon emission levels close to RCP 6.0 or lower, then the Arctic will not likely undergo a sudden dramatic rise in annual average temperature. However, if carbon emissions are much worse than RCP 6.0, such a change is likely, and the cause is a saddle-node bifurcation of the stable cold equilibrium. Such a change would clearly have catastrophic effects on the Arctic environment leading to massive global effects. These results are in agreement with those of Årthun et al. (2021) who, from a study of various CMIP6 models of Arctic climate, predict that under a low emissions scenario sea ice loss will be seasonal, but for a high emissions scenario it will be year-round for all areas of the Arctic. Further, the hysteresis displayed by the model indicates that a change of this nature may be practically irreversible. Although some comfort may be taken that only the worst of the four carbon pathway scenarios ends in such a catastrophic change, the model shows that both RCP 6.0 and RCP 4.5 lie in the region of bistability from the year 2070 onward. In a bistable situation, external disturbances could cause the system to jump into the basin of attraction of the warm equilibrium, effectively bringing about the catastrophic change prior to the bifurcation. The model is too simple to allow for any reasonable measure of the likelihood of such an occurrence, but the important thing is that the model exhibits bistability in the parameter region where the system is likely to reside within 50 year's time. This bistability has been shown to persist regardless of the values of the two biggest uncertainties in the model, the atmospheric and ocean heat transport to the Arctic from the mid-latitudes.

Our model addresses the equilibrium state only and represents the Arctic temperature as an annual average. The real Arctic climate undergoes massive seasonal changes, which effectively means that the system is actually oscillating around the equilibrium temperatures of our model. Such temperature oscillations may be sufficient to effectively push a system located on a cold solution in the bistable regime to “above” the unstable solution and so into the basin of attraction of the warm solution.

Seasonal variations in Arctic temperatures and sea ice are studied in many works including Eisenman and Wettlaufer (2009) who argue that a tipping point to a year-round ice-free Arctic is not likely to occur while the Arctic is ice-covered for a significant portion of the year, but once the Arctic is seasonally ice-free for a sufficient number of months in a year, such a tipping point becomes more likely. A possible enhancement to our model would be to include seasonal solar variation and an ice model as in Eisenman and Wettlaufer (2009).

Appendix A: Model Derivation Details

This appendix provides details regarding the model. The model is written in a nondimensional form in Section A1, defining nondimensional variables and parameters. This system is then transformed in Section A2 to a standard form of nine first order ordinary differential equations and corresponding boundary conditions. This standard form makes it evident that the system is numerically stiff due to the fact that the thermal conductivity, k , of air is very small. To remove this stiffness, the limit as $k \rightarrow 0$ is applied, which reduces the system by one dimension. Section A3 discusses our choice of the functional forms for the dependence of relative humidity $\delta(z)$, atmospheric meridional heat transport $F_A(z)$, and mass flux $M_b(z)$, on altitude.

A1 Non-Dimensionalization

Define the non-dimensional variables

$$\begin{aligned} \hat{z} &= \frac{z - z_B}{z_T - z_B}, & y_1 &= \frac{w c_v \rho_0 T_R}{\sigma T_R^4}, & y_2 &= \frac{\rho}{\rho_0}, & y_3 &= \frac{I_+}{\sigma T_R^4}, & y_4 &= \frac{I_-}{\sigma T_R^4}, \\ y_5 &= \frac{I_S}{\sigma T_R^4}, & y_6 &= \frac{F_C}{\sigma T_R^4}, & y_7 &= \frac{T}{T_R}, & y_8 &= \left(\frac{z_T - z_B}{T_R} \right) \frac{dT}{dz}, & y_9 &= \frac{T_S}{T_R}, \end{aligned} \quad (\text{A1})$$

where P_0 (N m^{-2}) is the standard atmospheric pressure, T_R is the freezing point of water in Kelvin, $\rho_0 = P_0/(R_A T_R)$ is the density at standard pressure and freezing temperature, and $\sigma T_R^4/(c_v \rho_0 T_R)$ represents the vertical velocity required to move a parcel of air with standard density at freezing temperature so that the power transferred is equal to the radiative power for a black body at the same temperature. This comes to be about 1 mm/s.

Applying the change of variables to the troposphere BVP, (3)–(7), (10), (11), (31), we get

$$\frac{d}{d\hat{z}}(y_1 y_2) = D\phi(\hat{z}), \quad (\text{A2})$$

$$440 \quad H \frac{d}{d\hat{z}}(y_1^2 y_2) = -J \frac{d}{d\hat{z}}(y_2 y_7) - E y_2, \quad (\text{A3})$$

$$\frac{dy_3}{d\hat{z}} = -\hat{\kappa}(y_2, y_7) (y_3 - y_7^4), \quad (\text{A4})$$

$$\frac{dy_4}{d\hat{z}} = \hat{\kappa}(y_2, y_7) (y_4 - y_7^4), \quad (\text{A5})$$

$$\frac{dy_5}{d\hat{z}} = G_S y_2 y_5, \quad (\text{A6})$$

$$\frac{dy_6}{d\hat{z}} = -B_1 y_6, \quad (\text{A7})$$

$$445 \quad \frac{dy_7}{d\hat{z}} = y_8, \quad (\text{A8})$$

$$\begin{aligned} \epsilon \frac{dy_8}{d\hat{z}} = & \frac{d}{d\hat{z}} \left(\frac{H}{2} y_1^3 y_2 + E y_1 y_2 \hat{z} + y_1 y_2 y_7 \right) + J \frac{d}{d\hat{z}}(y_1 y_2 y_7) - DE \hat{z} \phi(\hat{z}) - D\phi(\hat{z}) y_7 \\ & - \hat{\kappa}(y_2, y_7) (y_3 + y_4 - 2y_7^4) - G_S y_2 y_5 - B_1 y_6 - \hat{F}_A(\hat{z}), \end{aligned} \quad (\text{A9})$$

$$\frac{dy_9}{d\hat{z}} = 0, \quad (\text{A10})$$

for $\hat{z} \in [0, 1]$. The boundary conditions, (22)–(30), become

$$450 \quad y_1(0) y_2(0) = D\Phi_B, \quad (\text{A11})$$

$$y_2(0) y_7(0) = 1, \quad (\text{A12})$$

$$y_3(0) = (y_9(0)^4 - y_7(0)^4) e^{-\hat{\kappa}(y_2(0), y_7(0))\zeta} + y_7(0)^4, \quad (\text{A13})$$

$$y_6(0) = \hat{F}_{C0}(y_2(0), y_7(0), y_9(0)) e^{-B_1 \zeta}, \quad (\text{A14})$$

$$0 = F - y_9(0)^4 + (y_4(0) - y_7(0)^4) e^{-\hat{\kappa}(y_2(0), y_7(0))\zeta} + y_7(0)^4$$

$$455 \quad + y_5(0) e^{-G_S y_2(0)\zeta} (1 - \hat{\alpha}(y_9(0))) - \hat{F}_{C0}(y_2(0), y_7(0), y_9(0)), \quad (\text{A15})$$

$$0 = F - y_3(0) + y_4(0) + y_5(0) - y_5(0) e^{-G_S y_2(0)\zeta} \hat{\alpha}(y_9(0)) - y_6(0) + \epsilon y_8(0)$$

$$- \frac{1}{2} H D \Phi_B y_1(0)^2 - \frac{1}{2} D E \Phi_B \zeta, \quad (\text{A16})$$

$$y_4(1) = 0, \quad (\text{A17})$$

$$y_5(1) = K_S, \quad (\text{A18})$$

$$460 \quad y_8(1) = 0, \quad (\text{A19})$$

where

$$\hat{F}_{C0}(y_2, y_7, y_9) = B_2 y_2 (y_9 - y_7) + \frac{B_3}{y_7} \left[\exp \left(G_{W1} \left(1 - \frac{1}{y_9} \right) \right) - \delta(0) \exp \left(G_{W1} \left(1 - \frac{1}{y_7} \right) \right) \right], \quad (\text{A20})$$

$$\hat{\kappa}(y_2, y_7) = \kappa(y_2 \rho_0, y_7 T_R)(z_T - z_B) = G_{Cl} + G_C \hat{\mu} y_2 + \frac{G_{W2} \delta(\hat{z})}{y_7} \exp \left(G_{W1} \left(1 - \frac{1}{y_7} \right) \right), \quad (\text{A21})$$

$$\hat{F}_A(\hat{z}) = \frac{(z_T - z_B) F_A(\hat{z}(z_T - z_B) + z_B)}{\sigma T_R^4}, \quad (\text{A22})$$

$$465 \quad \hat{\alpha}(y_9) = \frac{1}{2} \left[(\alpha_w + \alpha_c) + (\alpha_w - \alpha_c) \tanh \left(\frac{y_9 - 1}{\omega} \right) \right], \quad (\text{A23})$$

are nondimensionalized functions describing the sensible/latent heat flux from the surface, the absorption of long wave radiation due to greenhouse gases, the atmospheric heat transport, and the surface albedo; and α_w , α_c , ω , Φ_B , and

$$\begin{aligned} B_1 &= b(z_T - z_B), & B_2 &= \frac{c_v \rho_0 C_D U}{\sigma T_R^3}, & B_3 &= \frac{\rho_W^{sat}(T_R) C_D U L_v}{\sigma T_R^4}, \\ D &= \frac{M_{tot} c_v}{\sigma T_R^3}, & E &= \frac{g(z_T - z_B)}{c_v T_R}, & F &= \frac{F_O}{\sigma T_R^4}, \\ G_{Cl} &= k_{Cl}(z_T - z_B), & G_C &= k_C \frac{M_{CO_2}}{M_A} \rho_0 (z_T - z_B), & G_{W1} &= \frac{L_v}{R_W T_R}, \\ G_{W2} &= k_W \rho_W^{sat}(T_R)(z_T - z_B), & G_S &= k_S \rho_0 (z_T - z_B), & H &= \frac{\sigma^2 T_R^5}{c_v^3 \rho_0^2}, \\ J &= \frac{R_A}{c_v}, & K_S &= \frac{I_S^{TP} - Q_R}{\sigma T_R^4}, & \epsilon &= \frac{k}{(z_T - z_B) \sigma T_R^3}, \\ \hat{\mu} &= \frac{\mu}{10^6}, & \zeta &= \frac{z_B}{z_T - z_B}, \end{aligned} \quad (\text{A24})$$

470 are nondimensional constants. Values of the physical parameters are given in the tables in Appendix B. It turns out that all of the above nondimensional constants are close to order one (range 0.04 to 26) except $\zeta = 5.6 \times 10^{-3}$, which is the relative boundary layer thickness, $\epsilon = 2.35 \times 10^{-6}$, which is the nondimensional conductance, and $H = 7.96 \times 10^{-12}$, which is a factor on the kinetic energy term. As shown in the next section, the fact that ϵ is small causes system stiffness.

A2 Standard Form and Vanishing Conduction Limit

The BVP given by (A2)–(A19) contains derivatives of products of some of the variables. It can be put in standard ODE form 475 via algebraic manipulations. First we expand the derivative on the left side of (A3) and use (A2) to simplify it:

$$\frac{d}{d\hat{z}}(y_1^2 y_2) = y_1 \frac{d}{d\hat{z}}(y_1 y_2) + (y_1 y_2) \frac{dy_1}{d\hat{z}} = y_1 D \phi(\hat{z}) + y_1 y_2 \frac{dy_1}{d\hat{z}}.$$

Thus (A3) may be replaced with

$$H y_1 y_2 \frac{dy_1}{d\hat{z}} + J y_7 \frac{dy_2}{d\hat{z}} = -J y_2 y_8 - H D y_1 \phi(\hat{z}) - E y_2, \quad (\text{A25})$$

480 where we have used (A8) to replace the derivative of y_7 . We can also expand the derivative in (A2) to get the equivalent equation

$$y_2 \frac{dy_1}{d\hat{z}} + y_1 \frac{dy_2}{d\hat{z}} = D \phi(\hat{z}). \quad (\text{A26})$$

Equations (A25) and (A26) are a linear system in $\frac{dy_1}{d\hat{z}}$ and $\frac{dy_2}{d\hat{z}}$, namely

$$\begin{bmatrix} Hy_1y_2 & Jy_7 \\ y_2 & y_1 \end{bmatrix} \begin{bmatrix} \frac{dy_1}{d\hat{z}} \\ \frac{dy_2}{d\hat{z}} \end{bmatrix} = \begin{bmatrix} -Jy_2y_8 - HDy_1\phi(\hat{z}) - Ey_2 \\ D\phi(\hat{z}) \end{bmatrix}.$$

Solving this yields

$$485 \quad \frac{dy_1}{d\hat{z}} = \frac{(Jy_7 + Hy_1^2)D\phi(\hat{z}) + y_1y_2(E + Jy_8)}{y_2(Jy_7 - Hy_1^2)}, \quad (\text{A27})$$

$$\frac{dy_2}{d\hat{z}} = -\frac{2Hy_1D\phi(\hat{z}) + y_2(E + Jy_8)}{Jy_7 - Hy_1^2}. \quad (\text{A28})$$

Now expanding the derivatives on the right hand side of (A7) and simplifying we obtain

$$\begin{aligned} \epsilon \frac{dy_8}{d\hat{z}} &= \left[H \frac{y_1^2}{2} + \frac{J^2y_7^2 + H^2y_1^4}{Jy_7 - Hy_1^2} \right] D\phi(\hat{z}) + (J + 1)y_1y_2y_8 + Jy_1y_2 \frac{Ey_7 + Hy_1^2y_8}{Jy_7 - Hy_1^2} \\ &\quad - \hat{\kappa}(y_2, y_7)y_2(y_3 + y_4 - 2y_7^4) - G_Sy_2y_5 - B_1y_6 - \hat{F}_A(\hat{z}). \end{aligned} \quad (\text{A29})$$

490 In summary the BVP for the troposphere in standard form is given by

$$\frac{dy_1}{d\hat{z}} = \frac{(Jy_7 + Hy_1^2)D\phi(\hat{z}) + y_1y_2(E + Jy_8)}{y_2(Jy_7 - Hy_1^2)}, \quad (\text{A30})$$

$$\frac{dy_2}{d\hat{z}} = -\frac{2Hy_1D\phi(\hat{z}) + y_2(E + Jy_8)}{Jy_7 - Hy_1^2} \quad (\text{A31})$$

$$\frac{dy_3}{d\hat{z}} = -\hat{\kappa}(y_2, y_7)(y_3 - y_7^4), \quad (\text{A32})$$

$$\frac{dy_4}{d\hat{z}} = \hat{\kappa}(y_2, y_7)(y_4 - y_7^4), \quad (\text{A33})$$

$$495 \quad \frac{dy_5}{d\hat{z}} = G_Sy_2y_5, \quad (\text{A34})$$

$$\frac{dy_6}{d\hat{z}} = -B_1y_6, \quad (\text{A35})$$

$$\frac{dy_7}{d\hat{z}} = y_8, \quad (\text{A36})$$

$$\begin{aligned} \epsilon \frac{dy_8}{d\hat{z}} &= \left[H \frac{y_1^2}{2} + \frac{J^2y_7^2 + H^2y_1^4}{Jy_7 - Hy_1^2} \right] D\phi(\hat{z}) + (J + 1)y_1y_2y_8 + Jy_1y_2 \frac{Ey_7 + Hy_1^2y_8}{Jy_7 - Hy_1^2} \\ &\quad - \hat{\kappa}(y_2, y_7)(y_3 + y_4 - 2y_7^4) - G_Sy_2y_5 - B_1y_6 - \hat{F}_A(\hat{z}), \end{aligned} \quad (\text{A37})$$

$$500 \quad \frac{dy_9}{d\hat{z}} = 0, \quad (\text{A38})$$

with boundary conditions

$$y_1(0)y_2(0) = D\Phi_B, \quad (\text{A39})$$

$$y_2(0)y_7(0) = 1, \quad (\text{A40})$$

$$y_3(0) = (y_9(0)^4 - y_7(0)^4) e^{-\hat{\kappa}(y_2(0), y_7(0))\zeta} + y_7(0)^4, \quad (\text{A41})$$

$$505 \quad y_6(0) = \hat{F}_{C0}(y_2(0), y_7(0), y_9(0)) e^{-B_1\zeta}, \quad (\text{A42})$$

$$0 = F - y_9(0)^4 + (y_4(0) - y_7(0)^4) e^{-\hat{\kappa}(y_2(0), y_7(0))\zeta} + y_7(0)^4 \\ + y_5(0) e^{-G_S y_2(0)\zeta} (1 - \hat{\alpha}(y_9(0))) - \hat{F}_{C0}(y_2(0), y_7(0), y_9(0)), \quad (\text{A43})$$

$$0 = F - y_3(0) + y_4(0) + y_5(0) - y_5(0) e^{-G_S y_2(0)\zeta} \hat{\alpha}(y_9(0)) - y_6(0) + \epsilon y_8(0) \\ - \frac{1}{2} H D \Phi_B y_1(0)^2 - \frac{1}{2} D E \Phi_B \zeta, \quad (\text{A44})$$

$$510 \quad y_4(1) = 0, \quad (\text{A45})$$

$$y_5(1) = K_S, \quad (\text{A46})$$

$$y_8(1) = 0. \quad (\text{A47})$$

As mentioned above, of the nondimensional constants all are close to order one except $\zeta = 5.6 \times 10^{-3}$, $H = 7.96 \times 10^{-12}$, and $\epsilon = 2.3 \times 10^{-6}$ (with $z_T = 9000$). The constant ζ only appears in the boundary conditions. The constant H occurs in the system only in summations with other non-derivative terms, hence the fact it is small only means that those terms contribute little. It does not cause stiffness. The relatively small value of the constant ϵ however, does cause stiffness in Eqn. (A37) of the system. Because ϵ is small, the variable y_8 , which is the dimensionless rate of vertical temperature change dT/dz , will approach the y_8 -nullcline very rapidly (which means in a very short z -distance from either boundary). To simplify numerical computation we take the limit as ϵ goes to zero, which is equivalent to saying that conduction is negligible. In this limit, (A37) is an algebraic expression from which we can isolate y_8 :

$$y_8(\hat{z}) = - \left[\left(H \frac{y_1^2}{2} + \frac{J^2 y_7^2 + H^2 y_1^4}{J y_7 - H y_1^2} \right) D \phi(\hat{z}) + \frac{E J y_1 y_2 y_7}{J y_7 - H y_1^2} - G_S y_2 y_5 \right. \\ \left. - \hat{\kappa}(y_2, y_7) (y_3 + y_4 - 2y_7^4) - B y_6 - \hat{F}_A(\hat{z}) \right] \left[(J + 1) y_1 y_2 + \frac{J H y_1^3 y_2}{J y_7 - H y_1^2} \right]^{-1}. \quad (\text{A48})$$

With this expression for y_8 the system is reduced by eliminating Eqn. (A37) and boundary condition (A47). Since y_1 , the dimensionless vertical velocity, is a factor in the denominator it is necessary that y_1 be nonzero throughout the atmosphere in order not to introduce a singularity. For this reason we select $\Phi_T > 0$, $\Phi_B < 0$ and ϕ such that it is nonnegative in the upper atmosphere and nonpositive in the lower atmosphere.

A3 Modelling Choices for Functional Forms

In this section we describe the various functional forms that we used for relative humidity, atmospheric heat transport, and mass flux. For the latter two, several different forms were tried and these are detailed below. Calibration to empirical data, described in Appendix B, was used to select specific functional forms for the heat transport and mass flux.

A3.1 Relative humidity

The relative humidity is modelled as a linear function decreasing with altitude from a higher surface value, δ_B , to a lower value at the tropopause, δ_T . Specifically,

$$\delta(\hat{z}) = \delta_B(1 - \hat{z}) + \delta_T\hat{z}. \quad (\text{A49})$$

A3.2 Atmospheric heat transport

Atmospheric heat transport is primarily due to large scale turbulent mixing of the column with its environment. This mixing is not modelled explicitly, but, instead, is incorporated into the model via the function $F_A(z)$, which represents the thermal energy supplied to the column by turbulent mixing. The integral of $F_A(z)$ over the atmosphere thickness represents the total atmospheric heat transport in/out of the system. So given a set amount of such energy, F_A^{tot} we have

$$F_A^{\text{tot}} = \int_{z_B}^{z_T} F_A(z) dz = \sigma T_R^4 \int_0^1 \hat{F}_A(\hat{z}) d\hat{z}. \quad (\text{A50})$$

This provides one restriction on the function $F_A(z)$, but its precise form is a modelling choice. However, the exact form must be chosen with care, because with certain choices of $F_A(z)$, the boundary conditions can only be satisfied with unrealistic solutions. In particular, if the values of $F_A(z)$, near the tropopause (z near z_T , that is, \hat{z} near 1) are too small, then the temperature drops precipitously toward absolute zero; if they are too big, the temperature turns around and climbs rapidly.

Thus, in order to automate an appropriate choice of $F_A(z)$, we have proceeded as follows:

1. Choose $\hat{F}_A(1)$ such that the temperature gradient at the tropopause is zero, that is, re-impose boundary condition (A47). Thus use (A48) with $y_8(1) = 0$, to solve for $\hat{F}_A(1)$. Let \hat{F}_{A1} denote this value of $\hat{F}_A(1)$, and let F_{A1} denote the corresponding dimensional value of $F_A(z_T)$, that is, $F_{A1} = \hat{F}_{A1}\sigma T_R^4/(z_T - z_B)$.

2. Assume that $F_A(z)$ is of the form

$$F_A(z) = F_{Ab}(z) + \frac{F_A^{\text{tot}}}{z_T - z_B} \psi\left(\frac{z - z_B}{z_T - z_B}\right), \quad (\text{A51})$$

where the base function $F_{Ab}(z)$ is a linear function passing through zero at the midpoint of the atmosphere and equal to F_{A1} at z_T , that is,

$$F_{Ab}(z) = F_{A1} \frac{2z - (z_T + z_B)}{z_T - z_B}, \quad \iff \quad \hat{F}_{Ab}(\hat{z}) = \hat{F}_{A1}(2\hat{z} - 1).$$

555 This base portion of F_A contributes no net heat to the column, it is simply a factor that essentially moves heat around in the column in order to assure the temperature gradient at the top is zero.

The remaining portion of F_A is the actual atmospheric heat transport entering the column from outside. We assume that $\psi(\hat{z})$ satisfies

$$\psi(1) = 0 \quad \text{and} \quad \int_0^1 \psi(x) dx = 1,$$

560 so that the value of F_A is not altered at the tropopause, and so that F_A^{tot} (W m^{-2}) represents the total energy flux of atmospheric heat transport entering the column. Thus the non-dimensional function is

$$\hat{F}_A(\hat{z}) = \hat{F}_{A1}(2\hat{z} - 1) + \hat{F}_A^{\text{tot}}\psi(\hat{z}),$$

$$\text{where } \hat{F}_A^{\text{tot}} = \frac{F_A^{\text{tot}}}{\sigma T_R^4}.$$

565 3. As a numerical issue, since the boundary condition value \hat{F}_{A1} is needed in the computation of the vector field, and since the MATLAB solver we are using does not have a way of making boundary condition information available to the vector field computation function, we circumvented this issue by adding another variable to the problem y_{10} , with differential equation $\frac{dy_{10}}{d\hat{z}} = 0$ and boundary condition $y_{10}(1) = \hat{F}_{A1}$.

4. Choice of the functional form of $\psi(\hat{z})$ is somewhat open, we tested the following two forms:

$$\psi(\hat{z}) = g_1(1 - \hat{z}, L_\psi), \quad (\text{A52})$$

$$\psi(\hat{z}) = g_2(1 - \hat{z}, L_\psi), \quad (\text{A53})$$

570 where the functions g_1 and g_2 are defined as

$$g_1(x, L) = \frac{L\pi}{1 - \cos(L\pi)} \sin(L\pi x),$$

$$g_2(x, L) = \frac{2L\pi}{2L\pi - \sin(2L\pi)} (1 - \cos(2L\pi x)), \quad (\text{A54})$$

and where L is a parameter in $(0, 1]$ free to be chosen. (The functions g_1 and g_2 will also be utilized in the modelling of $\phi(\hat{z})$.) The primary difference between these two forms is that the first has a non-zero slope at $x = 0$, while the latter has a zero slope there.

575 A3.3 Mass flux

The function ϕ dictates the mass flux across the vertical boundary (negative outward), and, along with the fluxes Φ_B and Φ_T across the bottom and top of the column, drives the vertical air movement in the column. The only general restrictions on these fluxes are given by (2). To model the situation in the Arctic, we want a downward flow of air with a vertical wind speed, w , on the order of 1 mm s^{-1} in the column. A reasonable assumption at the tropopause would be to set $w = 0$, however, since

580 our model has a singularity when $w = 0$, we impose a small wind speed at the tropopause by ensuring Φ_T is positive. At the surface boundary layer we impose $\Phi_B \leq 0$. Further, to simplify matters and to ensure a downward flow throughout the column, we assume that

$$\begin{aligned}\phi(\hat{z}) &\leq 0, & \text{if } \hat{z} \in [0, z_c), \\ \phi(\hat{z}) &\geq 0, & \text{if } \hat{z} \in [z_c, 1],\end{aligned}$$

585 where z_c is some point in $[0, 1]$. The following forms were tested for $\phi(\hat{z})$:

$$\phi(\hat{z}) = \begin{cases} \frac{-1-\Phi_B}{z_c} & \text{if } \hat{z} \in [0, z_c), \\ \frac{1-\Phi_T}{1-z_c} & \text{if } \hat{z} \in [z_c, 1], \end{cases} \quad \text{piecewise constant,} \quad (\text{A55})$$

$$\phi(\hat{z}) = \begin{cases} \frac{2(-1-\Phi_B)}{z_c^2} (z_c - \hat{z}) & \text{if } \hat{z} \in [0, z_c), \\ \frac{2(1-\Phi_T)}{(1-z_c)^2} (\hat{z} - z_c) & \text{if } \hat{z} \in [z_c, 1], \end{cases} \quad \text{piecewise linear,} \quad (\text{A56})$$

$$\phi(\hat{z}) = \begin{cases} \frac{-1-\Phi_B}{z_c} g_1 \left(1 - \frac{\hat{z}}{z_c}, L_{\phi B}\right), & \text{if } \hat{z} \in [0, z_c), \\ \frac{1-\Phi_T}{1-z_c} g_1 \left(\frac{\hat{z}-z_c}{1-z_c}, L_{\phi T}\right), & \text{if } \hat{z} \in [z_c, 1], \end{cases} \quad \text{piecewise sine,} \quad (\text{A57})$$

$$\phi(\hat{z}) = \begin{cases} \frac{-1-\Phi_B}{z_c} g_2 \left(1 - \frac{\hat{z}}{z_c}, L_{\phi B}\right), & \text{if } \hat{z} \in [0, z_c), \\ \frac{1-\Phi_T}{1-z_c} g_2 \left(\frac{\hat{z}-z_c}{1-z_c}, L_{\phi T}\right), & \text{if } \hat{z} \in [z_c, 1], \end{cases} \quad \text{piecewise cosine,} \quad (\text{A58})$$

590 where g_1 and g_2 are defined by (A54). (In the case that $z_c = 0$ or $z_c = 1$, it is understood that only the non-empty interval for ϕ in the above definitions is used and that it is closed at both ends.)

Appendix B: Model Parameters and Calibration

This appendix lists the parameter values used in the model and discusses how some of them were calibrated to empirical data. Section B2 gives the calculation of the average annual insolation for the Earth north of 70° N.

595 B1 Parameter Values and Calibration

Values of the model parameters are given in Tables B1 and B2. The parameters in Table B1 are physical constants. The parameters in Table B2 are those that have been assigned from empirical data and knowledge, or whose values have resulted from fitting the model to empirical data. The model was applied to the globally averaged situation for the purposes of calibration of some parameters (see below) and then also applied to the Arctic.

600 Here we provide justification and explanation of our choice of parameter values in Table B2. The height of the boundary layer was set to $z_B = 50$ m. The model is not very sensitive to this parameter. The height of the tropopause is about 9 km at the poles and 17 km at the equator so we used the lower value for the Arctic, and a middle value of 14 km for the global average. The globally averaged insolation, the atmospheric solar reflection, and the average surface albedo are all obtained from Wild

Table B1. Physical constants used in the model.

Parameter	Symbol	Value	Units
reference temperature	T_R	273.15	K
Stefan-Boltzmann constant	σ	5.67037×10^{-8}	$\text{W m}^{-2} \text{K}^{-4}$
latent heat of vapourization for water	L_v	2.2558×10^6	$\text{m}^2 \text{s}^{-2}$
specific heat capacity of dry air at T_R	c_v	716.4	$\text{J K}^{-1} \text{kg}^{-1}$
saturated vapour density at T_R	$\rho_w^{sat}(T_R)$	4.849×10^{-3}	kg m^{-3}
universal gas constant	R	8.31446	$\text{J mol}^{-1} \text{K}^{-1}$
molecular weight of CO_2	M_{CO_2}	4.4009×10^{-2}	kg mol^{-1}
molecular weight of dry air	M_A	2.89644×10^{-2}	kg mol^{-1}
molecular weight of water	M_W	1.80153×10^{-2}	kg mol^{-1}
gas constant for air	$R_A = R/M_A$	287.058	$\text{m}^2 \text{s}^{-2} \text{K}^{-1}$
specific gas constant of water vapour	$R_W = R/M_W$	461.4	$\text{m}^2 \text{s}^{-2} \text{K}^{-1}$
pressure at surface	P_0	101,325	Pa
standard dry density at T_R	$\rho_0 = P_0/(R_A T_R)$	1.29225	kg m^{-3}
conductivity of air	k	24.35×10^{-3}	$\text{W m}^{-1} \text{K}^{-1}$
gravitational acceleration	g	9.8	m s^{-2}

et al. (2013). For the Arctic, the insolation is the annual average for the region north of 70°N , the calculation of which is shown
605 in Section B2. For the average global situation there is no ocean or atmospheric heat transport; the corresponding values for
the Arctic come from Mayer et al. (2019); Serreze et al. (2007). The Arctic atmospheric reflection and surface albedo come
from Kalnay et al. (1996); National Centers for Environmental Protection/National Weather Service/NOAA/U.S. Department
of Commerce (accessed Nov., 2021); Mayer et al. (2019). Relative humidity is low at the top of the troposphere so in both the
global and Arctic cases was set to 10%. The surface relative humidity was set to 75% for the global average and 70% for the
610 Arctic. The surface horizontal wind speed, U , which is a factor in the sensible and latent heat transport from the surface, was
set to 10 m s^{-1} . The exact value is not too important since U always appears multiplied by the drag coefficient factor C_D ,
which we calibrate to data below. For the global model it is appropriate to assume that the average vertical wind speed is zero,
but since the model requires a nonzero wind speed we set $M_{\text{tot}} = 2.0 \times 10^{-6} \text{ kg m}^{-2} \text{ s}^{-1}$, and we set $\Phi_B = -1$, $\Phi_T = 0.2$,
and assumed $\phi(\hat{z})$ is given by (A57) with $z_c = 0$ and $L_{\phi T} = 1$. ($L_{\phi B}$ is irrelevant since $z_c = 0$.) These settings make the wind
615 speed relatively constant and on the order of 10^{-3} mm/s , which is far enough away from the singularity to avoid convergence
issues, but is small enough so all of the convection-related terms in the model become negligible. Mass flux for the Arctic
situation was set with trial and error to $8.0 \times 10^{-4} \text{ kg m}^{-2} \text{ s}^{-1}$, which gave vertical wind speeds in the column on the order of
0.5 mm/s. Since $F_A^{\text{tot}} = 0$ for the global case, the form of ψ and therefore also the parameter L_ψ are not relevant.

Table B2. Other model parameters. Some parameters are geographically dependent and have different values for the global and Arctic situations. An empty value in the “Arctic” column indicates the global value is used in both cases. Values that were fitted by the calibration steps described in the text are indicated in bold.

Parameter	Symbol	Global Value	Arctic Value	Units	Equation(s)
height of boundary layer	z_B	50		m	(3), (8), (11), (17)–(27), (A1), (A21), (A22), (A24), (A50), (A51)
top of troposphere	z_T	14,000	9,000	m	(3), (8), (11), (28)–(30), (A1), (A21), (A22), (A24), (A50), (A51)
insolation	Q	340	185	W m^{-2}	(29)
atmospheric solar reflection	Q_R	76	20	W m^{-2}	(29), (A24)
ocean heat transport	F_O	0	15	W m^{-2}	(15), (17), (26), (27), (A24)
total atmospheric heat transport	F_A^{tot}	0	100	W m^{-2}	(A22), (A50), (A51)
relative humidity at tropopause	δ_T	0.1		—	(A49)
relative humidity at bottom	δ_B	0.75	0.7	—	(A49)
surface horizontal wind speed	U	10		m s^{-1}	(14), (A24)
drag coefficient	C_D	3.180×10^{-3}		—	(14), (A24)
shortwave absorption	k_S	4.035×10^{-5}		$\text{m}^2 \text{kg}^{-1}$	(7), (11), (21), (26), (27), (A24)
longwave CO_2 absorption	k_C	0.1552		$\text{m}^2 \text{kg}^{-1}$	(8), (A24)
longwave H_2O absorption	k_W	0.04969		$\text{s}^2 \text{kg}^{-1}$	(8), (A24)
longwave cloud absorption	k_{Cl}	7.020×10^{-5}		m^{-1}	(8), (A24)
latent & sensible heat decay	b	4.153×10^{-4}		m^{-1}	(10), (11), (19), (25), (A24)
mass flux	M_{tot}	2.0×10^{-6}	8.0×10^{-4}	$\text{kg m}^{-2} \text{s}^{-1}$	(3), (11), (17), (27)
relative mass flux through top	Φ_T	0.2	0.05	—	(2), (A55)–(A58)
relative mass flux through bottom	Φ_B	−1	−0.4287	—	(2), (17), (22), (27), (A11), (A16), (A39), (A44), (A55)–(A58)
relative zero location of ϕ	z_c	0	0.2708	—	(A55)–(A58)
ϕ length scale bottom	$L_{\phi B}$	—	1.000	—	(A57), (A58)
ϕ length scale top	$L_{\phi T}$	1	0.5727	—	(A57), (A58)
ψ length scale	L_{ψ}	—	0.7744	—	(A52), (A53)
CO_2 level	μ	390*	390*	molar ppm	(8), (A24)
cold surface albedo	α_c^\dagger	24/185	0.667	—	(15), (16), (17), (26), (27), (A23)
warm surface albedo	α_w^\dagger	24/185	0.1	—	(15), (16), (17), (26), (27), (A23)
albedo transition steepness	ω	—	0.01942	—	(15), (16), (17), (26), (27), (A23)

*Values used for calibration only.

†For the purposes of calibrating the model to global data, both α_c and α_w were set to the empirical value of 24/185. When calibrating the model’s mass flux and atmospheric heat transport parameters to Arctic data, both α_c and α_w were set to the empirical value of 2/3.

Table B3. Global Average Energy Fluxes (W m^{-2}) from Wild et al. (2013) and contribution fractions for absorption from Schmidt et al. (2010).

	$I_+(z_T)$	$I_+(0)$	$I_-(0)$	$I_S(0)$	$F_C(0)$	C_{contrib}	Cl_{contrib}	W_{contrib}
Data	239	397	342	185	105	0.25	0.25	0.50
Model	239.7	397.4	341.7	184.9	105.2	0.2332	0.2130	0.5538

Calibration of the other model parameters was done in two steps. First, the absorption coefficients, k_S , k_C , k_W , and k_{Cl} , the
620 decay for sensible and latent heat transport, b , and the drag coefficient, C_D , (which is a multiplicative factor of both B_2 and B_3) were calibrated using global average energy fluxes obtained from Wild et al. (2013). In addition, this calibration attempted to match estimates from Schmidt et al. (2010) indicating that 25% of absorption is due to carbon dioxide, 25% due to clouds and 50% due to water vapour. These fractions are determined from the model via

$$\begin{aligned}
 Tot &= \int_0^1 \hat{\kappa}(y_2, y_7) y_4 d\hat{z}, & C_{\text{contrib}} &= \frac{1}{Tot} \int_0^1 G_C \mu y_2 y_4 d\hat{z}, \\
 625 \quad Cl_{\text{contrib}} &= \frac{1}{Tot} \int_0^1 G_{Cl} y_4 d\hat{z}, & W_{\text{contrib}} &= \frac{1}{Tot} \int_0^1 \frac{G_W \delta}{y_7} e^{G_{W1}(1-1/y_7)} y_4 d\hat{z}.
 \end{aligned}$$

The relevant data are given in Table B3.

Using these parameter settings we minimized the sum of squares of the differences between the data from Table B3 (after
nondimensionalization) with the model outputs allowing the parameters k_S , k_C , k_W , k_{Cl} , b , and C_D to vary. In the mini-
630 mization calculation, the terms associated with the contribution values (last three columns of Table B3) were given a heuristic weight of 0.01, since these values are less reliable than the other data. The resulting calibrated values for these parameters are given in Table B2; the results of the fitting are given in Table B3 and Figure B1. As can be seen in Table B3, the minimization achieved very good agreement with the globally averaged data.

Using the calibrated values for k_S , k_C , k_W , k_{Cl} , b , and C_D obtained from the first step, the second calibration step was to
select the parameters for the functions $\phi(\hat{z})$ and $\psi(\hat{z})$ to attempt to match the annual temperature profile for the Arctic from
635 Cronin and Jansen (2016, Fig. 1). Data from that figure of their paper is reproduced in Table B4. The Arctic values of the geographic-dependent parameters from Table B2 were used. Using each of the four forms for ϕ , given by Eqns. (A55)–(A58), and each of the two forms for F_A , given by Eqns. (A52)–(A53), the sum of the square of the differences between the model and the data in Table B4 was minimized by allowing the parameters z_c , Φ_B , $L_{\phi T}$, $L_{\phi B}$, and L_ψ to vary. The fit quality is shown in Figure B2. From this figure it is evident that the first two forms for ϕ , namely equations (A55) and (A56) do not give adequate
640 fits. The other two forms for ϕ are similar, and both forms of ψ only give small changes. The best fit is the third form for ϕ and the second form for ψ , that is, equations (A57) and (A53) and so these forms were chosen for the model. The calibrated parameters for these forms of ϕ and ψ are given in Table B2 and the corresponding functions ϕ and ψ are shown in Figure B3.

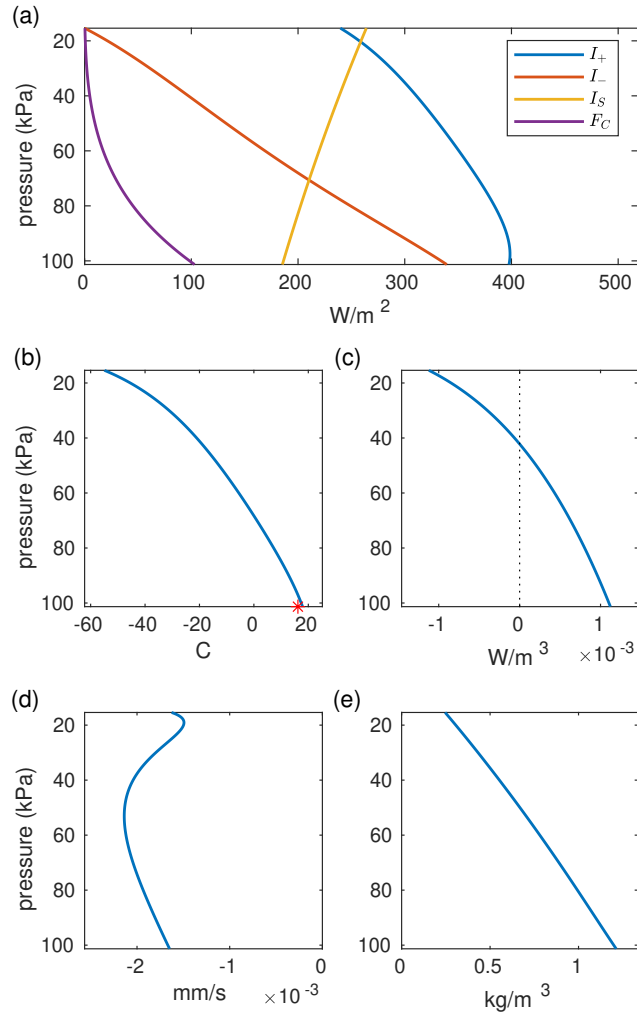


Figure B1. Model calibrated to global average values. The vertical axis in all plots is the pressure. (a) Energy transport via radiative terms, and latent and sensible heat. (b) Atmospheric temperature. The red asterisk marks the surface temperature, T_S . (c) Atmospheric heat transport, F_A . (d) Vertical wind speed. (e) Density.

Table B4. Annual Arctic temperature data from Figure 1 of Cronin and Jansen (2016).

Pressure (kPa)	100	95	90	85	80	75	70	65	60	55
Temperature (K)	260.1	261.8	262.1	261.3	260.1	258.1	255.7	252.7	249.5	245.4
Pressure (kPa)	50	45	40	35	30	25	20	15	10	
Temperature (K)	241.4	236.3	231.4	225.7	221.3	220.1	221.8	222.5	221.5	

For the above calibrations, the albedo values α_c and α_w in (A23) were set at the same constant value (24/185 for calibration to global data, 2/3 for calibration to Arctic data) in order to ensure the empirical albedo value was matched regardless of
645 the surface temperature. Now that the calibrations are complete, we wish to apply the model to the Arctic using different values of the carbon dioxide level. As such, we must choose values of α_c , α_w , and ω that allow for the albedo to change appropriately with surface temperature. The empirical albedo of 2/3, which was used in the Arctic calibration, resulted in a surface temperature of $T_S = 253.4$ K. Since this temperature is well below freezing, the albedo at this temperature should be near the maximum albedo, so α_c was set to 0.667. Second, α_w was set to 0.1, corresponding to the fact that north of 70°
650 latitude, the Earth is mostly ocean covered (albedo 0.06 for open water) and partly land (albedo 0.1–0.4). The value of ω was then calculated from (16) so that $\alpha(253.4) = 2/3$. This resulted in $\omega = 0.01942$.

B2 Insolation

This section presents the calculation of the insolation used in the Arctic model, where, in particular, the insolation is taken as an annual average over the region north of 70° N.

Select a Cartesian coordinate system (x, y, z) for the solar system with the sun at the origin, with the z -axis perpendicular to the Earth's orbital plane, and with the positive x -axis defined by the direction in the orbital plane from the sun to the centre of the Earth when the Earth's north pole is furthest from the sun (northern hemisphere winter solstice). Let (r, θ) be the usual polar coordinates for the centre of the Earth on the orbital plane. Approximate the incoming solar radiation to the Earth as parallel rays traveling from the direction $q = -[\cos \theta, \sin \theta, 0]^T$ with energy flux $S_0 = 1366 \text{ W m}^{-2}$ (the solar constant). Let
660 ϕ and ψ be the latitude and "longitude" of a location on the Earth's surface, where we assume that $\psi = 0$ is aligned with the positive x -axis of the solar coordinate system, not with some fixed location on the Earth's surface. If the Earth's axis of rotation was parallel to the z -axis (no tilt) then the unit outward normal to the Earth's surface in the solar coordinate system would be

$$\begin{bmatrix} \cos \phi \cos \psi \\ \cos \phi \sin \psi \\ \sin \phi \end{bmatrix}, \quad -\frac{\pi}{2} \leq \phi \leq \frac{\pi}{2}, \quad -\pi < \psi \leq \pi.$$

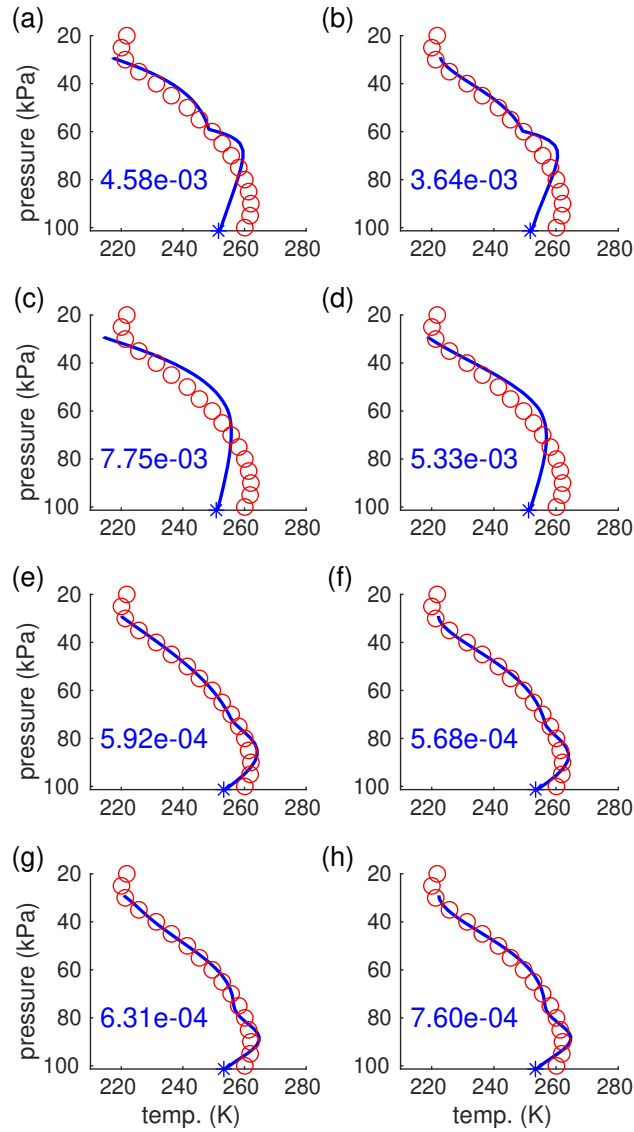


Figure B2. Temperature profiles for best fits for each of the forms of ϕ and ψ . The top row, panels (a) and (b), correspond to Eqn. (A55) for ϕ , the second row, panels (c) and (d), to Eqn. (A56), and similarly, the panels in the third and fourth rows correspond to Eqns. (A57) and (A58), respectively. The left column, panels (a), (c), (e), and (g), correspond to Eqn. (A52) for ψ , while panels in the right column correspond to Eqn. (A53). The data from Cronin and Jansen (2016) are the red circles. The numbers in the bottom left are the residuals for the fits.

665 However, the Earth's axis is actually tilted by an angle $\beta = \frac{23.5}{180}\pi$ in the negative sense in the (x, z) -plane. Applying this tilt gives the unit outward normal as

$$n = \begin{bmatrix} \cos \beta & 0 & \sin \beta \\ 0 & 1 & 0 \\ -\sin \beta & 0 & \cos \beta \end{bmatrix} \begin{bmatrix} \cos \phi \cos \psi \\ \cos \phi \sin \psi \\ \sin \phi \end{bmatrix} = \begin{bmatrix} \cos \beta \cos \phi \cos \psi + \sin \beta \sin \phi \\ \cos \phi \sin \psi \\ -\sin \beta \cos \phi \cos \psi + \cos \beta \sin \phi \end{bmatrix}.$$

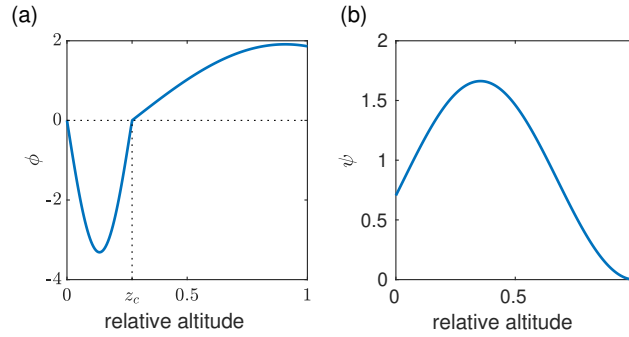


Figure B3. Calibrated functions: (a) ϕ , and (b) ψ .

The insolation striking the Earth's surface at (ϕ, ψ) is then $S_0 \max(q \cdot n, 0)$, where the maximum is due to the fact that the dot product is negative for points on the dark side of the Earth, away from the sun, and hence the insolation there is zero.

Let S be the region of the Earth between latitudes ϕ_1 and ϕ_2 . The area element is $dS = R^2 \cos \phi d\psi d\phi$, where R is the radius of the Earth. Therefore the average annual (θ runs 0 to 2π) insolation on a region, S , of the Earth is

$$Q = \frac{\int_0^{2\pi} \int_S S_0 \max(q \cdot n, 0) dS d\theta}{\int_S dS} = \frac{S_0}{2\pi(\sin(\phi_2) - \sin(\phi_1))} \int_0^{2\pi} \int_{\phi_1}^{\phi_2} \int_{-\pi}^{\pi} \max(q \cdot n, 0) \cos \phi d\psi d\phi d\theta.$$

Numerical integration of the above formula yields $Q = 341 \text{ W m}^{-2}$ for the entire globe, which agrees well with Wild's value of 340 (the difference is likely due to some ambiguity in the precise value of S_0), and yields $Q = 185 \text{ W m}^{-2}$ for the Arctic region north of 70° latitude. If one considers the limit as the region of interest shrinks to size zero around the North Pole, the value of the insolation limits to 173.85 W m^{-2} .

Code availability. Analysis code available from authors on request.

Competing interests. The authors have no competing interests.

Author contributions. KK roles: formal analysis, investigation, methodology, software, validation, visualization, writing — review & editing. WL roles: conceptualization, funding acquisition, writing — review & editing. GL roles: methodology, software (supporting), funding acquisition, writing — review & editing. AW roles: methodology, writing — original draft, software (supporting), funding acquisition, visualization, writing — review & editing.

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