2	Non-Gaussian statistics in global atmospheric dynamics: a
3	study with a 10240-member ensemble Kalman filter using an
4	intermediate AGCM
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#### 19 Abstract.

We previously performed local ensemble transform Kalman filter (LETKF) experiments with up to 20 10240 ensemble members using an intermediate atmospheric general circulation model (AGCM). 21 22 While the previous study focused on the impact of localization on the analysis accuracy, the present study focuses on the probability density functions (PDFs) represented by the 10240-member 23 ensemble. The 10240-member ensemble can resolve the detailed structures of the PDFs and indicates 24 25 that non-Gaussianity is caused in those PDFs by multimodality and outliers. The results show that the spatial patterns of the analysis errors are similar to those of non-Gaussianity. While the outliers appear 26 randomly, large multimodality corresponds well with large analysis error, mainly in the tropical 27regions and storm track regions where highly nonlinear processes appear frequently. Therefore, we 28 further investigate the lifecycle of multimodal PDFs, and show that the multimodal PDFs are mainly 29 30 generated by the on-off switch of convective parameterization in the tropical regions and by the instability associated with advection in the storm track regions. Sensitivity to the ensemble size 31 suggests that approximately 1000 ensemble members be necessary in the intermediate AGCM-32 LETKF system to represent the detailed structures of the non-Gaussian PDF such as skewness and 33 kurtosis; the higher-order non-Gaussian statistics are more vulnerable to the sampling errors due to a 34 smaller ensemble size. 35

# 36 1 Introduction

Data assimilation is a statistical approach to estimate a posterior probability density function (PDF) using 37 information of a prior PDF and observations. Based on the posterior PDF estimate, the optimal initial state is 38 given for numerical weather prediction (NWP). The ensemble Kalman filter (EnKF; Evensen 1994) 39 is an ensemble data assimilation method based on the Kalman filter (Kalman 1960) and approximates 40 the background error covariance matrix by an ensemble of forecasts. The EnKF can explicitly 41 42 represent the PDF of the model state, where the ensemble size is essential because the sampling error contaminates the PDF represented by the ensemble. Although the sampling error is reduced by 43 increasing the ensemble size, the EnKF is usually performed with a limited ensemble size up to 44 O(100) due to the high computational cost of ensemble model runs. Recently, EnKF experiments with 45 a large ensemble have been performed using powerful supercomputers. Miyoshi et al. (2014; hereafter 46 47 MKI14) implemented a 10240-member EnKF with an intermediate atmospheric general circulation model (AGCM) known as the Simplified Parameterizations, Primitive Equation Dynamics model 48 (SPEEDY; Molteni 2003), and found meaningful long-range error correlations. In addition, they 49 reported that sampling errors in the error correlation were reduced by increasing the ensemble size. 50 Further, Miyoshi et al. (2015) assimilated real atmospheric observations with a realistic model known 51as the Nonhydrostatic Icosahedral Atmospheric Model (NICAM; Tomita and Satoh 2004; Satoh et al. 52 2008; 2014) using an EnKF with 10240 members. Kondo and Miyoshi (2016; hereafter KM16) 53 investigated the impact of covariance localization on the accuracy of analysis using a modified 54 version of the MKI14 system. 55

56	MKI14 also focused on the PDF and reported strong non-Gaussianity, such as a bimodal PDF.
57	Previous studies investigated the impact of non-Gaussianity on the EnKF. Anderson (2010) reported
58	that an $N$ -member ensemble could contain an outlier and a cluster of $N$ -1 ensemble members under
59	nonlinear scenarios using the ensemble adjustment Kalman filter (EAKF; Anderson 2001). Anderson
60	(2010) called this phenomenon ensemble clustering (EC), which leads to degradation of analysis
61	accuracy. Amezcua et al. (2012) investigated EC with the ensemble transform Kalman filter (ETKF;
62	Bishop et al. 2001) and local ensemble transform Kalman filter (LETKF; Hunt et al. 2007), and found
63	that random rotations of the ensemble perturbations could avoid EC. Posselt and Bishop (2012)
64	explored the non-Gaussian PDF of microphysical parameters using an idealized one-dimensional
65	(1D) model of deep convection and showed that the non-Gaussianity of the parameter was generated
66	by nonlinearity between the parameters and model output.
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76 behavior and lifecycle of the non-Gaussian PDF. To the best of the authors' knowledge, this is the first study investigating the non-Gaussian PDF using a 10240-member ensemble of an intermediate 77AGCM. This study also discusses how many ensemble members are necessary to represent non-78 Gaussian PDF without contamination by the sampling error, since in general higher-order non-79 Gaussian statistics are more vulnerable to the sampling error due to a limited ensemble size. This 80 paper is organized as follows. Section 2 describes measures for the non-Gaussian PDF. Section 3 81 describes experimental settings, and Section 4 presents the results. Finally, summary and discussions 82 are provided in Section 5. 83

84

#### 85 2 Non-Gaussian measures

Sample skewness  $\beta_1^{1/2}$  and sample excess kurtosis  $\beta_2$  are well-known parametric properties of a non-Gaussian PDF, and are defined as follows:

$$\beta_1^{1/2} = \frac{N}{(N-1)(N-2)} \frac{\sum_{i=1}^N (x_i - \bar{x})^3}{\sigma^3} \tag{1}$$

$$\beta_2 = \frac{N(N+1)}{(N-1)(N-2)(N-3)} \frac{\sum_{i=1}^{N} (x_i - \bar{x})^4}{\sigma^4} - \frac{3(N-1)^2}{(N-2)(N-3)}$$
(2)

where  $x_i$  and  $\bar{x}$  denote the *i*th ensemble member and *N*-member ensemble mean, respectively;  $\sigma$ denotes the sample standard deviation, i.e.,  $\sigma = \sqrt{\frac{1}{N-1}\sum_{i=1}^{N}(x_i - \bar{x})^2}$ , and skewness  $\beta_1^{1/2}$  represents the asymmetry of the PDF. Positive (negative) skewness  $\beta_1^{1/2}$  corresponds to the PDF with the longer tail on the right (left) side. Positive (negative) kurtosis  $\beta_2$  corresponds to the PDF with a more pointed (rounded) peak and longer (shorter) tails on both sides. When the PDF is Gaussian, both skewness  $\beta_1^{1/2}$  and kurtosis  $\beta_2$  go to zero in the limit of infinite sample size. In addition, we also use Kullback–Leibler divergence (KL divergence, Kullback and Leibler 1951) from the Gaussian PDF. KL divergence is a direct measure of the difference between two PDFs. Let p(x) and q(x) be two PDFs. The KL divergence  $D_{KL}$  between the two PDFs is defined as

$$D_{KL} = \int p(x) \log \frac{p(x)}{q(x)} dx$$
(3)

97 Here, we obtain p(x) from the histogram based on the ensemble, and q(x) from the theoretical Gaussian function with the ensemble mean  $\bar{x}$  and standard deviation  $\sigma$ , respectively.  $D_{KL}$  measures 98 the difference between the ensemble-based histogram and the fitted Gaussian function. Figure 1 99 shows examples of ensemble-based histograms and corresponding skewness  $\beta_1^{1/2}$ , kurtosis  $\beta_2$ , and 100 KL divergence  $D_{KL}$  with 10240 samples. Here, the Scott's choice method (Scott 1979) is applied to 101 decide the bin width for histograms. The histogram with KL divergence  $D_{KL} = 0.01$  looks 102 approximately Gaussian while the other three histograms with larger  $D_{KL}$  values show significant 103 discrepancies from the Gaussian function. The skewness and kurtosis measure the degrees of 104 symmetry and tailedness, respectively, while the KL divergence  $D_{KL}$  is more suitable for measuring 105 the degrees of difference between a given PDF and the fitted Gaussian function. Based on the 106 subjective observation of Fig. 1, hereafter, the PDF is considered to be non-Gaussian when  $D_{KL}$  > 1070.01. 108

A non-Gaussian PDF can also be caused by outliers. Although detailed results are shown in
 Section 4, one or several ensemble members are detached from the main cluster; this also results in

111 the large KL divergence  $D_{KL}$ , as well as large skewness and kurtosis, shown in Fig. 2b. We tested two outlier detection methods: the standard deviation-based method (SD method) and the local outlier 112factor method (LOF method; Breunig et al. 2000). Here, univariate PDFs are considered, so that SD 113 and LOF methods are computed for each variable at each grid point separately. 114In the SD method, the ensemble members beyond a prescribed threshold in the unit of SD are 115defined as outliers. If we make 10240 independent random draws from a Gaussian PDF, statistically 116 27.6, 0.65, and 0.0059 samples (0.270, 0.00633, and 0.0000573 %) are expected beyond the  $\pm 3\sigma$ ,  $\pm 4\sigma$ , 117 and  $\pm 5\sigma$  thresholds, respectively. Namely, with the threshold of  $\pm 3\sigma$ , we would expect to detect 27.6 118 outliers at every grid point. With the  $\pm 4\sigma$  threshold, we would expect to detect 1.3 outliers in two grid 119 120 points (20480 random draws). With the  $\pm 5\sigma$  threshold, we would expect to detect 1.18 outliers in 200 121 grid points (2048000 random draws). Since outliers appear frequently with  $\pm 3\sigma$  and  $\pm 4\sigma$  thresholds, we choose the  $\pm 5\sigma$  threshold for the SD method in this study. 122

Unlike the SD method, the LOF method is based on the local density, not on the distance from the sample mean. For a given two-dimensional dataset D, let d(p, o) denote the distance between two objects  $p \in D$  and  $o \in D$ . For any positive integer k, define k-distance(p) to be the distance between the object p and the kth nearest neighbor. The k-distance neighborhood of p, or simply  $N_k$ (p), is defined as the k nearest objects:

$$N_k(p) = \{q \in D \mid q \neq p, d(p,q) \le k\text{-}distance(p)\}$$

$$\tag{4}$$

The cardinality of  $N_k(p)$ , or  $|N_k(p)|$ , is greater than or equal to the number of objects (except for the object *p* itself) within *k*-distance(*p*). We define the *reachability distance* of *p* with respect to the object

$$reach-dist_k(p, o) = \max\{k-distance(o), d(p, o)\}$$
(5)

That is, if the object *p* is sufficiently distant from the object *o*, *reach-dist*<sub>k</sub>(*p*, *o*) is *d*(*p*, *o*). If they are sufficiently close to each other, *reach-dist*<sub>k</sub>(*p*, *o*) is replaced by *k-distance*(*o*) instead of *d*(*p*, *o*). Figure 3 shows a schematic diagram of *reach-dist*<sub>k</sub>(*p*, *o*) with k = 3.  $N_k(p)$  includes  $o_1$ ,  $o_2$ ,  $o_3$ , and  $o_4$ , and  $|N_k(p)|$  is 4. In Fig. 3 (a), *reach-dist*<sub>k</sub>(*p*, *o*<sub>1</sub>) is *k-distance*(*o*<sub>1</sub>) = *d*(*o*<sub>1</sub>, *o*<sub>4</sub>) because *k-distance*(*o*<sub>1</sub>) is greater than *d*(*p*, *o*<sub>1</sub>). In contrast, in Fig. 3 (b), *reach-dist*<sub>k</sub>(*p*, *o*<sub>1</sub>) is *d*(*p*, *o*<sub>1</sub>). We further define the *local reachability density* of *p*, or simply *lrd*<sub>k</sub>(*p*), as the inverse of the average of *reachability distance* of *p*:

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$$lrd_{k}(p) = \frac{|N_{k}(p)|}{\sum_{o \in N_{k}(p)} reach-dist_{k}(p, o)}$$
(6)

Finally, the *local outlier factor* of p, denoted as  $LOF_k(p)$ , is defined as:

140 
$$LOF_k(p) = \frac{\sum_{o \in N_k(p)} \frac{lrd_k(o)}{lrd_k(p)}}{|N_k(p)|}.$$
 (7)

Given a lower *local reachability density* of *p* and a higher *local reachability density* of *p*'s *k*-nearest neighbors,  $LOF_k(p)$  becomes higher.  $LOF_k(p)$  or simply LOF is approximately 1 for an object deep within a cluster, and LOF becomes larger around the edge of the cluster due to sparse objects on the far side from the cluster. To summarize, the LOF method focuses on the local densities of objects, and outliers are detected by comparing the local densities. For instance, when k = 3 in Fig. 3a, the local densities of the objects *p* and  $o_{1,2,3,4,5}$  have all similar values because the *k*-*distance*(*p*) is similar to the *k*-*distances*( $o_{1,2,3,4,5}$ ). Therefore, they are not identified as outliers. In contrast, in Fig. 3b the

148	object p has a smaller local density than the other objects $o_{1, 2, 3, 4, 5}$ because k-distance(p) > k-
149	<i>distances</i> ( $o_{1, 2, 3, 4, 5}$ ). Therefore, the object <i>p</i> has a larger <i>LOF</i> and is identified as an outlier. An object
150	with LOF much larger than 1 may be categorized as an outlier, but it is not clear how to determine
151	the threshold for outliers because the threshold also depends on the dataset. The threshold of LOF is
152	chosen to be 8.0 in this study, and Section 4 shows the results with different values of the threshold
153	and discusses why we choose this value. k is a control parameter for the LOF method and depends on
154	the dataset, as shown by Breunig et al. (2000), who suggested that choosing $k$ from 10 to 20 work
155	well for most of the datasets. If we choose $k$ too small, some objects deeply inside a cluster have a
156	large <i>LOF</i> , and the LOF method does not work. In fact, using the dataset of KM16, $k = 10$ showed
157	this problem, while $k = 20$ did not. Therefore, we chose $k = 20$ in this study. Similar to the SD method,
158	the LOF method is applied to a one-dimensional dataset consisted of 10240 ensemble members.
159	The statistics of the KL divergence, SD and LOF methods with 10240 samples are evaluated
160	numerically with 1 million trials of 10240 random draws from the standard normal distribution by
161	the Box-Muller's method (Box and Muller 1958). The results show that the expected value of KL
162	divergence $D_{KL}$ is 0.0025, and its standard deviation is 0.00048. As for outlier detections, 5767 and
163	16088 trials have at least one outlier for SD and LOF methods, respectively. Namely, the probabilities
164	to detect at least one outlier at a grid point are 0.58 % for the SD method and 1.6 % for the LOF
165	method. Here, the threshold for the SD method is $\pm 5\sigma$ . For the LOF method, we choose $k = 20$ and,
166	as discussed below in Section 4, the threshold value $LOF = 8.0$ .

# 168 **3 Experimental settings**

169 We use the 10240-member global atmospheric analysis data from an idealized LETKF experiment of KM16. That is, the experiment was performed with the SPEEDY-LETKF system (Miyoshi 2005) 170consisting of the SPEEDY model (Molteni 2003) and the LETKF (Hunt et al. 2007; Miyoshi and 171 Yamane 2007). The SPEEDY model is an intermediate AGCM based on the primitive equations at 172T30/L7 resolution, which corresponds horizontally to  $96 \times 48$  grid points and vertically to seven 173174levels, and has simplified forms of physical parametrization schemes including large-scale condensation, cumulus convection (Tiedtke 1993), clouds, short- and long-wave radiation, surface 175 fluxes, and vertical diffusion. Due to the very low computational cost, the SPEEDY model has been 176 used in many studies on data assimilation (e.g., Miyoshi 2005; Greybush et al. 2011; Miyoshi 2011; 177 Amezcua et al. 2012; Miyoshi and Kondo 2013; Kondo et al. 2013; MKI14; KM16). 178179 The LETKF applies the ETKF (Bishop et al. 2001) algorithm to the local ensemble Kalman filter (LEKF; Ott et al. 2004). The LETKF can assimilate observations at every grid point independently, 180 which is particularly advantageous in high-performance computation. In fact, Miyoshi and Yamane 181 (2007) showed that the parallelization ratio reached 99.99% on the Japanese Earth Simulator 182 supercomputer, and KM16 performed 10240-member SPEEDY-LETKF experiments within 5 183 minutes for one execution of LETKF, not including the forecast part on 4608 nodes of the Japanese 184

185 K supercomputer. The LETKF is computed as follows. Let **X** ( $\delta$ **X**) denote an  $n \times m$  matrix, whose

186 columns are composed of *m* ensemble members (deviations from the mean of the ensemble) with the

187 system dimension n. The superscripts a and f denote the analysis and forecast, respectively. The

188 analysis ensemble  $\mathbf{X}^a$  is written as:

$$\mathbf{X}^{a} = \bar{\mathbf{x}}^{f} \mathbf{1} + \delta \mathbf{X}^{f} \left[ \widetilde{\mathbf{P}}^{a} (\mathbf{H} \delta \mathbf{X}^{f})^{\mathrm{T}} \mathbf{R}^{-1} (\mathbf{y}^{o} - \mathbf{H} \bar{\mathbf{x}}^{f}) \mathbf{1} + \sqrt{m - 1} (\widetilde{\mathbf{P}}^{a})^{1/2} \right]$$
(8)

[cf. Eqs. (6) and (7) of Miyoshi and Yamane 2007]. Here,  $\bar{\mathbf{x}}^f$ ,  $\mathbf{y}^o$ , **H**, and **R** denote the background ensemble mean, observations, linear observation operator, and observation error covariance matrix, respectively. **1** is an *m*-dimensional row vector with all elements being 1. The  $m \times m$  analysis error covariance matrix  $\tilde{\mathbf{P}}^a$  in the ensemble space is given as

$$\widetilde{\mathbf{P}}^{a} = [(m-1)\mathbf{I}/\rho + (\mathbf{H}\delta\mathbf{X}^{f})^{\mathrm{T}}\mathbf{R}^{-1}(\mathbf{H}\delta\mathbf{X}^{f})]^{-1} = \mathbf{U}\mathbf{D}^{-1}\mathbf{U}^{\mathrm{T}}$$
(9)

193 [cf. Eqs. (3) and (9) of Miyoshi and Yamane 2007]. Here,  $\rho$  denotes the covariance inflation factor. 194 As  $\tilde{\mathbf{P}}^a$  is real symmetric, U is composed of the orthonormal eigenvectors, such that  $\mathbf{U}\mathbf{U}^{\mathrm{T}} = \mathbf{I}$ . The 195 diagonal matrix **D** is composed of the non-negative eigenvalues.

196 KM16 performed a perfect-model twin experiment for 60 days from 0000 UTC 1 January in the second year of the nature run, which was initiated at 0000 UTC 1 January from the standard 197 atmosphere at rest (zero wind). The first year of the nature run was discarded as spin-up. To resolve 198detailed PDF structures, the ensemble size was fixed to 10240. No localization was applied, yielding 199 the best analysis accuracy as shown by KM16 who compared five 10240-member experiments with 200 201 different choices of localization: step functions with 2000-km, 4000-km and 7303-km localization radii, a Gaussian function with a 7303-km localization radius, and no localization. The observations 202 for horizontal wind components (U, V), temperature (T), specific humidity (Q), and surface pressure 203 (Ps) were simulated by adding observational errors to the nature run every 6 h at radiosonde-like 204locations (cf. Fig. 8, crosses) for all seven vertical levels, but the observations of specific humidity 205

were simulated from the bottom to the fourth model level (about 500 hPa). The observational errors were generated from independent Gaussian random numbers, and the observational error standard deviations were fixed at 1.0 m s<sup>-1</sup>, 1.0 K, 0.1 g kg<sup>-1</sup>, and 1.0 hPa for U/V, T, Q, and Ps, respectively. The non-Gaussian measures, skewness  $\beta_1^{1/2}$ , kurtosis  $\beta_2$ , and KL divergence  $D_{KL}$ , are calculated at each grid point for each variable. Outliers are diagnosed similarly at each grid point for each variable with the SD method and LOF method.

212

# 213 **4 Results**

Figure 4 shows the spatial distributions of the analysis absolute error, ensemble spread, background 214 skewness  $\beta_1^{1/2}$ , kurtosis  $\beta_2$ , and KL divergence  $D_{KL}$  for temperature at the fourth model level (~500 215hPa) at 0600 UTC 22 February. When the analysis absolute error is large, the background non-216Gaussian measures also tend to be large, especially in the tropics. The peaks for skewness  $\beta_1^{1/2}$ , 217 kurtosis  $\beta_2$ , and KL divergence  $D_{KL}$  tend to coincide. Although grid point A (16.7°S, 90.0°E) has a 218 large KL divergence  $D_{KL}$  with large analysis absolute error, at grid point B (35.3°N, 146.3°E) with a 219 large KL divergence  $D_{KL}$  the analysis absolute error is small (< 0.08 K). This result shows that the 220large analysis error is not always associated with the strong non-Gaussianity at a specific time. The 221 PDFs at grid points A and B are shown in Fig. 2a, b, respectively. The histogram at the grid point A 222 is clearly a multimodal PDF with KL divergence  $D_{KL} > 0.01$ , and the right mode captures the truth 223(yellow star). At grid point B, although the PDF seems to be closer to Gaussian, skewness  $\beta_1^{1/2}$  and 224

225	kurtosis $\beta_2$ are much larger than those at grid point A. In fact, the PDF does not fit to the Gaussian
226	function calculated by the ensemble mean and standard deviation. Zooming in on the left side of Fig.
227	2b shows a small cluster composed of 76 members detached from the main cluster; 74 members of
228	the small cluster exceed $-5\sigma$ and are categorized as outliers in the SD method. This small cluster
229	causes the standard deviation to become large and results in the Gaussian function having a longer
230	tail than the histogram. The small cluster should not be considered as consisting of outliers because
231	it may have some physical significance. Scatter diagrams of LOF versus distance from ensemble
232	mean for all ensemble members at grid points A and B are shown in Fig. 5a, b, respectively. At grid
233	point A, $LOF$ is not so large even at the edge of the cluster (< 4), and the multimodal PDF does not
234	influence <i>LOF</i> . In addition, all members are within $\pm 3\sigma$ . Therefore, there are no clear outliers at grid
235	point A. At grid point B, although most of the small cluster exceeds $-5\sigma$ , the maximum <i>LOF</i> in the
236	small cluster is still smaller than 3. This indicates that all members of the small cluster should not be
237	outliers in the LOF method. As an outlier case, we pick up the grid point C (35.3°N, 112.5°W) in Fig.
238	4. The PDF at the grid point C fits the Gaussian function well, and the non-Gaussian measures are
239	quite small (Fig. 2c). A member on the left edge of the scatter diagram in Fig. 5c has the largest LOF
240	> 8.0, but the member is within $\pm 3\sigma$ . As mentioned in Section 2, the threshold of <i>LOF</i> for outliers
241	depends on the dataset. Figure 6 shows the number of outliers for thresholds of 5.0, 8.0, and 11.0 at
242	0600 UTC 22 February. There are too many outliers with threshold $= 5.0$ , but in contrast, the number
243	of outliers decreases markedly with threshold = $8.0$ or $11.0$ . Based on these results, and as already
244	mentioned in Section 2, we adopt $LOF = 8.0$ as a threshold for outliers.

245	Figure 7 shows the spatial distributions of the time-mean analysis RMSE, ensemble spread, the
246	background absolute skewness $\beta_1^{1/2}$ , absolute kurtosis $\beta_2$ , and KL divergence $D_{KL}$ . As mentioned in
247	KM16, the time-mean ensemble spread corresponds well to the RMSE, which is larger in the tropics.
248	The pattern correlation between the RMSE and ensemble spread is 0.97. Moreover, the distributions
249	of non-Gaussian measures are similar to each other and also correspond well to the RMSE and
250	ensemble spread. The RMSE and non-Gaussian measures differ in that the non-Gaussianity is large
251	in storm tracks, such as the North Pacific Ocean and the North Atlantic Ocean. This may be because
252	the LETKF inhibits growing errors well in storm tracks regardless of the strong non-Gaussianity. To
253	investigate the non-Gaussianity in more detail, Figs. 8 and 9 show the frequencies of non-Gaussian
254	PDF with high KL divergence $D_{KL} > 0.01$ and identifying at least one outlier with high $LOF > 8.0$
255	on a 10240-member ensemble, respectively. The frequency of non-Gaussian PDF is defined as the
256	ratio of non-Gaussianity appearance at every grid point during the 36-day period from 0000 UTC 25
257	January to 1800 UTC 1 March. The spatial distribution of frequency of non-Gaussianity for
258	temperature is similar to that of the time mean RMSE and $D_{KL}$ (Figs. 7 a, e, and 8 b), and the pattern
259	correlation between the spatial distribution of mean RMSE and $D_{KL}$ is 0.68. We find high frequency
260	of non-Gaussian PDF in the tropics and storm track regions for temperature, specific humidity, and
261	surface pressure, although non-Gaussian PDF seldom appears in the densely observed regions. In the
262	tropics, the frequency reaches up to 90%, and in South America the frequency reaches the highest
263	value over 95%, i.e., the non-Gaussian PDF appears for 34 days out of 36 days. In contrast, the non-
264	Gaussian PDF for zonal wind hardly appears (Fig. 8 a), and the intensity of the non-Gaussianity, as

evaluated by other measures, is also weak (not shown). On the other hand, the outliers appear almost
randomly and do not clearly depend on the region for any of the variables (Fig. 9), and most outliers
disappear within only one or a few analysis steps. Moreover, there are no correlations between the
frequency of outliers and analysis RMSE.

To investigate how the non-Gaussian PDF is generated, we plot the forecast and analysis update 269processes at 1.9°N, 168.7°E for 256 members chosen randomly from 10240 members from the 270analysis at 0000 UTC 9 February (157th analysis cycle) to the forecast at 0000 UTC 10 February 271 (161<sup>st</sup> analysis cycle, Fig. 10a). That is, Fig. 10a shows the lifecycle of the non-Gaussian PDF. As the 272 vertical axis, we introduce the convective instability  $d\theta_e$ , which is defined as a difference between 273274 equivalent potential temperature  $\theta_e$  at the fourth model level (~500 hPa) and  $\theta_e$  at the second model 275level (~850 hPa). Negative (Positive)  $d\theta_e$  indicates a convectively unstable (stable) atmosphere. The non-Gaussian PDF appears in the background at the 159<sup>th</sup> cycle (1200 UTC 9 February), and the 276model forecast increases the KL divergence  $D_{KL}$  for  $d\theta_e$  up to 0.154 with a bimodal PDF of clusters 277 A and B. We find many lines crossing in the forecast step from the analyses at the 158<sup>th</sup> cycle to the 278background at the 159<sup>th</sup> cycle. Namely, many of the upper side cluster A at the 159<sup>th</sup> cycle come from 279 the lower side analyses in the previous 158<sup>th</sup> cycle, generally reducing the instability (increasing 280 values of  $d\theta_e$ ) in the forecast step, and vice versa for the lower side cluster B. In the background 281 temperature at the fourth model level, the KL divergence  $D_{KL}$  also increases from 0.003 to 0.299 for 282 6 h (Figs. 10b, c). Finally, the non-Gaussian PDF almost disappears at the 161<sup>st</sup> cycle (0000 UTC 10 283February). Figure 11 shows a scatter diagram of 0600 UTC versus 1200 UTC 9 February for 284

285 background temperature in the fourth model level for each member at 1.9°N, 168.7° E, and also shows histograms corresponding to the scatter diagrams. The PDF at 0600 UTC is almost Gaussian. 286 However, at 1200 UTC, the bimodal structure with KL divergence  $D_{KL} = 0.299$  appears. The dot 287 colors show  $d\theta'_e$  evaluated from 0600 UTC to 1200 UTC 9 February, namely,  $d\theta'_e =$ 288  $(d\theta_{e\ 1200\ UTC} - d\theta_{e\ 0600\ UTC}) - (d\bar{\theta}_{e\ 1200\ UTC} - d\bar{\theta}_{e\ 0600\ UTC})$ , where  $\bar{\theta}_{e}$  indicates the equivalent 289 potential temperature calculated from the ensemble mean. That is, a red (blue) dot shows more 290 stability (instability) than the ensemble mean. The red and blue dots are clearly divided into the right 291 and left side modes, respectively. Most members with mitigated (enhanced) instability move to the 292 right (left) side mode. The members with larger (smaller) temperature values at 1200 UTC correspond 293294 to larger (smaller) values of stability as shown by the warmer (colder) color. In addition, both right 295 and left modes correspond to the opposite side modes in the specific humidity, respectively (not shown). That is, the members with higher (lower) temperature have lower (higher) humidity than the 296ensemble mean. The instability is driven by precipitation. Figure 12 is similar to Fig. 11, but for 297 precipitation. The 10240 members are clearly divided into three clusters at 1200 UTC by the 298instability. The three clusters indicate the number of times cumulus parameterization is triggered. 299 300 Most members in the right (left) cluster are red (blue) and show mitigation (enhancement) of the instability. Figure 13 is also similar to Fig. 11, but for zonal wind at the fourth model level. In 301 agreement with what has been seen on Fig. 8a, the non-Gaussianity of zonal wind is weak, and the 302 bimodal structure appearing in temperature and humidity seldom affects the PDF of zonal wind. We 303 found no relationship between the atmospheric instability and zonal wind. Therefore, the genesis of 304

305 non-Gaussian PDF in the tropics is deeply related to precipitation process, which is driven by convective instability through cumulus parameterization in the SPEEDY model. As a result, the 306 precipitation process mitigates the instability, with rising temperature and decreasing humidity. 307 Similar results are generally obtained at other grid points with non-Gaussian PDF. 308 In the extratropics, non-Gaussian PDF is generated differently. To investigate the genesis of non-309 Gaussian PDF in the extratropics, we focus on a case around an extratropical cyclone over the Atlantic 310 Ocean. A non-Gaussian PDF appears at 0600 UTC 15 February at 42.7°N, 48.8°W, and the KL 311 divergence  $D_{KL}$  of background temperature increases from 0.003 to 0.460 (Fig. 14, crosses). Figure 312 15 is similar to Fig. 11, but for background specific humidity at the second model level (~850 hPa) 313 314 versus precipitation at 42.7°N, 48.8°W at 0006 UTC 15 February. Trimodal PDFs appear in both 315 specific humidity and precipitation. The three modes of specific humidity are clearly separated by the color, i.e., instability  $d\theta'_e$ . Namely, modes with larger humidity has colder colors (smaller  $d\theta'_e$ ) 316corresponding to more instability). However, the three modes of precipitation show no clear 317 dependence on  $d\theta'_{e}$ . Therefore, the trimodal PDF of specific humidity would not be driven by the 318 cumulus parameterization. Next, the relationship between background specific humidity and 319 320 meridional wind at the second model level (~850 hPa) is shown in Fig. 16. The members in the left mode have lower specific humidity with relatively stronger northerly wind. If we look at the fourth 321 model level (~500 hPa) for these members with lower humidity, they have relatively weaker northerly 322 wind and warm temperature (not shown). Namely, instabilities are mitigated by the northerly 323 advection of dry air at the lower troposphere and by warm temperature at the mid troposphere. In this 324

case study, the non-Gaussianity genesis in the extratropics is associated with the advections. This is only an example, and the non-Gaussianity genesis in the extratropics is generally more complicated and would be affected by not only vertical stratification but also larger-scale atmospheric phenomena such as extratropical cyclones and advections. Here, we do not go into details for different cases of non-Gaussianity genesis, but instead, this is further discussed in Section 5.

The non-Gaussian measures are sensitive to the ensemble size due to sampling errors. Figure 17 330 shows the spatial distributions of the skewness  $\beta_1^{1/2}$ , kurtosis  $\beta_2$ , and KL divergence  $D_{KL}$  for 331 temperature at the fourth model level (~500 hPa) at 0600 UTC 22 February with 80, 320, and 1280 332 subsamples from 10240 members, respectively. Skewness  $\beta_1^{1/2}$ , kurtosis  $\beta_2$ , and KL divergence  $D_{KL}$ 333 with 80 members contain high levels of contaminating errors originating from sampling errors, and 334 335 the non-Gaussian measures are difficult to distinguish from the contaminating errors. With increasing the ensemble size up to 1280, the sampling errors become smaller by gradation. With 1280 members, 336 the sampling errors are essentially removed, and the distributions are comparable to those with 10240 337 members (see Fig. 4). Therefore, a sample size of about 1000 members is necessary to represent non-338 Gaussian PDF. The outliers also depend on the sample size. Figure 18 shows LOF with 80, 320, 1280, 339 340 and 5120 subsamples from 10240 members for temperature at the fourth model level at the grid point B (35.3°N, 146.3°E), as in Fig. 5b. With 80 members, there are no outliers as the LOF of each member 341 is much smaller than the outlier threshold of 8.0. When the ensemble size is 320, four members with 342 high LOF > 8.0 are identified as outliers. With the ensemble sizes of 1280 and 5120, 13 and 41 343 members construct a small cluster, respectively, but they are not outliers with the threshold of LOF = 344

8.0. With increasing the ensemble size up to 10240, the *LOF*s of the small cluster and main cluster
show almost the same value (Fig. 5b).

347 We saw a good agreement between the RMSE and ensemble spread (Figs. 7a, b), but it is useful to further evaluate the 10240-member ensemble using ranked probability scores. The rank histogram 348(Hamill and Collucci 1997, Talagrand et al. 1999, Anderson 1996, Hamill 2001) evaluates the 349350 reliability of ensemble statistically. Figure 19 shows almost flat rank histograms at all grid points and the grid points with non-Gaussian PDF. The truth is known in this study and used as a verifying 351 analysis. The flat rank histograms correspond to healthy background ensemble distributions. The 352continuous ranked probability score (CRPS, Hersbach 2000) is another method to evaluate ensemble 353 354 distributions, decomposed into reliability, resolution and uncertainty as

$$CRPS = Reli - Resol + U.$$
(10)

Here, the reliability Reli becomes zero under the perfectly reliable system. The resolution Resol indicates the degree to which the ensemble distinguishes situations with different frequencies of occurrence, and is associated with the accuracy or sharpness. The uncertainty U measures the climatological variability. The reliability, resolution and uncertainty are given on the prescribed area as

$$Reli = \sum_{i=0}^{N} \bar{g}_i (\bar{o}_i - p_i)^2,$$

$$p_i = \frac{i}{N'},$$
(11)

$$U - \text{Resol} = \sum_{i=0}^{N} \bar{g}_i \bar{o}_i (1 - \bar{o}_i), \qquad (12)$$

$$U = \sum_{k,l < k} w_k w_l |y^k - y^l|,$$
(13)

[cf. Eqs 36, 37 and 19 in Hersbach 2000, respectively]. Here,  $\bar{g}_i$  is the area-weighted average width 360 of the bin *i* between consecutive ensemble members  $x_i$  and  $x_{i+1}$ , and  $\bar{o}_i$  is the area-weighted average 361 frequency that the verifying analysis is less than  $(x_{i+1} + x_i)/2$ . N denotes an ensemble size. In this 362 study,  $y^k$  and  $y^l$  indicate the anomalies between the background ensemble mean and monthly 363 climatology computed from a 30-year nature run at the grid points k and l, respectively. The weights 364  $w_{k_1}$ ,  $w_l$  are proportional to the cosine of latitude. Table 1 shows that the reliability is closer to zero and 365 that the resolution is much higher at all grid points than at the grid points with non-Gaussian PDF. 366 Therefore, the non-Gaussian PDF has a negative impact on updating the state variables for the LETKF. 367 The smaller uncertainty at the grid points with non-Gaussian PDF reflects generally smaller variations 368 369 in the tropics where the non-Gaussian PDFs frequently appear. Similar results are obtained for the 370 other variables.

371

## 372 **5** Summary and discussions

Kalman filters provide the minimum variance estimator, which coincides with the maximum
 likelihood estimator if the PDFs are Gaussian. This study investigated the non-Gaussian PDF and its
 behavior using the SPEEDY-LETKF system with 10240 members. Non-Gaussian PDFs appear

376 frequently in the areas where the RMSE and ensemble spread are larger. Moreover, an ensemble size 377 of about 1000 is necessary to identify the possible non-Gaussianity of PDFs, which may be difficult 378 to detect in the presence of sampling error.

The non-Gaussian PDF appears frequently in the tropics and the storm track regions over the 379 Pacific and Atlantic Oceans, particularly for temperature and specific humidity, but not for winds. 380 With the SPEEDY model, the genesis of non-Gaussian PDF in the tropics is mainly associated with 381 the convective instability. These results suggest that the non-Gaussianity be mainly caused by 382 precipitation processes such those associated with cumulus convection, but much less by dynamic 383 processes. Generally, the atmosphere in the tropics tends to become unstable, and the convective 384 instability is mitigated by vertical convection with precipitation. In the SPEEDY model, a simplified 385 386 mass-flux scheme developed by Tiedtke (1993) is applied. Convection occurs when either the specific or relative humidity exceeds a prescribed threshold (Molteni 2003). The members that hit the 387 threshold have precipitation, and this process mitigates their own convective instability resulting in a 388 temperature rise and humidity decrease. In contrast, the members with no or little precipitation 389 enhance or cannot mitigate their own convective instability. Therefore, convective instability is a key 390 391 to non-Gaussianity genesis in the tropics in the SPEEDY model.

In the extratropics, non-Gaussianity is generally weak and seldom appears except in the storm track regions, where the genesis of non-Gaussian PDF is also associated with instabilities, but with different processes from the tropics. This study focused on a case near the extratropical cyclone in the North Atlantic, and the results showed that the instability was associated with the horizontal advections. The members with reduced instabilities had lower humidity at the lower troposphere and higher temperature at the mid troposphere by meridional advections. In contrast, the members with higher humidity at the lower troposphere and lower temperature at the mid troposphere enhanced their instability. Moreover, the precipitation process through the cumulus parameterization did not explain the non-Gaussian PDF. Precipitation associated with extratropical cyclones is usually caused by synoptic-scale baroclinic instabilities and does not mitigate the local instability completely.

As mentioned in Section 4, to generalize the process of non-Gaussianity genesis in the extratropics 402 is not simple. The non-Gaussianity genesis is generally associated with instability from various 403 processes such as the convection, advection and larger-scale atmospheric phenomena, so that it is 404 very difficult to find general mechanisms of the non-Gaussianity genesis in the extratropics even for 405 406 the simple SPEEDY model. Furthermore, if we use more realistic models with complex physics schemes, the process of non-Gaussianity genesis would be much more diverse and complicated. This 407 is partly why we did not go into details to investigate different cases of non-Gaussianity genesis with 408 the SPEEDY model. 409

The non-Gaussianity is less frequent in the wind components not only in the time scale of 1 month but also for the snapshot, although the dynamic process of the atmosphere is a nonlinear system. Moreover, the non-Gaussian PDFs of temperature and specific humidity seldom affect the PDFs of the wind components. We hypothesize that the model complexity may be a reason for this. The SPEEDY model could not resolve some local interactions between wind components and other variables due to its coarse resolution and simplified processes. With more realistic models, physical

processes are much more complex, and the local interactions can also be represented. Indeed, we 416 obtained widely distributed non-Gaussianity with a 10240-member NICAM-LETKF system with 417 112-km horizontal resolution assimilating real observations from the National Centers for 418 Environmental Prediction (NCEP) known as PREPBUFR from 0000 UTC 1 November to 0000 UTC 419 8 November (Miyoshi et al. 2015). Figure 20 shows the spatial distributions of background KL 420 divergence of zonal wind and temperature at the second model level (~850 hPa) for SPEEDY at 0000 421 UTC 1 March and zonal wind and temperature at the eighth model level (~850 hPa) for the NICAM 422 at 0000 UTC 8 November 2011. With NICAM, the non-Gaussianity appears globally not only in the 423 temperature field but also in the zonal wind although we should account for the model errors of 424 NICAM. This result implies that the NICAM has various sources of non-Gaussianity such as smaller 425 426 scale physical and dynamical processes with various interactions among different model variables, and suggests the limitation of this study using the SPEEDY model. In the realistic situation, we would 427 428 presumably have more frequent occurrence of non-Gaussianity.

The outliers appear almost randomly regardless of locations, levels, and variables, and the lifetime is about a few analysis steps. When the outliers appear, the number of outliers is basically one per grid point, but sometimes the number is more than one. Anderson (2010) also reported similar results using a low-order dry atmospheric model. These results seem not to be consistent with Amezcua et al. (2012) who reported that just one outlier appeared with the ensemble square root filters in lowdimensional models and that the outlier did not rejoin the cluster easily. These properties of their outlier and our outliers in the SPEEDY model are somewhat different. In the low-dimensional models,

a certain ensemble member tends to become an outlier at all grid points and all variables. In contrast, 436 the outliers in the SPEEDY model appear at just some grid points but not all grid points and do not 437 appear in all variables simultaneously. In addition, the negative influence of outliers on the analysis 438 accuracy may be sufficiently small in high-dimensional models due to the randomness and short 439 longevity of outliers. In fact, the results showed no clear correspondence between the outlier 440 frequency and analysis accuracy. These are the results from the simple SPEEDY model. It remains to 441 be a subject of future research how the outliers behave with a more realistic model and real 442 observations. 443

As measures of non-Gaussianity, skewness, kurtosis, and KL divergence for the non-Gaussianity, 444and the SD and LOF methods for outliers, are introduced and compared with each other. The KL 445 446 divergence is a more suitable measure because it measures the direct difference between the ensemble-based histogram and the fitted Gaussian function. The LOF method is better than the SD 447 method because it can detect the outliers depending on the density of objects. Although it is easy to 448 detect the outliers using the SD method, misdetection of outliers is possible because this method 449 categorizes a small cluster far from the main cluster into outliers. The small cluster may be generated 450 through physical processes and have physical significance; this should not be treated as outliers. The 451 measures of non-Gaussianity are evaluated in the univariate field in this study. An extension to 452 multivariate fields with multivariate analysis remains as a subject of future research. 453

454 Non-Gaussian measures tend to be more sensitive to the sampling error due to the limited 455 ensemble size (see Figs. 17, 18). When the ensemble size is small, it is difficult to determine whether

456	a split member is a real outlier or a sample from a small cluster. Amezcua et al. (2012) discussed the
457	outliers by skewness using the 20-member SPEEDY-LETKF and reported that the skewness is clearly
458	large in the tropics and the Southern Hemisphere for the temperature and humidity fields. These
459	results were not consistent with those of the present study because the outliers appear randomly.
460	However, this inconsistency may have been due to the small ensemble size. The large skewness of
461	Amezcua et al. (2012) could possibly indicate the non-Gaussianity rather than the outliers with a large
462	ensemble size. Having a sufficient ensemble size, suggested to be about 1000 according to this study,
463	would be essential when discussing about non-Gaussianity and outliers.
464	
465	Data availability
466	All data and source code are archived in RIKEN Center for Computational Science and are available
467	upon request from the corresponding authors under the license of the original providers. The original
468	source code of the SPEEDY-LETKF is available at https://github.com/takemasa-miyoshi/letkf.

469

468

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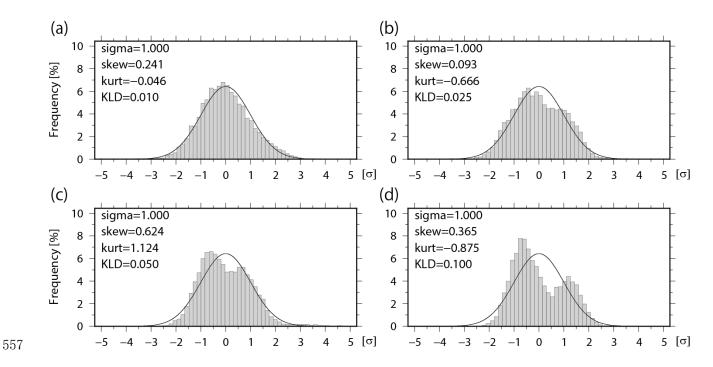
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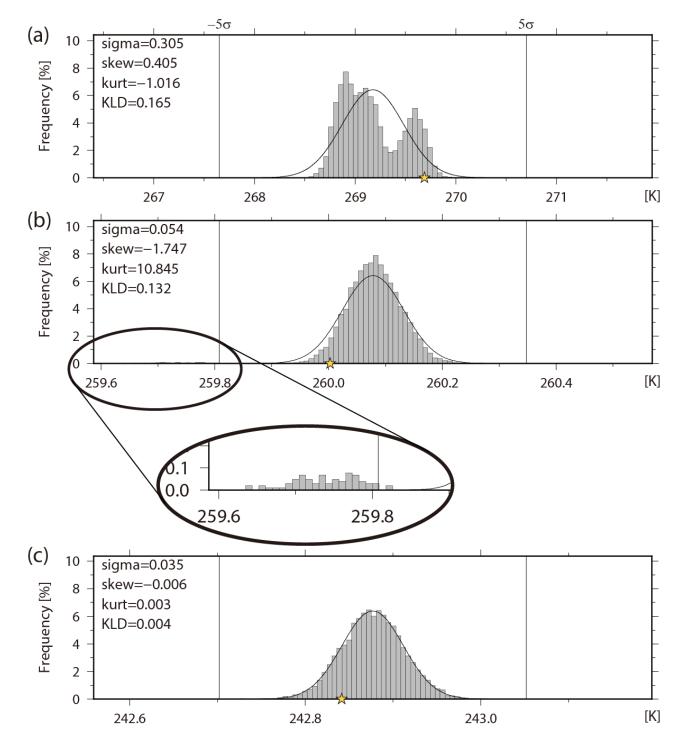
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558 Figure 1: Ensemble-based histograms with 10240 ensemble members when the Kullback–Leibler

559 (KL) divergence  $D_{KL}$  = (a) 0.010, (b) 0.025, (c) 0.050, and (d) 0.100. Solid lines indicate fitted

560 Gaussian functions. Skewness (skew) and kurtosis (kurt) are also shown in the figure.

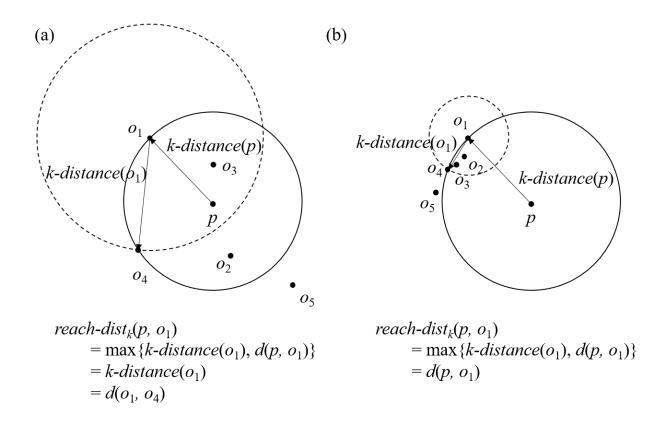


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563 Figure 2: Histograms of background temperature (K) at the fourth model level (~500 hPa) at (a)

<sup>564</sup> grid point A (16.7°S, 90.0°E), (b) grid point B (35.3°N, 146.3°E), and (c) grid point C (35.3°N,

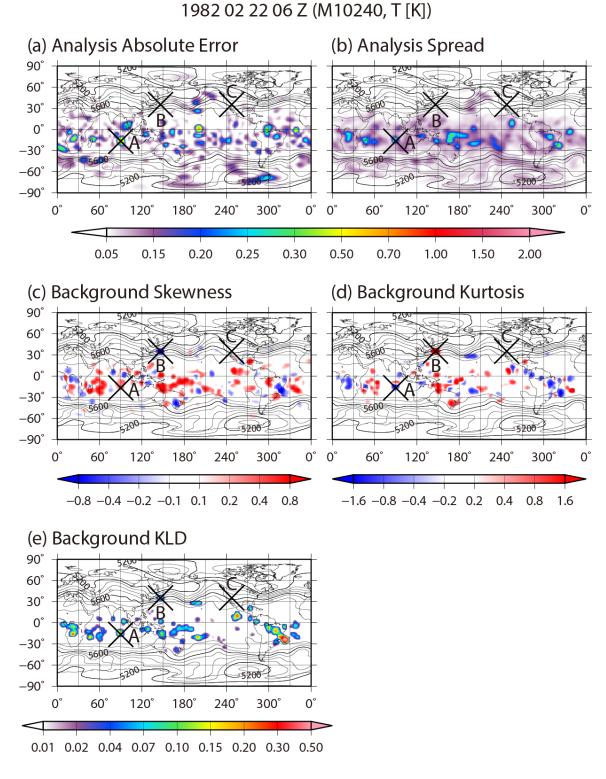
565 112.5°W). The yellow star shows the truth.

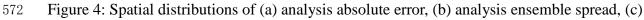




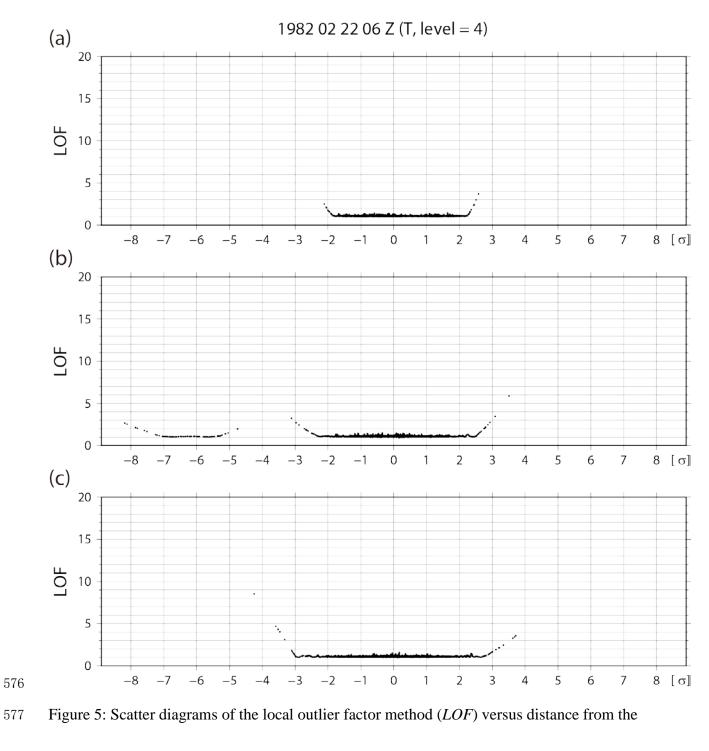
568 Figure 3: Schematic diagrams of *reach-dist*<sub>k</sub>(p, o) with k = 3 for (a) uniformly distributed data and

569 (b) data with an asymmetrical distribution.





- 573 background skewness, (d) background kurtosis, and (e) background KL divergence for temperature
- at the fourth model level (~500 hPa) at 0600 UTC 22 February. Contours indicate geopotential
- height of the ensemble mean at the 500 hPa level.



ensemble mean for all ensemble members for background temperature at the fourth model level
(~500 hPa) at (a) grid point A (16.7°S, 90.0°E), (b) grid point B (35.3°N, 146.3°E), and (c) grid
point C (35.3°N, 112.5°W).

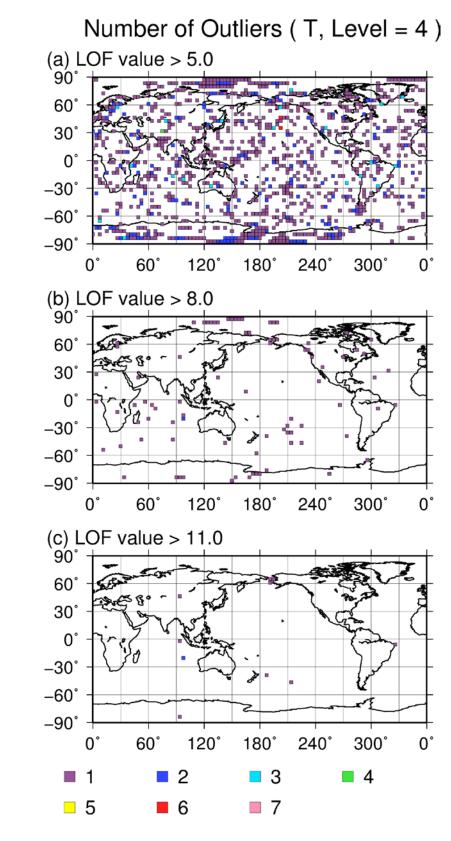


Figure 6: Spatial distributions of the number of outliers for background temperature at the fourth
model level (~500 hPa) at 0600 UTC 22 February for *LOF* thresholds of (a) 5.0, (b) 8.0, and (c)

**11.0**.

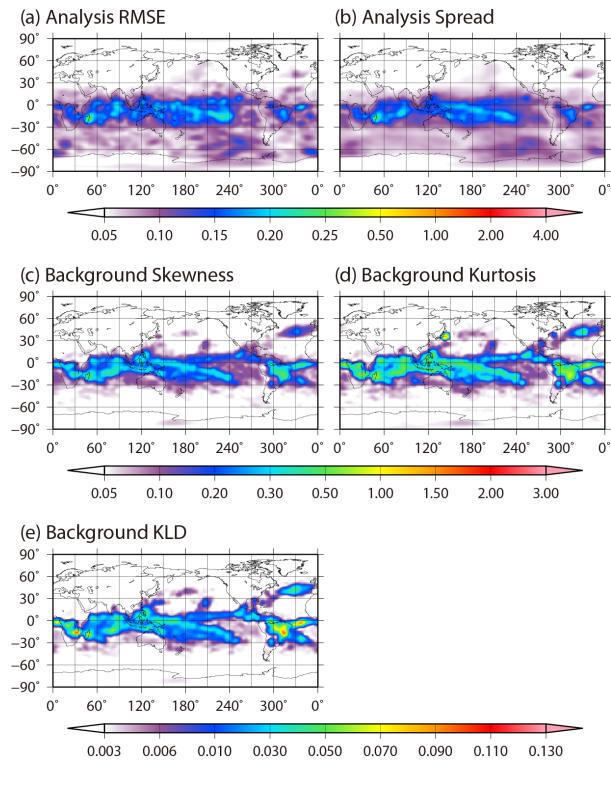
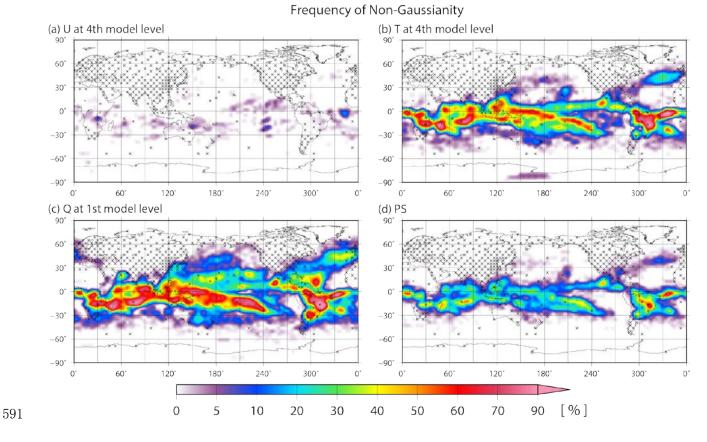


Figure 7: Spatial distributions of the time-mean (a) analysis RMSE, (b) analysis ensemble spread,
(c) background absolute skewness, (d) background absolute kurtosis, and (e) background KL
divergence for temperature at the fourth model level (~500 hPa) from 0000 UTC 25 January to
1800 UTC 1 March.



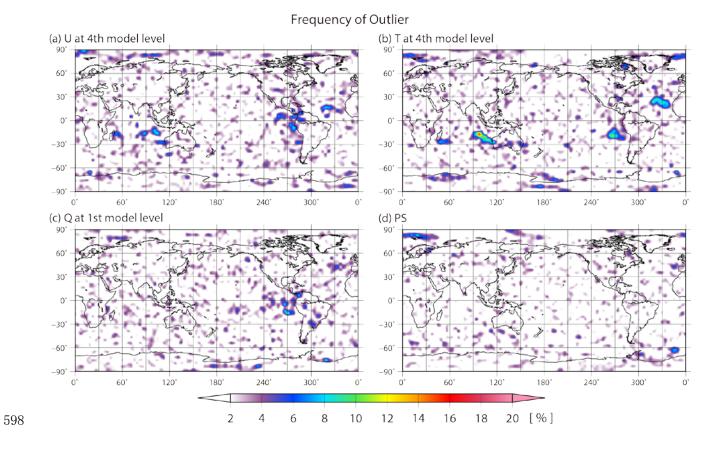
592 Figure 8: Spatial distributions of frequency of non-Gaussian PDF with high KL divergence  $D_{KL}$  >

593 0.01 for (a) zonal wind at the fourth model level, (b) temperature at the fourth model level, (c)

specific humidity at the lowest model level, and (d) surface pressure. The frequency is defined as a

ratio of high KL divergence D<sub>KL</sub> appearance from 0000 UTC 25 January to 1800 UTC 1 March. The

596 crosses indicate the radiosonde-like locations.



599 Figure 9: Similar to Fig. 8, but showing the frequency of identifying at least one outlier with high

LOF > 8.0 on a 10240-member ensemble.

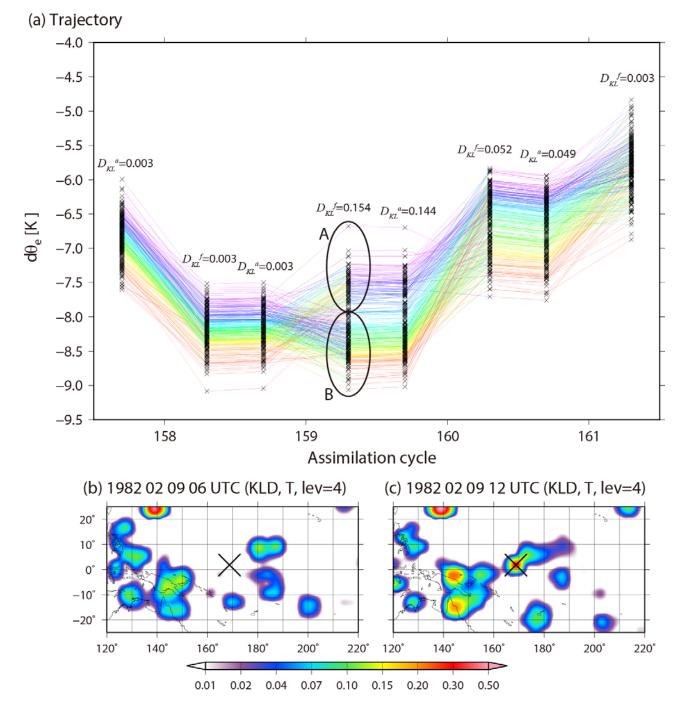
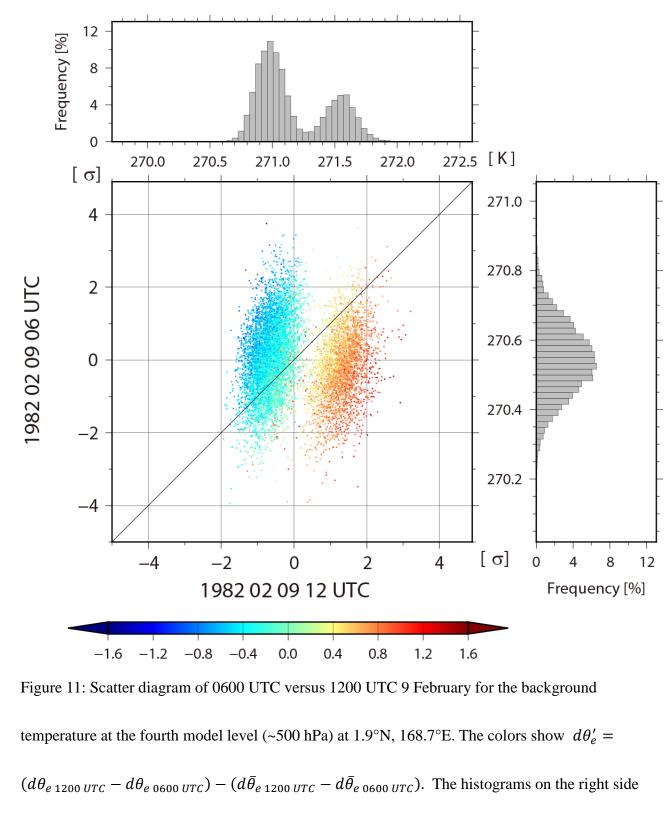


Figure 10: Lifecycle of non-Gaussianity at 1.9°N, 168.7°E. (a) Trajectories of 256 randomly chosen members from 10240 members for  $d\theta_e$  (see text for definition) from analysis at the 157th analysis cycle (0000 UTC 9 February) to forecast the 161st analysis cycle (0000 UTC 10 February). The colors show the order of  $d\theta_e$  for every analysis.  $D_{KL}$  shows KL divergence for  $d\theta_e$ , and the superscripts *a* and *f* indicate analysis and forecast, respectively. (b, c) Spatial distributions of KL

- divergence for background temperature at the fourth model level (~500 hPa) at the 158th analysis
- 609 cycle (0600 UTC 9 February) and the 159th analysis cycle (1200 UTC 9 February), respectively.
- 610 The cross shows the location of the point considered in panel a.



and upper side show the background temperature at the same grid point.

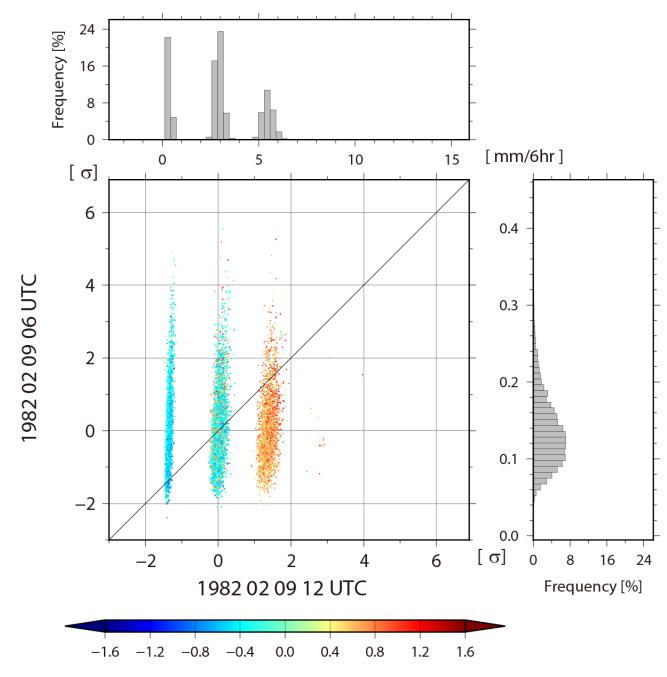
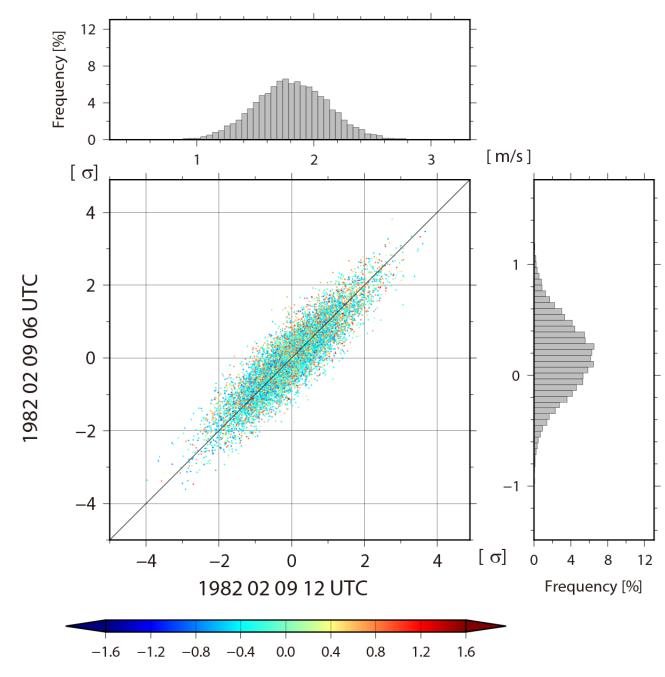


Figure 12: Similar to Fig. 11, but for 0600 UTC versus 1200 UTC 9 February for background

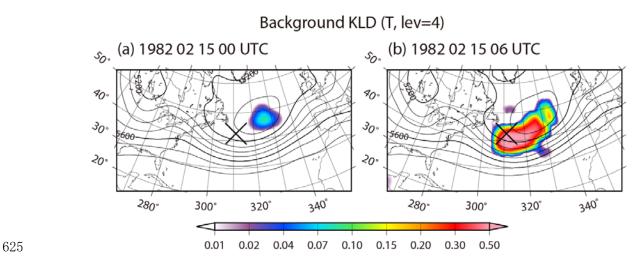
619 precipitation.

620



622 Figure 13: Similar to Fig. 11, but for 0600 UTC versus 1200 UTC 9 February for background zonal

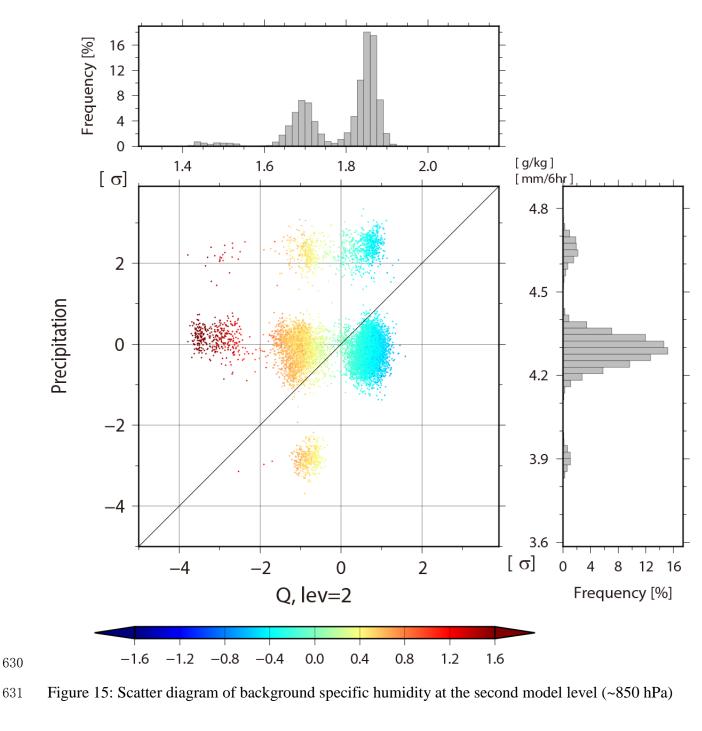
<sup>623</sup> wind at the fourth model level (~500 hPa).



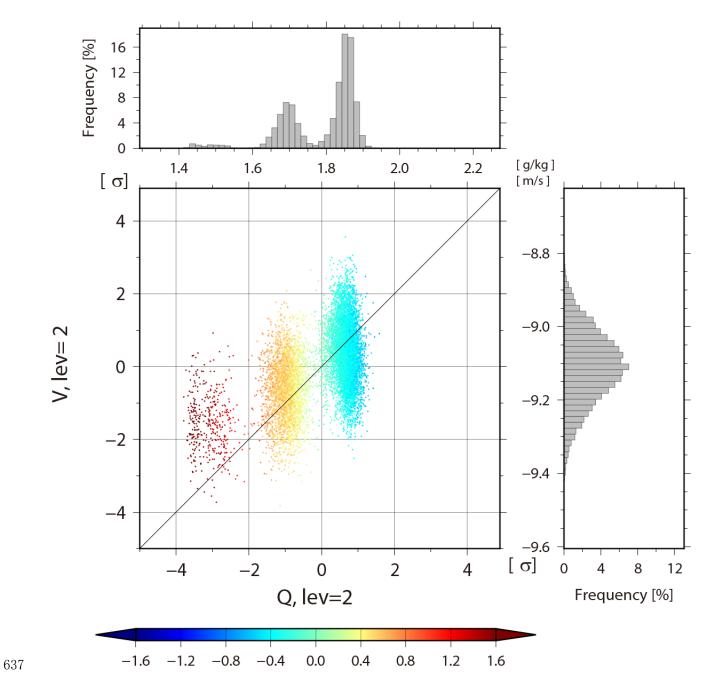
626 Figure 14: Spatial distributions of the KL divergence for background temperature at the fourth

model level (~500 hPa) (a) at 0000 UTC 15 February and (b) at 0600 UTC 15 February. Contours

628 show geopotential height of the ensemble mean at the 500 hPa level.



versus background precipitation at 42.7°N, 48.8°W (311.3°E) at 0600 UTC 15 February. The colors show  $d\theta'_e = (d\theta_{e\ 0600\ UTC} - d\theta_{e\ 0000\ UTC}) - (d\bar{\theta}_{e\ 0600\ UTC} - d\bar{\theta}_{e\ 0000\ UTC})$ . The histograms on the right side and on top show background precipitation and temperature at the same grid point, respectively.



638 Figure 16: Similar to Fig. 14, but for background specific humidity versus meridional wind

<sup>639</sup> background at the second level (~850 hPa).

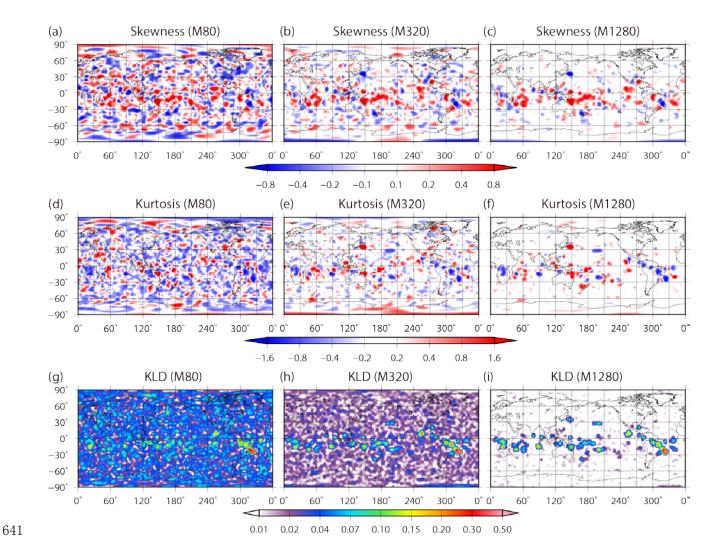


Figure 17: Spatial distributions of (a-c) skewness, (d-f) kurtosis, and (g-i) KL divergence for

temperature at the fourth model level (~500 hPa) at 0600 UTC 22 February. The left, center, and

right columns show 80, 320, and 1280 subsamples from 10240 members, respectively.

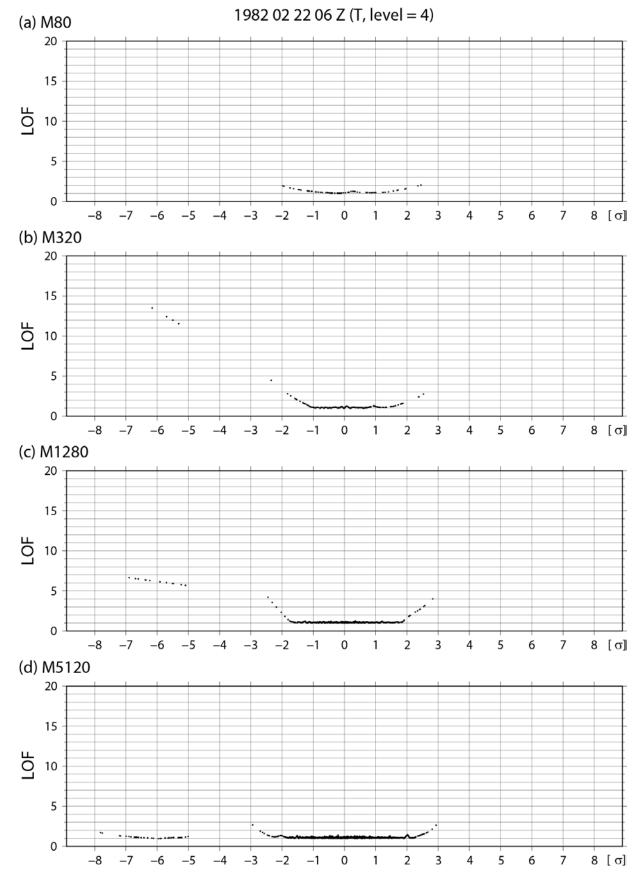
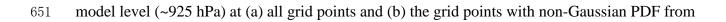


Figure 18: Similar to Fig. 5b, but for the ensemble sizes (a) 80, (b) 320, (c) 1280, and (d) 5120.

## Rank Histogram (Q, Level = 1) (a) All grid points (b) Grid points with non-Gaussian PDF 0.0004 Frequency 0.0002 0.0000 4000 6000 8000 0 2000 10000 0 2000 4000 6000 8000 10000 Rank Rank 649

650 Figure 19: Rank histograms verified against truth for background specific humidity at the lowest



652 0000 UTC 25 January to 1800 UTC 1 March.

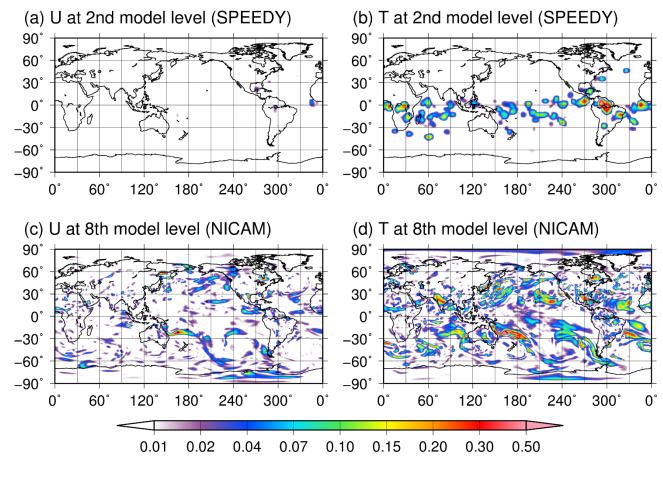


Figure 20: Spatial distributions of background KL divergence for SPEEDY model and NICAM.

Upper panels show (a) zonal wind and (b) temperature at the second model level (~850 hPa) for the
SPEEDY model at 0000 UTC 1 March. Bottom panels show (c) zonal wind and (d) temperature at
the eighth model level (~850 hPa) for NICAM at 0000 UTC 8 November 2011.

659

660 Table. 1: CRPS and its three components (reliability, resolution and uncertainty) for background

661 specific humidity at the lowest model level (~925 hPa) from 0000 UTC 25 January to 1800 UTC 1

662	March.

	CRPS	Reli	Resol	U
	[g kg <sup>-1</sup> ]			
All grid points	0.0214	0.0000101	0.525	0.547
Grid points with	0.0475	0.0000244	0.030	0.077
non-Gaussian PDF				