

1 **A Local Particle Filter for High Dimensional Geophysical** 2 **Systems**

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10 **Abstract**

11 A local particle filter (LPF) is introduced that outperforms traditional ensemble Kalman filters
12 in highly nonlinear/non-Gaussian scenarios, both in accuracy and computational cost. The
13 standard Sampling Importance Resampling (SIR) particle filter is augmented with an
14 observation-space localization approach, for which an independent analysis is computed
15 locally at each gridpoint. The deterministic resampling approach of Kitagawa is adapted for
16 application locally and combined with interpolation of the analysis weights to smooth the
17 transition between neighboring points. Gaussian noise is applied with magnitude equal to the
18 local analysis spread to prevent particle degeneracy while maintaining the estimate of the
19 growing dynamical instabilities. The approach is validated against the Local Ensemble
20 Transform Kalman Filter (LETKF) using the 40-variable Lorenz-96 model. The results show
21 that: (1) the accuracy of LPF surpasses LETKF as the forecast length increases (thus
22 increasing the degree of nonlinearity), (2) the cost of LPF is significantly lower than LETKF
23 as the ensemble size increases, and (3) LPF prevents filter divergence experienced by LETKF
24 in cases with non-Gaussian observation error distributions.

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1 1 Introduction

2 The Particle Filter (PF) has been explored in the data assimilation community since the
3 introduction of its Gaussian linear variant, the Ensemble Kalman Filter (EnKF) in the mid-
4 1990's (*Evensen, 1994*). While general PFs have been intractable for high dimensional
5 systems, the EnKF has experienced great success in numerical weather prediction (NWP)
6 (*e.g. Kleist 2012; Hamrud et al., 2014*) and ocean data assimilation (*e.g. Penny et al. 2015*).
7 However, at least two limitations are on the horizon for EnKFs. Perhaps counter-intuitively,
8 these limitations arise due to *increased* computational resources, and have already become
9 challenges at the RIKEN Advanced Institute for Computational Science (AICS, *e.g.*,
10 *Miyamoto et al. 2013; Miyoshi et al. 2014; 2015*). First, global models will be pushed to
11 higher resolutions in which they begin to resolve highly nonlinear processes. To maintain the
12 Gaussian linear assumption required for the EnKF, much smaller timesteps are needed. For
13 example, the standard 6-hour analysis cycles used for the atmosphere may need to be
14 decreased to 5 minutes or even 30 seconds. Second, large ensembles (*e.g.* with ensemble size
15 $k > 10,000$ members) will become feasible for lower-resolution models. While at first this
16 may seem an advantage rather than a limitation, the computational cost of the local ensemble
17 transform Kalman filter (LETKF) (*Hunt et al., 2007*), for example, increases at a rate $O(k^3)$
18 with the ensemble size k . Thus as the ensemble size increases, the cost of computing the
19 analysis increases at a much greater rate. Alternative EnKFs feasible for large geophysical
20 systems scale in computational cost with the observation dimension, which is typically
21 multiple orders of magnitude larger than the ensemble dimension.

22 The PF is generally applicable to nonlinear non-Gaussian systems, including cases with multi-
23 modal distributions or nonlinear observation operators. With little difficulty, PFs can
24 explicitly include representation of model error, nonlinear observation operators (*Nakano*
25 *2007; Lei and Bickel 2011*), non-diagonal observation error covariance matrices, and non-
26 Gaussian likelihood functions. For example, observed variables such as precipitation are
27 inherently non-Gaussian and cannot be effectively assimilated by standard EnKF techniques
28 (*e.g. Lien et al. 2013; 2015*). In the expansion to sea-ice and land data assimilation
29 applications, the non-Gaussian quantities such as ice concentration, ice thickness, snow cover,
30 and soil moisture outnumber those that can be modeled with Gaussian error. *Bocquet et al.*
31 *(2010)* further review the difficulties using observations with non-Gaussian error
32 distributions. All of these problem-specific variations can create great difficulties for standard

1 methods, such as the EnKF or variational approaches (3D-Var/4D-Var), as used in current
2 operational systems.

3 Sampling Importance Resampling (SIR) (also known as the bootstrap filter, *Gordon et al.*,
4 1993) is a commonly used enhancement to the basic Sequential Importance Sampling (SIS)
5 particle filter. However, even with resampling the number of ensemble members required by
6 the SIR particle filter to capture the high probability region of the posterior in high-
7 dimensional geophysical applications is too large to make SIR usable (*Ades and van Leeuwen*
8 2013). *Snyder et al. (2008)* found that the number of required ensemble members scales
9 exponentially with the size of the system, giving the example that a 200 dimensional system
10 would require 10^{11} members. However, *Snyder et al.* note that clever choices of the proposal
11 distribution could overcome the need for these exponentially large ensemble sizes in high-
12 dimensional systems, which has been more recently explored by *Snyder et al. (2015)*.
13 Applying such an approach, *van Leeuwen (2003)* considers a model for the Agulhas Current
14 with dimension of roughly 2×10^5 . Further, *Beskos et al. (2012)* discuss recursive methods for
15 estimating the proposal densities, similar to the Running-in-Place algorithm (*Yang et al.*
16 2012a/b, *Penny et al. 2013*) that has been used with LETKF in meteorological and
17 oceanographic data assimilation. *Xiong et al. (2006)* presented techniques related to the ETKF
18 that may provide an alternative approach to the common SIR method for generating particle
19 estimates for the posterior distribution.

20 Techniques such as localization and inflation are typically applied as modifications to make
21 the EnKF operationally feasible. Inspired by this practice, we introduce a local particle filter
22 (LPF) designed for geophysical systems that is scalable to high dimensions and has
23 computational cost $O(k)$. Spatial localization is typically justified by the fact that long
24 distance correlations are either spurious or weak in comparison to nearby correlations,
25 particularly when the ensemble is under-sampled. We use this same approach to reduce the
26 required ensemble size for the LPF. The findings of *Jardak et al. (2000)* indicate that the
27 EnKF with localization works well in the case of a linear observation operator but has
28 difficulties with nonlinear observation operators.

29 **2 Methodology**

30 Localization is used in most operational NWP data assimilation systems, either through a
31 direct scaling of the background error covariance matrix (*e.g. Whitaker and Hamill, 2002*) or
32 by a scaling of the observation error covariance matrix (*Hunt et al. 2007*). Because the

1 computation of a background error covariance matrix is not needed for the PF, the latter
 2 approach is applied here to develop an effective PF for high-dimensional geophysical
 3 systems. Localization reduces the dimensionality of the solution space, thus requiring fewer
 4 ensemble members to sample the phase space. Gaussian noise is applied as an additive
 5 inflation to prevent particle degeneracy.

6 **2.1 The Standard SIR Particle Filter**

7 There are many variations of the PF (*Stewart and McCarty, 1992; Gordon et al., 1993;*
 8 *Kitagawa, 1996; Hurzeler and Kunsch, 1998; Liu and Chen, 1998*). In essence it is simply a
 9 Monte Carlo estimation of Bayes Theorem, reformulated as a recursion (*Doucet et al., 2001*),

$$10 \quad p(\mathbf{x}_t | \mathbf{y}_{1:t}) = \frac{p(\mathbf{y}_t | \mathbf{x}_t) p(\mathbf{x}_t | \mathbf{y}_{1:t-1})}{p(\mathbf{y}_t | \mathbf{y}_{1:t-1})}, \quad (1)$$

11 where $p(x_t | y_{1:t})$ is the probability of the state x at time t , given all observations y up to time t .
 12 We consider the model domain to be described by vectors x with dimension m , the
 13 observation domain to be described by vector y with dimension l , and an ensemble of size k .
 14 The term in the numerator can be expressed using the Chapman-Kolmogorov equation as,

$$15 \quad p(\mathbf{x}_t | \mathbf{y}_{1:t-1}) = \int p(\mathbf{x}_t | \mathbf{x}_{t-1}) p(\mathbf{x}_{t-1} | \mathbf{y}_{1:t-1}) dx_{t-1}, \quad (2)$$

16 and similarly the term in the denominator can be expressed as,

$$17 \quad p(\mathbf{y}_t | \mathbf{y}_{1:t-1}) = \int p(\mathbf{y}_t | \mathbf{x}_t) p(\mathbf{x}_t | \mathbf{y}_{1:t-1}) dx_t. \quad (3)$$

18 The two factors in the numerator of Equation (1) are sampled using a numerical model f ,

$$19 \quad p(\mathbf{x}_t | \mathbf{y}_{1:t-1}) \approx \frac{1}{k} \sum_{i=1}^k \delta(\mathbf{x}_t - f(\mathbf{x}_{t-1}^i)), \quad (4)$$

$$20 \quad p(\mathbf{y}_t | \mathbf{x}_t) = g(\mathbf{y}_t | \mathbf{x}_t). \quad (5)$$

21 The term in eqn. (5) is typically called the likelihood, because the probability of y given x is
 22 equivalent to the likelihood of x given y , i.e. $p(\mathbf{y} | \mathbf{x}) \equiv \ell(\mathbf{x} | \mathbf{y})$. The function g is general and
 23 can represent any distribution for the observations.

1 For the experiments here we generate two experiment cases, each with a different likelihood
 2 function. First we use a Gaussian likelihood function corresponding to that used for EnKFs,

$$3 \quad g(\mathbf{y}_t | \mathbf{x}_t) = \frac{1}{\sqrt{(2\pi)^m |\mathbf{R}|}} \exp \left[-\frac{1}{2} (\mathbf{y}_t - h(\mathbf{x}_t))^T \mathbf{R}^{-1} (\mathbf{y}_t - h(\mathbf{x}_t)) \right], \quad (6)$$

4 where the function h is a general, possibly nonlinear, observation operator mapping from the
 5 model state space to the observation space. For the LPF, it is straightforward to generalize to
 6 arbitrary non-Gaussian likelihood functions. As an example, we also apply a multivariate
 7 Gaussian mixture model (GM₂) following *Fowler and van Leeuwen (2013)* with pdf,

$$8 \quad p(\mathbf{y} | \mathbf{x}) \propto \nu_w \exp \left[(\mathbf{y} + \nu_1 \mathbf{1} - h(\mathbf{x}))^T \mathbf{R}^{-1} (\mathbf{y} + \nu_1 \mathbf{1} - h(\mathbf{x})) \right] \\
 + (1 - \nu_w) \exp \left[(\mathbf{y} + \nu_2 \mathbf{1} - h(\mathbf{x}))^T \mathbf{R}^{-1} (\mathbf{y} + \nu_2 \mathbf{1} - h(\mathbf{x})) \right]. \quad (7)$$

9 Let each ensemble member be identified with an index, i . Normalized weights are evaluated
 10 for each member,

$$11 \quad w_t^i = \frac{p(\mathbf{y}_t | \mathbf{x}_t^i)}{\sum_{j=1}^k p(\mathbf{y}_t | \mathbf{x}_t^j)}. \quad (8)$$

12 Then the posterior is,

$$13 \quad p(\mathbf{x}_t | \mathbf{y}_{1:t}) \approx \sum_{i=1}^k w_t^i \delta(\mathbf{x}_t - f(\mathbf{x}_{t-1}^i)). \quad (9)$$

14 Based on Liouville's theorem, the evolution of a probability measure in a dynamical system
 15 satisfies the property that "the probability of finding trajectories inside the time-variant
 16 volume $W(t)$ is constant during the evolution of the dynamical system." (Property 2,
 17 http://www.ulb.ac.be/di/map/gbonte/ftp/bontempi_fpde.pdf). If the solution manifold expands
 18 in some directions, so will the pdf represented by the particles. Thus, the fidelity of the
 19 distribution will quickly become insufficient to sample a solution manifold around the true
 20 trajectory. A resampling procedure is used to refocus the particles on the densest areas of the
 21 distribution at each analysis step. For the experiments here, we use a resampling procedure
 22 that resembles resampling with replacement. After resampling we have,

$$p(\mathbf{x}_t | \mathbf{y}_{1:t}) \approx \frac{1}{k} \sum_{i=1}^k \delta(\mathbf{x}_t - \mathbf{x}_t^i) \quad (10)$$

2.2 The Transform Interpretation

The PF can be interpreted similarly to the Ensemble Transform Kalman Filter (ETKF) of Bishop et al. (2001). The transform interpretation has been explored by Reich 2013 and Metref et al. 2014. Namely we define the PF solution as a transformation of the background ensemble to the analysis ensemble,

$$\mathbf{X}^a = \mathbf{X}^b \mathbf{T} \quad (11)$$

where each column of \mathbf{X}^b is a background ensemble member, and each column of \mathbf{X}^a is an analysis ensemble member. The transform function \mathbf{T} in (11) is used to generically refer to either \mathbf{E} in (16), or $\mathbf{E}^{(j)}$ in (19) in the text below.

Let \mathbf{b} be the vector of background particle indices and \mathbf{a} be the vector of analysis particle indices,

$$\mathbf{b} \in \left\{ \mathbf{z} \in \mathbb{Z}^k \mid \mathbf{z} = (1, 2, 3, \dots, k)^T \right\}, \quad (12)$$

$$\mathbf{a} \in \left\{ \mathbf{z} \in \mathbb{Z}^k \mid \mathbf{z} = (a_1, a_2, a_3, \dots, a_k)^T, a_i \in [1, k] \right\}. \quad (13)$$

Here, \mathbf{b} is simply vector composed of an ordered set of indices ranging from 1 to k , while \mathbf{a} is a vector composed of a subset of elements of the vector \mathbf{b} , with no particular ordering and possible repetitions. We also extend the definition of \mathbf{a} to allow the vector to vary by grid point. For example, for any grid point p , the vector \mathbf{a}_p may represent the resampling index at point p so that,

$$\mathbf{X}^a(p, i) = \mathbf{X}^b(p, a_{p,i}), \quad (14)$$

where the arguments indicate the row and column of the matrix \mathbf{X} .

We further define the vector \mathbf{e}_i , for which only the i^{th} element is nonzero,

$$\mathbf{e}_i \in \left\{ \mathbf{z} \in \mathbb{Z}_2^k \mid \mathbf{z} = (0, \dots, 0, 1_i, 0, \dots, 0)^T \right\}. \quad (15)$$

If \mathbf{e}_i are the canonical basis vectors then we can define,

$$\mathbf{E}_{k \times k} = \begin{bmatrix} \mathbf{e}_{a_1} & \mathbf{e}_{a_2} & \cdots & \mathbf{e}_{a_k} \end{bmatrix}. \quad (16)$$

For the standard PF, the indicator matrix \mathbf{E} is made up of k (not necessarily unique) standard basis vectors \mathbf{e}_i , with entries 0 and 1 that we will interpret as weights. Thus the analysis ensemble for the PF is defined simply by the transform,

$$\mathbf{X}_{m \times k}^a = \mathbf{X}_{m \times k}^b \mathbf{E}_{k \times k}. \quad (17)$$

We note that by using this approach, each new analysis member, with index i , maintains the continuity properties of its associated background member, a_i .

For reference in the next section, the components of the analysis matrix \mathbf{X}^a will have the form,

$$\mathbf{X}^a = \begin{bmatrix} x_{1,i} e_{i,1} & x_{1,i} e_{i,2} & \cdots & x_{1,i} e_{i,k} \\ x_{2,i} e_{i,1} & x_{2,i} e_{i,2} & \cdots & x_{2,i} e_{i,k} \\ \vdots & \vdots & \ddots & \vdots \\ x_{m,i} e_{i,1} & x_{m,i} e_{i,2} & \cdots & x_{m,i} e_{i,k} \end{bmatrix}. \quad (18)$$

Here we have used the Einstein tensor notation for the elements, in which $x_{1,i} e_{i,1}$ represents a summation over the index i (i.e. the inner product of row 1 of \mathbf{X}^b and column 1 of \mathbf{E}). While the summation index could be represented generically by any symbol, we reuse the symbol ‘ i ’ due to its correspondence with the background particle indices as defined above.

2.3 The Localization Approach

Snyder et al. (2008) note that when either the model dimension or observation count is large, the PF requires significantly more particles to give an adequate representation of the system. Localization, as introduced by Houtekamer and Mitchell (1998), reduces both the model and observation dimensions by dividing the problem into a series of sub-domains, thus reducing the required number of particles for accurate filtering. Bengtsson et al. (2003) were among the first to point to spatially local updating, using a local subset of observations, as a solution to difficulties of high-dimensional non-Gaussian filtering. *Lei and Bickel (2011)* introduced the notion of computing local weights in a non-Gaussian filter. The LPF follows *Hunt et al. (2007)* to select nearby observations for independent analysis at each grid point. Nearby grid points thus assimilate nearly identical sets of observations to derive their analyses.

1 We use the deterministic resampling of *Kitagawa (1996)*, with complexity $O(k)$,
 2 adapted for local use as described next. A uniform partition of $[0,1]$ with width $1/k$ is first
 3 generated globally, with an offset applied from a uniform distribution over $[0,1/k]$. The same
 4 partition is used locally for resampling at each grid point. Cumulative sums of the normalized
 5 weights (eqn. 8),

$$6 \quad \tilde{w}_t^j = \sum_{i=1}^j w_t^i, \quad (19)$$

7 are compared with the elements of the partition. Traversing $j=1, \dots, k$ all unassigned particles
 8 with index j having a corresponding cumulative sum with index j that surpasses the next
 9 element of the partition (ordered monotonically increasing) are assigned as particles of the
 10 resampled (analysis) ensemble.

11 For a given grid point, when the cumulative sums of the particle weights are near one
 12 of the partition values, there may be sensitivity in neighboring grid points that lead to
 13 discontinuities between local analyses. The analysis ensemble at this grid point consists of a
 14 subset of background particle indices (1 through k) with repetitions. To eliminate the
 15 discontinuities with neighboring grid points we associate weights of a local transform
 16 function T with the particle indices, nominally either 1.0 (full weight) or 0.0 (no weight). This
 17 is partially inspired by the “weight interpolation” of *Bowler (2006)*, applied to LETKF by
 18 *Yang et al. (2009)*, who found that interpolation of weights was more robust than interpolation
 19 of state values. At a single grid point, there are k pieces of background information about the
 20 possible system state at that point. In the standard PF, only 1 out of these k pieces of
 21 information is retained for each analysis ensemble member, based on the overall agreement
 22 with observations. In the LPF we use anywhere from 1 to k members to construct each
 23 analysis member based on the agreement with observations within a local radius.

24 Thus the modified local analysis at a given model grid point p is given as a linear
 25 combination of ensemble members at that point. For point p , ensemble member i , a set of N
 26 neighbor points N_p , and vector \mathbf{a}_p acting as a resampling index as in (13) but given as a
 27 function of the model grid point, we have,

$$28 \quad \mathbf{X}_{\{p,i\}}^{LPF} = \frac{1}{2} \mathbf{X}_{\{p,i\}}^a + \frac{1}{2N} \sum_{n \in N_p} \mathbf{X}_{\{p,\mathbf{a}(n,i)\}}^b. \quad (20)$$

29 The subscript indices indicate the row and column of the matrix.

1 A new transform can then be defined for the LPF at each point in the model domain to
 2 generate a set of m indicator matrices, $\mathbf{E}_{k \times k}^{(j)}$, so that for each point (x_p , for $p=1, \dots, m$),

$$3 \quad (\mathbf{X}_{m \times k}^a)^p = \mathbf{X}_{m \times k}^b \mathbf{E}_{k \times k}^{(p)}. \quad (21)$$

4 Using the summation tensor notation described in the previous section, the analysis ensemble
 5 can be written,

$$6 \quad \mathbf{X}^a = \begin{bmatrix} x_{1,i} e_{i,1}^{(1)} & x_{1,i} e_{i,2}^{(1)} & \cdots & x_{1,i} e_{i,k}^{(1)} \\ x_{2,i} e_{i,1}^{(2)} & x_{2,i} e_{i,2}^{(2)} & \cdots & x_{2,i} e_{i,k}^{(2)} \\ \vdots & \vdots & \ddots & \vdots \\ x_{m,i} e_{i,1}^{(m)} & x_{m,i} e_{i,2}^{(m)} & \cdots & x_{m,i} e_{i,k}^{(m)} \end{bmatrix}. \quad (22)$$

7 The transform matrix may have any degree of sophistication. We apply a smoothing operator
 8 to the ensemble subspace by modifying the weights \mathbf{e}_{a_i} associated with each analysis particle
 9 index $\mathbf{a}_{(i)}$ from a binary value to a continuous value between 0 and 1, while maintaining all
 10 column sums of equal to one, and call this new transform matrix \mathbf{W} .

11 We define the concept of a ‘neighbor point’ abstractly as a point near the analyzed grid
 12 point based on a specified distance metric. Through examination of (20), it is clear that the
 13 choice of neighbor points simply informs the weighting of indices, and that the values at these
 14 neighbor points otherwise have no impact on the analysis. If there are N neighbor points, then
 15 there will be at most $\min(N+1, k)$ collocated pieces of background information that can be
 16 utilized to construct each analysis ensemble member at this point. An example is given in
 17 **Figure 1**. With a sufficiently large set of observations the indices for these neighbor points
 18 are calculated from nearly identical observational innovations. Therefore, when there is a
 19 sufficiently large ensemble size (k) the difference between the states associated with different
 20 particle indices will be small. The transform function T is applied across all background
 21 indices (i.e. for particles $1, \dots, k$) at this grid point to compute the analysis.

22 We summarize that the total smoothing in the resulting analysis is achieved
 23 independently at each grid point due to a combination of localization in observation space and
 24 the formation of a convex combination of analysis weights in the ensemble space. There is no
 25 explicit smoothing in the physical model space such as might occur using a Gaussian
 26 smoother applied via a stencil of function values from neighboring grid points. Instead, the

1 neighbor points simply inform the choice of weights to apply to particle indices at a single
2 grid point. There is an implicit smoothing achieved by applying the same procedure to many
3 contiguous model grid points, each generating a different set of weights that vary only
4 slightly. This is similar to the effect of observation-space localization. However, as with most
5 localization techniques, the more distant information suffers from the poor sampling size of
6 the ensemble. Thus if holding the ensemble size k fixed, a larger radius of neighbor points
7 may lead to noisier results rather than increased smoothing.

8 **2.4 Particle Degeneracy**

9 The particle selection process of the PF reduces the rank of the ensemble. For a linear
10 deterministic system this leads to a rapid collapse of the ensemble and divergence of the filter.
11 For a sufficiently stochastic nonlinear system the members are made distinct after a single
12 forecast step. If the nonlinear system is not sufficiently stochastic, then we must address the
13 ensemble initialization problem at every analysis cycle. *Pazo et al. (2010)* discuss the
14 desirable properties in an initial ensemble, namely the members: (1) should be well-embedded
15 in the attractor, (2) should be statistically equivalent but have enough diversity to represent a
16 significant portion of the phase space, (3) should adequately represent the error between the
17 analysis and true state, and (4) should sample the fastest growing directions in phase space.
18 We wish to avoid particle degeneracy while also engendering some of these qualities.
19 Applying noise to the PF at the sampling step is a common empirical technique. Therefore we
20 employ a simple approach: at each cycle we add Gaussian noise with variance scaled locally
21 to a magnitude matching the analysis error variance and apply this to each analysis member
22 prior to the subsequent ensemble forecast. The amplitude of the additive noise was chosen to
23 conform to the dynamics of the growing error subspace, as estimated by the analysis
24 ensemble spread. We note that this amplitude varies spatially and temporally. The results
25 degraded when departing from this approach. The practice of applying noise to the PF
26 sampling step is a standard technique.

27 We caution that EnKFs have fundamentally different behavior than the general PF, in
28 that the former maintain a forcing term that drives the DA system toward the observations
29 even when the forecast may start far from the true state. The general PF with resampling
30 essentially requires random chance to generate a state with high enough probability to be
31 propagated through the resampling step. As the system size increases the probability of such
32 an event declines.

1 **2.5 Computational Complexity**

2 A data assimilation system is comprised of many components. We simplify the cost analysis
3 in order to gain an approximate relative measure of the algorithms presented here. Let m be
4 the model dimension, l be the observation dimension, and \bar{l}_i be the average local observation
5 dimension. The total cost (C_T) of an analysis cycle is equal to the overhead (C_H) of the
6 assimilation system plus m times the average local cost (C_L) of the assimilation method plus k
7 times the cost of one model forecast (C_M) of duration t ,

$$8 \quad C_T(k, l, m) = C_H(k, l, m) + m \cdot C_L(k, \bar{l}_i) + k \cdot C_M(\tau, m) \quad (23)$$

9 We will assume that between the two methods the overhead and model costs are
10 approximately equal. The primary difference in cost between the two systems is then the
11 average local cost,

$$12 \quad C_L^{LPF} = O(k\bar{l}_i), \quad (24)$$

$$13 \quad C_L^{LETKF} = O(k^2\bar{l}_i + k^3). \quad (25)$$

14 If as is typically the case, the system size m is large and the ensemble size k is small, then

$$15 \quad C_T(k, l, m) \approx C_H(k, l, m) + O(m), \quad (26)$$

16 and the difference in cost between LETKF and LPF is small. However for large k , we see that
17 the average local cost of LETKF,

$$18 \quad C_T^{LETKF}(k, l, m) = C_H(k, l, m) + O(mk^3), \quad (27)$$

19 exceeds that of the LPF,

$$20 \quad C_T^{LPF}(k, l, m) = C_H(k, l, m) + O(mk). \quad (28)$$

21 Subtracting the overhead costs, in this case the LPF is a factor of k^2 cheaper than LETKF.

22 **2.6 Data Assimilation Methods**

23 We enumerate the benefits of the LPF versus the benchmark LETKF, an ensemble square root
24 filter that performs its analysis in the ensemble space at each grid point using a geospatially

1 local selection of observations. The LETKF approach is very efficient as long as the ensemble
2 size is small relative to the number of observations and the model dimension.

3 We use LETKF as a proxy for a general EnKF. *Nerger (2015)* gives a comparison between
4 LETKF and the Ensemble Square Root Filter (ESRF) of *Whitaker and Hamill (2002)*, while
5 *Tippett et al. (2003)* indicate that the ESRF is identical to the Ensemble Adjustment Kalman
6 Filter (EAKF) of *Anderson (2001)* when using serial single observation processing.

7 **2.7 Experiment Design**

8 We demonstrate the algorithms on the Lorenz-96 system (*Lorenz 1996*), composed of $m=40$
9 grid points, using Lorenz’s original forcing, $F=8.0$. The Lorenz-96 system has frequently been
10 used to demonstrate PF and other data assimilation algorithms for the geosciences (*Nakano et*
11 *al, 2007; van Leeuwen, 2010; Lei and Bickel, 2011; Penny 2014*). Observations are sampled
12 from a nature run of Lorenz-96 after running the model for 14,400 timesteps to allow the
13 model to settle on the attractor. Gaussian noise is added to each observation with a standard
14 deviation of 0.5. For various experiments, the Lorenz-96 system is observed either at every
15 0.05 or 0.5 timesteps, reflective of a 6-hour and 60-hour forecast, respectively, based on
16 Lorenz’s original description of the system. Observations are sampled randomly on the
17 interval $[0,m]$, and a linear interpolation is used for the observation operator. The last
18 experiment case uses a bimodal Gaussian mixture distribution to represent observational
19 error.

20 **3 Results**

21 The standard SIR PF performs poorly with any ensemble size $O(m)$. For example, using 1500
22 particles and 20 randomly chosen observations per analysis cycle leads to rapid filter
23 divergence for the L96 system, even in a relatively linear regime ($dt=0.05$) of the system
24 (**Figure 2**). On the contrary, LETKF performs well even with few ensemble members and few
25 observations per cycle ($k=20, l=10$). Localization is consistent between each method, using
26 $r=2$ gridpoints. For a given ensemble size, increasing the localization radius degraded the
27 accuracy of both methods. Penny (2014) addressed the impact of varying localization radius
28 on LETKF for this L96 system. It was shown that for a fixed ensemble size, the analysis
29 errors reduce when the localization radius decreases, and for a fixed localization radius, the
30 analysis errors reduce when the ensemble size increases. The LPF showed similar behavior to
31 LETKF in this regard. To explore the relative advantages of each approach, we will describe a

1 series of cases in which the LETKF outperforms the LPF, and in which the LPF outperforms
2 LETKF.

3 **3.1 Case 1: typical forecast lengths (dt=0.05, or 6-hr)**

4 *Lorenz (1996)* introduced the $dt=0.05$ timescale as being comparable to the error doubling
5 taking place over 6 hours in the operational forecasting systems of the early 1990's. In this
6 relatively linear timescale of the L96 system, LETKF clearly outperforms the LPF at a given
7 ensemble size. This is expected as EnKFs take advantage of the Gaussian/linear assumption.
8 When the experiment parameters match such assumptions (even loosely), LETKF performs
9 quite well. However, using localization, the LPF can perform adequately (i.e. avoid filter
10 divergence) in a similar parameter regime (**Figure 3**). Thus for this case, we find that LETKF
11 attains higher accuracy than the LPF, but the LPF improves upon the accuracy and stability of
12 the standard SIR PF for a given ensemble size.

13 **3.2 Case 2: long forecast lengths (dt=0.50, or 60-hr)**

14 To increase the degree of nonlinearity in a data assimilation system using L96, it is typical to
15 increase the analysis cycle length (e.g. *Lei and Bickel, 2011*). The LPF has superior
16 performance for more nonlinear regimes of the L96 system (e.g. $dt=0.5$) provided there are
17 many ensemble members, e.g. $O(100)$. Using 80 observations per cycle and 100 ensemble
18 members, for example, LETKF produces occasional errors that propagate eastward (along the
19 positive x-direction). The LPF does not produce such effects, and the errors are generally
20 lower than with LETKF (**Figure 4**). We consider this a relevant scenario because the majority
21 of observations in operational weather forecasting are discarded (*Ochatta et al., 2005*).

22 Exploring a more complete parameter space, we examine the forecast error for LETKF
23 over a range of observation coverage ($l=2, \dots, 80$ per analysis cycle) and ensemble sizes
24 ($k=10, \dots, 400$), and compare the relative difference to LPF. **Figure 5** shows the average
25 absolute error over 600 analysis cycles of length $dt=0,5$ for 1600 different parameter
26 combinations of observation coverage (l) and ensemble size (k). The LPF is more accurate
27 than LETKF when using many observations (e.g. $l > 20$) and large ensemble sizes (e.g. $k >$
28 $100-200$). When using fewer than 20 observations per cycle in this case, both LETKF and
29 LPF experience filter divergence. Due to the unconstrained nature of the LPF as described in
30 section 2.4, large errors occur more frequently than for LETKF in this parameter regime.

1 Further, when examining the computational cost of the LPF versus LETKF, the
 2 relative costs reflect the analytical assessment given above in section 2.5. Namely, the elapsed
 3 time of the LETKF experiments grows with the cube of the ensemble size, while the elapsed
 4 time of the LPF is significantly lower at large ensemble sizes (**Figure 6**).

5 **3.3 Case 3: Non-Gaussian observation error.**

6 The previous section examined the impacts of nonlinearity and non-Gaussianity on the
 7 forecast. We now examine the impacts of non-Gaussian observation error. Using a
 8 multivariate Gaussian mixture model (GM₂) following *Fowler and van Leeuwen (2013)*, we
 9 apply a corresponding random error to the observation and compare the impacts on LETKF
 10 and LPF. We use the LETKF without modification, but modify the likelihood function of LPF
 11 to the definition of GM₂ as in section 2.1, Eqn. (7). Using $n_w=0.1$, $n_l=-1$, $n_2=1$, we create a
 12 bimodal distribution biased toward the second Gaussian mode. The analysis cycle is $dt=0.05$
 13 (6 hr) as in experiment case 1, section 3.1. **Figure 7**, compares LETKF and LPF using $k=100$
 14 ensemble members and $l=80$ observations. An additional results is given for LPF with $l=20$
 15 observations. The introduction of a strong non-Gaussianity in the observation error
 16 distribution disrupts LETKF and eventually creates errors that propagate throughout the entire
 17 domain. Using the same ensemble size and observation count, the LPF gains significant
 18 advantage in its ability to explicitly account for the non-Gaussian error structure of the
 19 observations. Even reducing the observation count by 75%, the LPF maintains its advantage.

20 **3.4 Case 4: Examining the impact of spatial smoothing.**

21 We consider the example of LPF applied to L96 with the analysis cycle $dt=0.5$ (producing a
 22 relatively non-linear error growth), a localization radius of 2 grid points, $k=100$ ensemble
 23 members, $l=40$ observations per cycle, and observation error of 0.5. By applying ensemble
 24 smoothing of the weights, we find that the mean absolute forecast error (averaged over the
 25 model domain) is reduced versus a non-smoothed analysis (**Figure 8**). Increasing the search
 26 radius to 2 gridpoints for the same example case does not produce any clear benefits.

27 Considering the impact on the ensemble, we use the smoothed analysis the mean
 28 effective ensemble size N_{eff} ,

$$29 \quad N_{eff} = \left[\sum_{i=1}^j (w_i^i)^2 \right]^{-1}, \quad (29)$$

1 with w_i^j defined as in (8), calculated at each gridpoint and averaged. We find that the effective
2 ensemble size is increased when using the smoothed versus the non-smoothed approach
3 (Figure 9).

4 **4 Conclusions**

5 The Local Particle Filter (LPF) has been shown to outperform a state of the art ensemble
6 Kalman filter (i.e. LETKF) in scenarios that violate the Gaussian/linear assumptions of the
7 Kalman filter. We showed the advantage of the LPF when forecast is more non-linear (via
8 longer analysis cycles, or less frequent observations), and when observation error is non-
9 Gaussian (using a bimodal error distribution). Further, upon transitioning to large ensembles
10 the LPF has a significant cost savings relative to LETKF.

11 The LPF maintains many of the attractive qualities that give Particle Filters (PFs)
12 advantages over standard EnKFs. While the PF provides a means of assimilating observations
13 with non-Gaussian errors (e.g. precipitation, sea-ice concentration), we caution that the
14 covariances utilized by the EnKF play a critical role in constraining the unobserved variables.
15 Thus while the LPF is not optimal for all possible data assimilation scenarios, there is great
16 potential for the LPF to be combined with more traditional approaches to create adaptive
17 hybrid systems that can avoid catastrophic filter divergence and manage multi-modal forecast
18 distributions, nonlinear observation operators, non-Gaussian observations.

19 We found that a large number of ensemble members (or particles) and observations
20 are sufficient for the LPF to match or surpass the accuracy of LETKF. We found large
21 ensemble sizes a relevant scenario for realistic systems running on large supercomputers such
22 as the K computer at RIKEN. The use of large observation sets is relevant in operational
23 weather forecasting as much of the dense satellite data is currently discarding in a thinning
24 process. Further, in this parameter regime the LPF has significantly lower computational cost
25 than LETKF.

26 In a realistic system, some mechanism is needed to drive the ensemble toward the
27 observations in the event of the ensemble drifting away from the true state. The PF itself has
28 no inherent mechanism to do this other than the brute force generation of more particles.
29 There are many techniques in the PF literature for managing filter divergence, but none of
30 them are foolproof. *Atkins et al. (2013)* presented a promising extension of the use of an
31 importance density that may connects effectively with the existing infrastructure of

1 variational solvers used by most operational centers. Another popular mechanism to achieve
2 this is regularization, which uses a kernel to sample from a continuous distribution at the
3 resampling stage.

4 Finally, while the inflation mechanism used here was effective for the Lorenz-96
5 system, it is not adequate for more realistic atmospheric or oceanic models. For such systems,
6 either geospatially correlated noise or stochastic physics parameterizations may be capable of
7 performing the same function. Stochastic physics parameterizations are an active area of
8 research, and are under development for a number of operational center models, including
9 NCEP (*Hou et al., 2006; Hou et al, 2010; Kolczynski et al., 2015*), ECMWF (*Berner et al.,
10 2009; Weisheimer et al., 2014; Watson et al., 2015*), and the Met Office (*Tennant et al., 2011;
11 Sanchez et al., 2014; Shutts and Paleres, 2014; Shutts 2015*).

12 **Acknowledgements**

13 We gratefully acknowledge the Japan Society for the Promotion of Science (JSPS) whose
14 FY2013 fellowship supported this work. We would also like to thank the RIKEN Advanced
15 Institute for Computational Science (AICS) for hosting Penny.

16 **Appendix**

17 We provide pseudo-code for the LPF:

```
18 % INPUTS:  
19 % k    :: ensemble size  
20 % m    :: model dimension  
21 % Yo   :: vector of global observations  
22 % yo   :: vector of local observations  
23 % Xb   :: background ensemble arranged as a matrix  
24 % R    :: observation error covariance matrix  
25 % HXb  :: Xb mapped from model space to observation space  
26 %  
27 % FUNCTIONS:  
28 % obsloc :: finds observations local to a grid point  
29 % cumsum :: cumulative sum of array  
30 % sort   :: sorting of array in descending order  
31 % repmat :: repeat array to form a matrix  
32 %  
33 % OUTPUTS:  
34 % Neff  :: effective ensemble sample size
```

```

1  % Xa  :: analysis ensemble arranged as a matrix
2
3  %
4  % Specify minimum of distribution tails
5  %
6  mintail = epsilon
7
8  %
9  % Setup global comb for resampling with replacement
10 %
11 interval = 1/k
12 start = interval * random_number
13 selection_points = [start : interval : start+(k-1)*interval]
14
15 %
16 % Loop over each grid point
17 %
18 for mi=1:m
19     % Find the observation points within range of focal grid point:
20     yo = obsloc(mi,Yo,local_radius)
21
22     % Update (calculate particle weights) using desired distribution
23     % (e.g. Gaussian shown here)
24     for ki=1:k
25         likelihood(ki) =
26             exp( -0.5* (yo-HXb(:,ki))' * R^{-1} * (yo-HXb(:,ki)) )
27     end
28
29     % Protect against numerical representation problems in the tails
30     likelihood = likelihood + mintail
31
32     %
33     % Normalize the weights
34     %
35
36     % Compute the weights
37     weights = likelihood/sum(likelihood)
38
39     % Calculate effective ensemble sample size
40     Neff(mi) = 1/sum(weights.^2)

```

```

1
2     % Form cumulative distribution
3     [wts, wi] = sort(weights)
4     weight = cumsum(wts)
5
6     %
7     % Apply the comb to resample analysis members
8     %
9     j=1
10    for i=1:k
11        while selection_points(i) >= weight(j)
12            j=j+1
13        end
14
15        % Specify the resampling index rs
16        rs(mi,i)=wi(j)
17
18        % Assign background value as global analysis
19        Xa(mi,i) = Xb(mi,rs(mi,i))
20    end
21
22 end % loop over model grid points
23
24 %
25 % Apply smoothing by weights in ensemble space.
26 %
27 x = 0
28 for ki=1:k
29     for mi=1:m
30         for ni={"set of all Np neighbor points"}
31             X(mi,ki) = X(mi,ki) + Xa(mi,ki) + Xb(mi,rs(ni,ki))
32         end
33         Xa(mi,ki) = X(mi,ki)/(2*Np);
34     end
35 end
36
37 %
38 % Apply additive inflation
39 %
40
41 % Compute the analysis ensemble spread to

```

```

1  % determine amplitude of inflation needed
2  if (mean(Neff) > k/2)
3      Xstd=std(Xa,2)
4  else
5      % add a little more noise to protect
6      % against ensemble collapse
7      Xstd=max(std(Xa,2),maximum_obs_error)
8  end
9
10 % Compute Gaussian random values with standard
11 % deviation equal to analysis ensemble spread
12 rmat=randn(m,k).*repmat(Xstd,1,k)
13
14 % Apply additive inflation (and remove sample mean)
15 Xa = Xa + rmat-repmat(mean(rmat,2),1,k)
16

```

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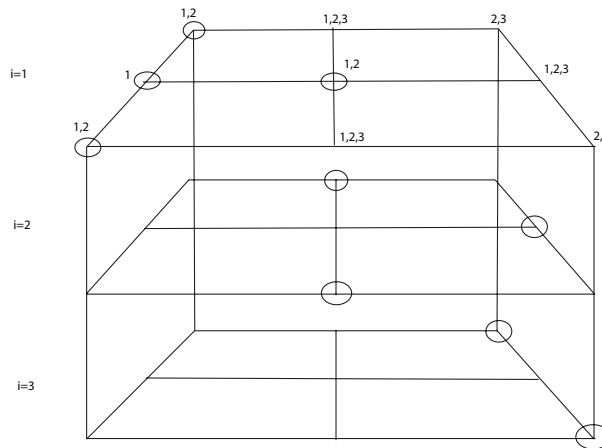
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1 Figures

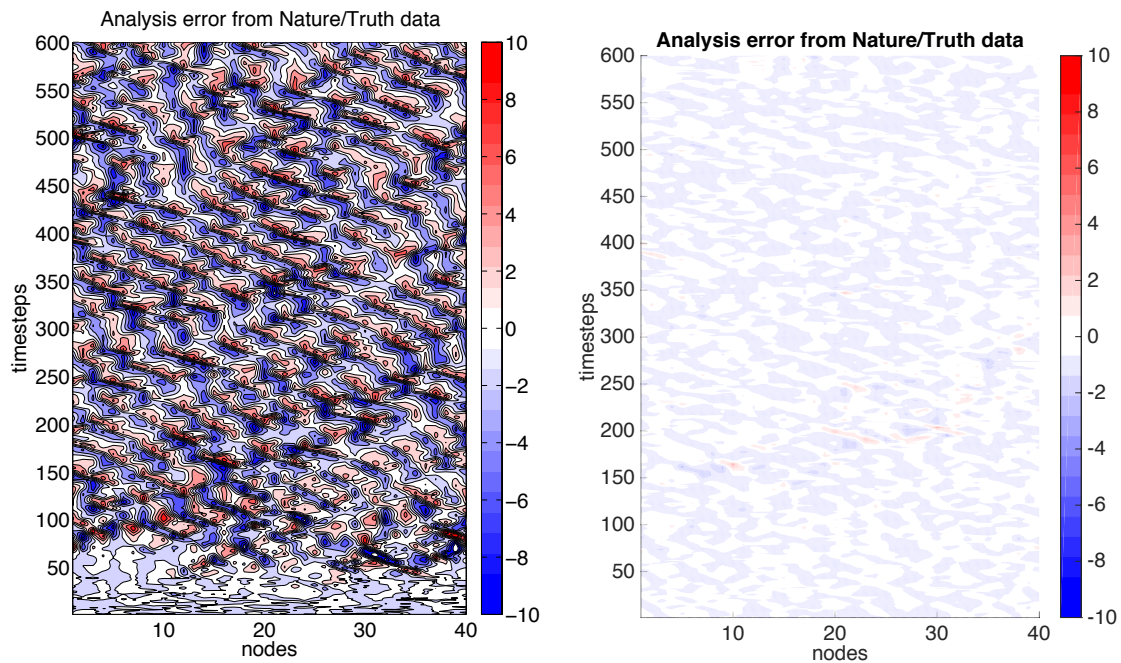
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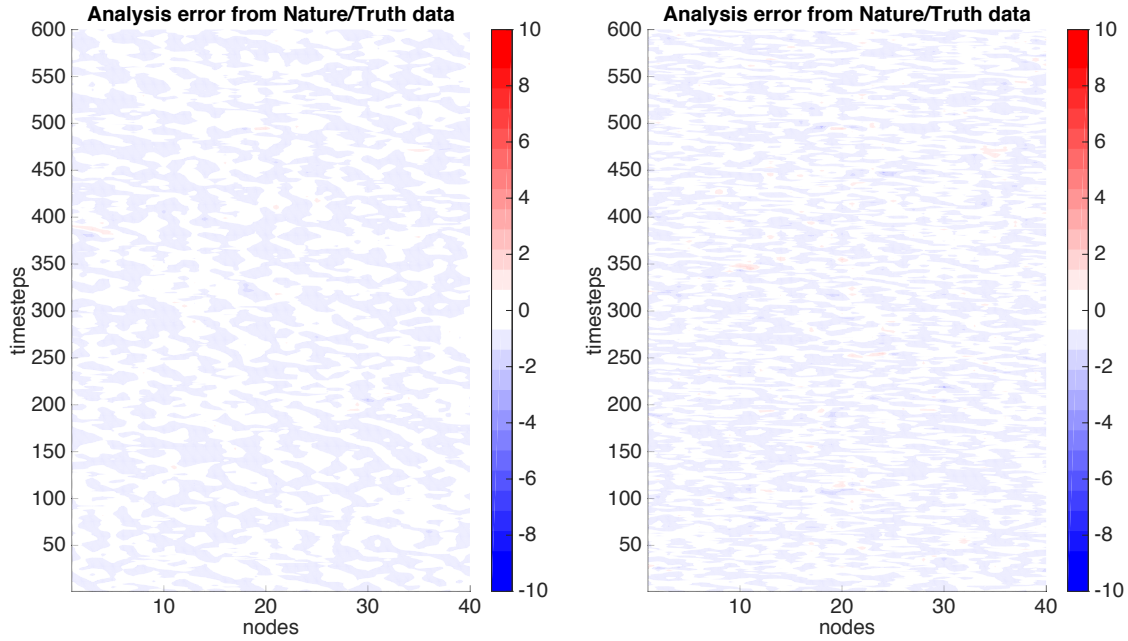
4 **Figure 1.** A hypothetical example depicting the construction of a single analysis member.
5 Each level represents a different background ensemble member (particle), with a model space
6 composed of a 3x3 grid. The nodes of the grid are circled if the member is chosen for the
7 construction of analysis member 1 by the LPF. The numerals indicate the ids for the
8 background members that will be averaged at the corresponding node, in this case based on
9 the immediately adjacent neighbor points of that node.

10

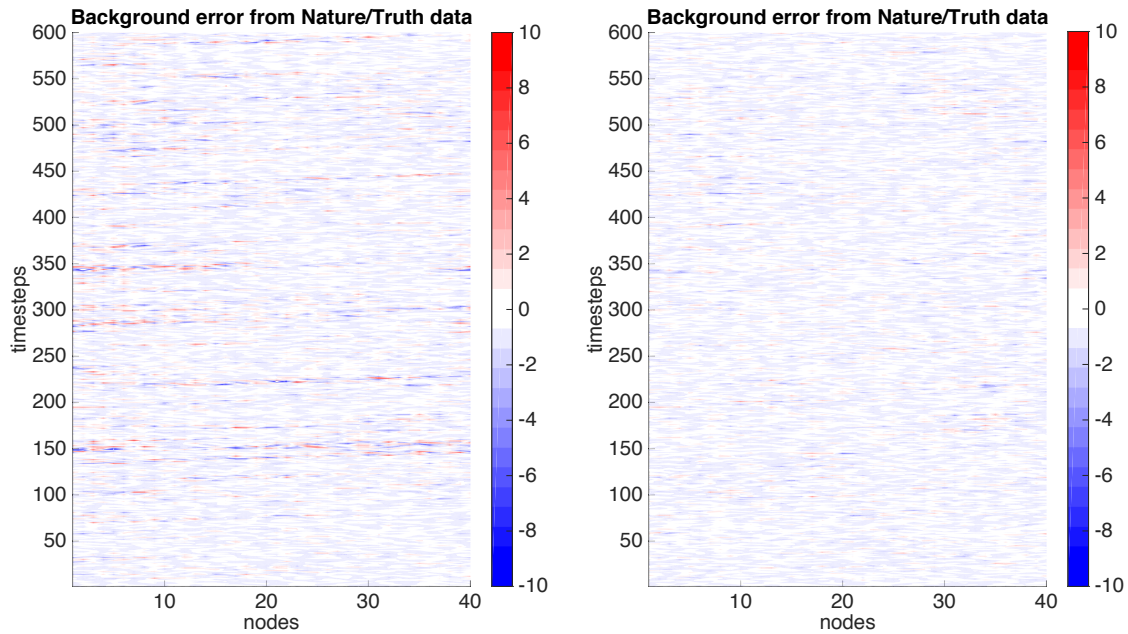


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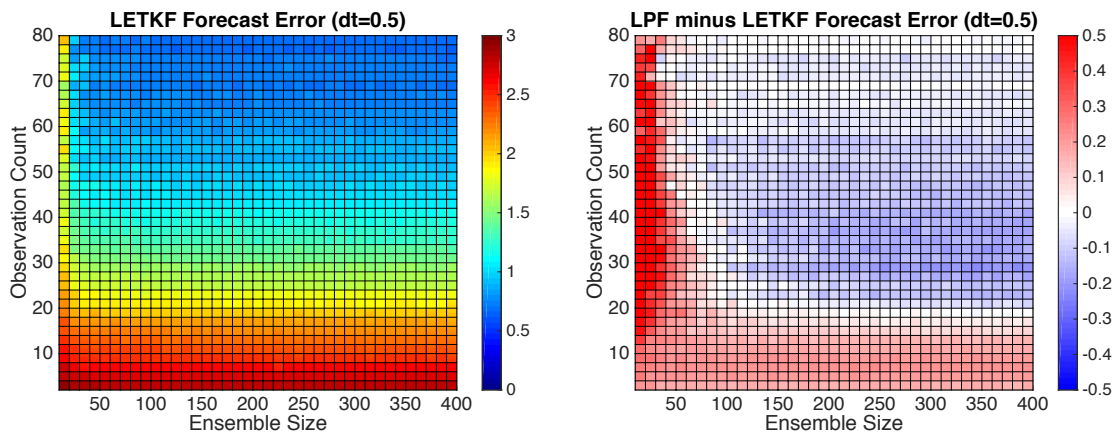
1 **Figure 2.** Analysis error using an analysis cycle window length $dt=0.05$ (6-hr) for (a) the
2 standard SIR PF using $k=1500$ particles with $l=20$ observations per analysis cycle, and (b)
3 LETKF with localization radius $r=2$, $k=20$ ensemble members, and $l=10$ observations per
4 analysis cycle, sampled randomly on the domain.



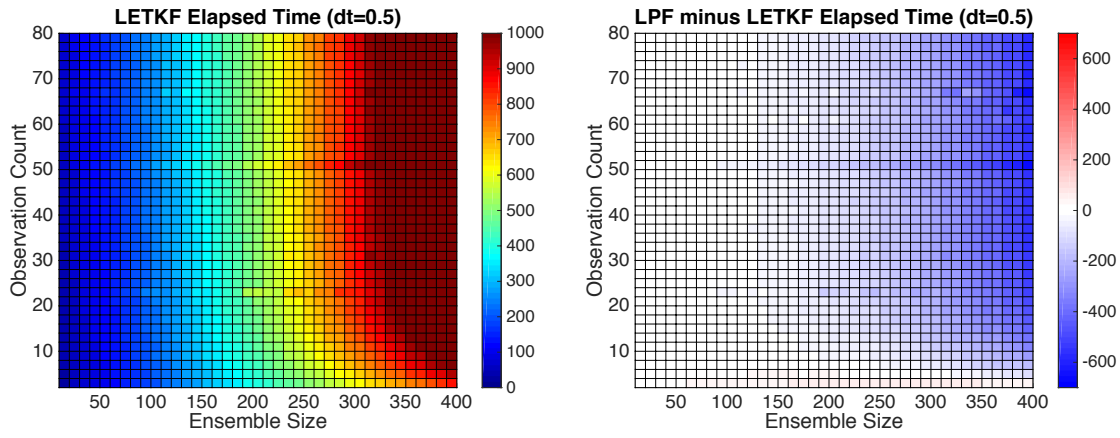
5
6 **Figure 3.** Analysis error for (a) LETKF and (b) LPF, using an analysis cycle window length
7 $dt=0.05$ (6-hr), localization radius $r=2$ grid points, $k=40$ ensemble members, and $l=20$
8 observations sampled randomly on the domain (prescribed observation locations and errors
9 are identical for both methods).



1
 2 **Figure 4.** Forecast error for (a) LETKF and (b) LPF, using an analysis cycle window length
 3 $dt=0.5$ (60-hr), localization radius $r=2$ grid points, $k=100$ ensemble members, and $l=80$
 4 observations sampled randomly on the domain (observation locations and errors are identical
 5 for both methods).

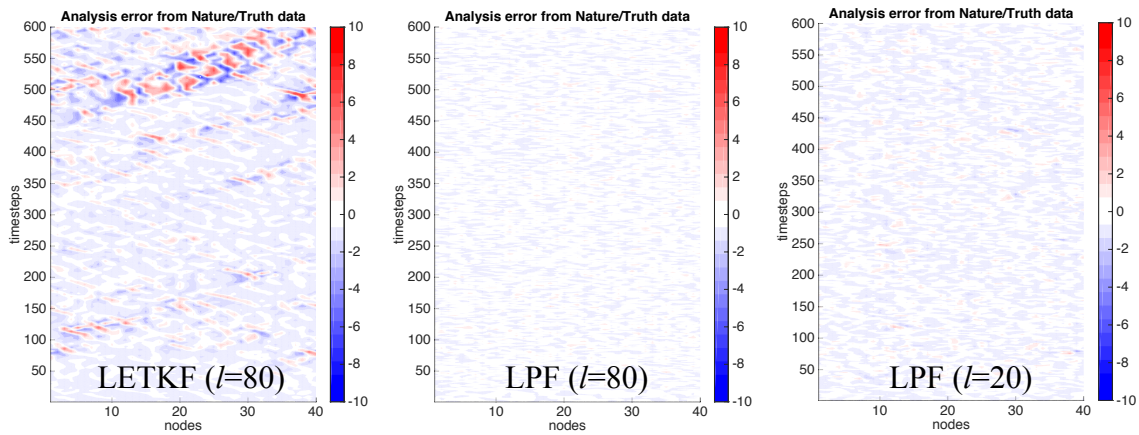


6
 7 **Figure 5.** Forecast error for (a) LETKF and (b) LPF minus LETKF. LPF reduces error in
 8 highly sampled cases with larger observation coverage. The LPF increases error in poorly
 9 sampled cases and with low observation coverage. Each cell represents one experiment case;
 10 absolute errors are averaged over the entire domain for 600 analysis cycles for each case.



1

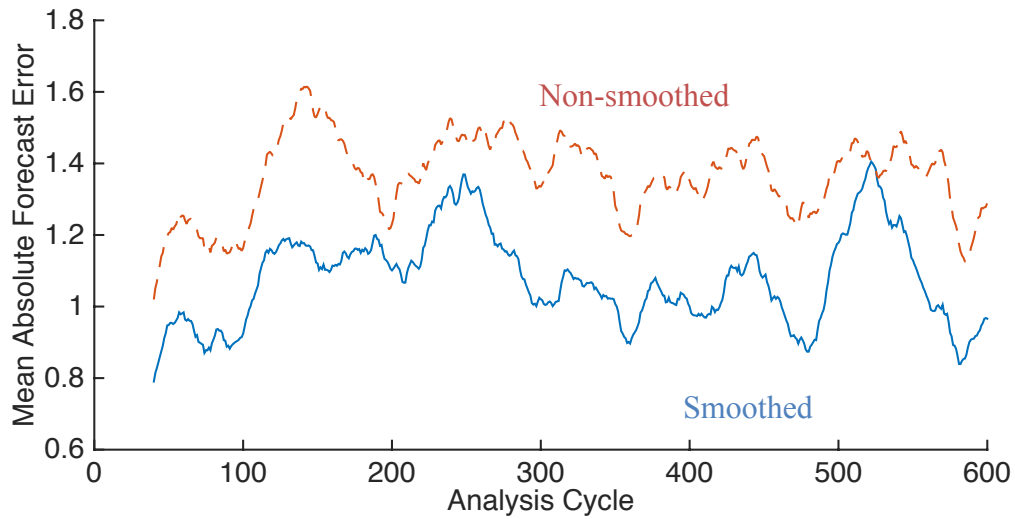
2 **Figure 6.** Elapsed time in seconds for (a) LETKF and (b) LPF minus LETKF.



3

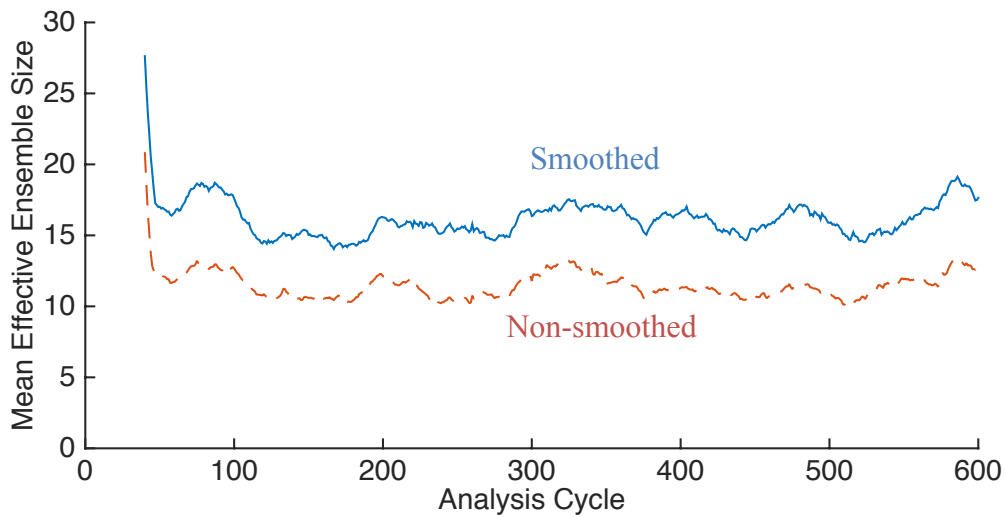
4 **Figure 7.** Analysis error for (a) LETKF and (b) LPF, using $l=80$ observations and $k=100$
 5 ensemble members. The observations used between (a) and (b) are identical. In (c), the
 6 number of observations for LPF is reduced to $l=20$, but improvement in accuracy versus
 7 LETKF remains.

8



1
 2 **Figure 8.** 40-cycle moving average of the mean absolute forecast error, comparing the Non-
 3 smoothed (dashed red) and Smoothed (solid blue) LPF analyses. The ensemble space
 4 smoothing improves the forecast accuracy.

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 7 **Figure 9.** Illustrating the impact of ensemble space smoothing on the effective ensemble size
 8 N_{eff} . Here we show the 40-cycle moving average of N_{eff} for the Smoothed (solid blue) and
 9 Non-smoothed (dashed red) cases.

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