

1 **Development of a hybrid variational-ensemble data**
2 **assimilation technique for observed lightning tested in**
3 **a mesoscale model**
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16 **Abstract**

17 Lightning measurements from the Geostationary Lightning Mapper (GLM) that
18 will be aboard the Geostationary Operational Environmental Satellite – R Series
19 will bring new information that can have the potential for improving the
20 initialization of numerical weather prediction models by assisting in the detection
21 of clouds and convection through data assimilation. In this study we focus on
22 investigating the utility of lightning observations in mesoscale and regional
23 applications suitable for current operational environments, in which convection
24 cannot be explicitly resolved. Therefore, we examine the impact of lightning
25 observations on storm environment. Preliminary steps in developing a lightning
26 data assimilation capability suitable for mesoscale modeling are presented in this
27 paper. World Wide Lightning Location Network (WWLLN) data was utilized as a
28 proxy for GLM measurements and was assimilated with the Maximum Likelihood
29 Ensemble Filter, interfaced with the Nonhydrostatic Mesoscale Model core of the
30 Weather Research and Forecasting system (WRF-NMM). In order to test this
31 methodology, regional data assimilation experiments were conducted. Results
32 indicate that lightning data assimilation had a positive impact on the following:
33 information content, influencing several dynamical variables in the model (e.g.
34 moisture, temperature, and winds), improving initial conditions during several

35 data assimilation cycles. However, the 6 h forecast after assimilation, did not
36 show a clear improvement in terms of root mean square errors.

37

38 **1 Introduction**

39 Thunderstorms are an important component of the climate system as they can
40 impact the atmospheric environment around them; they are capable of
41 redistributing moisture, heat, and wind patterns (Price, 2013). The assimilation of
42 lightning observations is a relatively new field. Several efforts to incorporate
43 lightning data into Numerical Weather Prediction (NWP) models have been made
44 recently (Alexander et al., 1999, Papadopoulos et al., 2005, Mansell et al., 2007,
45 Pessi and Bussinger, 2009, Fierro et al., 2012). In the vast majority of these
46 studies dynamical relaxation, or nudging techniques were applied. Even though
47 these studies highlighted the importance of utilizing lightning observations to
48 improve the representation of convection in models, they had less emphasis on
49 improving the environmental conditions.

50 Motivated by the initial success of nudging techniques in cloud-resolving
51 model applications, the objective of this study is to investigate if lightning
52 observations can be useful in mesoscale, regional, and global applications at a
53 coarse resolution, in which convection cannot be explicitly resolved. Therefore,
54 we would like to evaluate the impact of lightning observations on the environment
55 around storms, with potential implications to data assimilation, reanalysis, and
56 climate studies. As for any other observation, the information from lightning
57 observations can have impacts at several spatiotemporal scales. In the case of
58 lightning, one can assume that most of the information relates to cloud-resolving
59 processes. However, there should be also a fraction of lightning information that
60 can spread into larger scales (e.g., the storm environment). In this study we will
61 evaluate the large-scale component of information from lightning observations.

62 We anticipate that a myriad of applications can stem from monitoring
63 lightning activity. For instance, the lack of ground-based observations (e.g.
64 radiosondes, radars, etc.) over the open oceans can result in deficient
65 initialization of numerical weather and climate prediction models, especially if

66 weather systems that develop in these regions subsequently travel to continental
67 landmasses. Satellite radiances are an important source of observations over the
68 oceans. However, processing satellite observations requires considerably more
69 computational time due to the use of radiative transfer models, rather than just
70 processing lightning observations, which is computationally less intensive.
71 Therefore, the incorporation of this new type of data can provide useful
72 information for model initialization.

73 In addition, lightning may have a significant impact on the Earth's climate
74 by producing nitrogen oxides (NO_x) in the upper troposphere. NO_x is a precursor
75 of ozone, a major green house gas and pollutant (Price, 2013, Barthe et al.,
76 2010). The predicted concentrations of lightning- NO_x from NWP models coupled
77 with chemistry still contain large uncertainties. Incorporating geo-located lightning
78 data may assist these models in the simulation of convection, and consequently
79 NO_x production.

80 Lightning might be useful in future climate change monitoring studies due
81 to the interplay between lightning and atmospheric parameters, such as,
82 temperature, upper tropospheric water vapor, and cloud cover (Price, 2013).
83 Since lightning can be easily monitored through surface networks and satellite
84 platforms it can be a useful tool for tracking changes in important climate
85 parameters in the future (Price, 2009).

86 Satellite instruments have been launched in the past with the objective of
87 studying storm dynamics, cloud characteristics, annual and inter-annual
88 variability of thunderstorms, etc. (Adamo et al., 2009). In 1997, the Lightning
89 Imaging Sensor (LIS) was launched aboard de joint National Aeronautics and
90 Space Administration (NASA) and the Japan Aerospace Exploration Agency
91 (JAXA) Tropical Rainfall Measuring Mission (TRMM). This instrument can detect
92 lightning activity continuously at a horizontal resolution of 4 km over the tropics
93 (http://trmm.gsfc.nasa.gov/overview_dir/lis.html).

94 In the near future, mapping of lightning from geostationary orbit at cloud
95 scale resolution will be possible, thus complementing established surface
96 detection networks (Adamo et al., 2009, Finke, 2009). The launch of the

97 Geostationary Lightning Mapper (GLM) instrument that will be aboard the next
98 generation of the National Oceanic and Atmospheric Administration (NOAA)
99 geostationary satellites (i.e., GOES-R, [http://www.goes-
101 r.gov/spacesegment/glm.html](http://www.goes-
100 r.gov/spacesegment/glm.html)) will allow continuous day and night monitoring of
102 total lightning activity over the Americas and adjacent ocean regions up to 52
103 degrees north. One of the advantages over previous lightning mapping
104 instruments is that it will be able to monitor weather affecting the adjacent ocean
105 regions of the continental United States and not just the tropics. Some of the
106 mission objectives for the GLM instrument include: improvement in severe
107 thunderstorm lead times and false alarm reduction, advancements in the
108 initialization of NWP models through better identification of deep convection,
109 creation of lightning climatologies to track decadal changes in lightning activity,
among others (Adamo et al., 2008).

110 In this paper the possibility of assimilating lightning observations within a
111 hybrid variational-ensemble system in a mesoscale numerical weather prediction
112 model is explored, focusing on the typical resolution of operational weather
113 forecasting and climate models. The methodologies presented herein represent
114 an initial stage towards developing a comprehensive, multivariate, multi-scale,
115 multi-sensor data assimilation system that prepares for the assimilation of
116 lightning data along with other types of observations.

117 Eventually, this data assimilation technique will be tested in different
118 applications at various time and length scales. In the mean time, we intend to
119 investigate if the assimilation of lightning data can (1) add information content
120 into a mesoscale modeling system that can resolve a convective environment,
121 rather than explicit convection, (2) positively impact the dynamical variables of
122 the model, and (3) improve analysis and prediction. Note that a coarse
123 resolution is also typical of climate models, and thus assessing the utility of
124 lightning observations in data assimilation at these scales can be relevant for
125 climate studies as well. To our knowledge, lightning data have not been used in
126 operational weather prediction, in climate monitoring studies, or in reanalysis. By

127 assimilating lightning data in a coarse resolution model we are taking first steps
128 toward extending their use to weather and climate applications.

129 As a proof of concept case we chose the mesoscale convective system
130 that spawned numerous tornados over the southeastern United States on 27-28
131 April 2011. Lightning data from the World Wide Lightning Location Network
132 (WWLLN, <http://webflash.ess.washington.edu>) was used as a proxy to test the
133 potential impact of the assimilation of lightning flash rates measured by the GLM.
134 This data network has global coverage, including ocean regions. For North
135 America, this lightning detection network better approximates the coverage of the
136 upcoming GLM instrument compared to some surface networks that primarily
137 cover the continental United States.

138 The data assimilation system (DA) used in this study was the Maximum
139 Likelihood Ensemble Filter (MLEF – Zupanski, 2005; Zupanski et al., 2008),
140 which was interfaced with the non-hydrostatic core of the Weather and Research
141 Forecasting system (WRF-NMM - Janjić et al., 2010). The simplified
142 microphysics and low-resolution of the model defined the spatiotemporal scales
143 for data assimilation, as well as the options for the employed observation
144 operator. In this case, a 6-hour data assimilation window was chosen (± 3 hours
145 from a central time), in which the lightning observations were averaged at a
146 horizontal resolution of 10 km closely matching that of the innermost domain of
147 WRF-NMM.

148 This paper is organized in the following manner: the methodology for
149 using lightning observations is described in Sect. 2, details on the experimental
150 design are provided in Sect. 3, followed by results in Sect. 4, and finally a
151 summary and future work are presented in Sect. 5.

152

153 **2 Methodology for utilizing lightning observations**

154

155 **2.1 Data Assimilation System**

156

157 WRF-NMM was interfaced with MLEF, a hybrid ensemble-variational data
 158 assimilation method developed at Colorado State University. The solution of the
 159 analysis maximizes the likelihood of the posterior probability distribution,
 160 obtained by a minimization of a cost function that includes a general nonlinear
 161 observation operator. As in typical variational and ensemble data assimilation
 162 methods, a cost function is derived using a Gaussian probability density function
 163 framework. Like other ensemble data assimilation algorithms, MLEF produces an
 164 estimate of the analysis uncertainty (e.g., analysis error covariance). In addition
 165 to the common use of ensembles in calculations of the forecast error covariance,
 166 the ensembles in MLEF are exploited to efficiently calculate the Hessian
 167 preconditioning and the gradient of a cost function. The MLEF method is well
 168 suited for use with highly nonlinear observation operators, for a small additional
 169 computational cost of the minimization procedure. Relevant prognostic WRF-
 170 NMM variables were selected as control variables, as they can significantly
 171 impact the initial conditions, which can, in turn, influence the forecast. This
 172 selection includes the following variables: temperature (T), specific humidity (q),
 173 hydrostatic pressure depth (PD), the U and V components of the wind, and Cloud
 174 Water Mass (CWM – total cloud condensate in WRF-NMM) that combines all
 175 cloud hydrometeors into a total sum. The goal is to minimize the following cost
 176 function:

$$177 \quad J(x) = \frac{1}{2} [x - x^f]^T \mathbf{P}_f^{-1} [x - x^f] + \frac{1}{2} [y - h(x)]^T \mathbf{R}^{-1} [y - h(x)] \quad (1)$$

178
 179 where x represents the above defined control variables with a forecast error
 180 covariance \mathbf{P}_f , the index f denotes the forecast guess, y is the lightning flash rate
 181 observations with an error covariance \mathbf{R} , and h is the nonlinear lightning
 182 observation operator that maps the control variables to the lightning flash rate
 183 observations. The superscript T indicates the transpose of a matrix.

184

185 **2.2 Lightning flash rate observations**

186

187 Since the actual lightning measurements are lightning strikes, while the lightning
188 observation operator is commonly related to lightning flash rates, it was
189 necessary to transform lightning strikes into flash rates. In doing so, a subset
190 domain containing all lightning strikes was defined and subsequently partitioned
191 into a rectangular horizontal grid (different from the model grid), with a spacing of
192 0.1 degrees (~10 km) in order to be comparable with the horizontal grid spacing
193 of the smallest domain of our model configuration that will be discussed in Sect.
194 3.2. The choice of a regular grid that is not identical to the model grid is arbitrary.
195 In our case, it was motivated by a desire to keep the observation information
196 formally independent from the model, i.e. to not use any information about the
197 model when defining observations and observation errors. Lightning strikes
198 counted in each local area surrounding a grid point during a 6-hour time window
199 coinciding with the data assimilation interval were assigned to a particular grid
200 point, and then divided by a time interval to form lightning flash rates. Therefore,
201 the lightning flash rate observations are grid-point values that represent a
202 cumulative count of geo-located lightning strikes over the 6-hour assimilation
203 time window (± 3 hours from a central time), rather than the instantaneous
204 measurements. Note that the observed lightning flash rates were assumed to be
205 greater than zero, i.e., the observation grid points without any lightning strikes
206 were not included in the observations pool. Observations of zero lightning can be
207 important in pointing the location of misplaced convection events. However, it is
208 not clear how this information would impact convection events that are not
209 characterized by strong lightning. It is likely that additional information would be
210 needed in order to selectively define zero lightning observations. Even though,
211 this information is important, it needs further investigation. The non-negative
212 character of lightning observations introduces a skewness that points out to a
213 need for a non-Gaussian PDF in lightning data assimilation (e.g., Fletcher and
214 Zupanski, 2007; Lien et al. 2013). This issue will be examined in the future since
215 it can potentially improve the utility of lightning data.

216

217 **2.3 Lightning flash rate observation operator**

218

219 The lightning flash rate observation operator h (Eq. (1)) includes two operations:
220 a transformation (h_2) and an interpolation (h_1), i.e. $h = h_1 h_2$. In this study the
221 forward lightning transformation operator (h_2) was adopted by exploiting the
222 relationship between lightning and vertical velocity. This choice was influenced
223 by the properties of a bulk microphysics scheme used in the WRF-NMM model
224 (e.g., Ferrier, 2005), and by the coarse assimilation time window that effectively
225 restricts using the cloud-scale information about hydrometeors and their
226 interactions. A bi-linear interpolation technique was used to interpolate the guess
227 lightning flash rates to observation location (h_1).

228 As seen in previous studies, lightning is related to updrafts that support a
229 deep layer of super-cooled water droplets and a mixed phase region where
230 charge separation occurs (Black and Hallet, 1999). Based on Price and Rind
231 (1992), an empirical relationship between maximum updraft velocity (w_{\max}) and
232 lightning flash rate (f) given by:

$$233 \quad h_2 = f = c w_{\max}^{\beta} \quad (2)$$

234 was used, under the assumption that updrafts are positively correlated to cloud
235 top height. $c = 5 \times 10^{-6}$ and $\beta = 4.5$ are empirical parameters. β is a value derived
236 from satellite data climatologies for continental clouds as in Price and Rind
237 (1992). Both c and β are dimensionless.

238 The procedure to develop the lightning observation operator started with
239 an approximate calculation of vertical velocity from WRF-NMM, through the use
240 of a reduced version from the nonhydrostatic continuity equation

$$241 \quad w \approx \frac{1}{g} \left(\frac{\partial \Phi}{\partial t} + \mathbf{v} \cdot \nabla_{\sigma} \Phi + \dot{\sigma} \frac{\partial \Phi}{\partial t} \right) \quad (3)$$

242 where w is the vertical velocity, g is the gravity constant, Φ is the geopotential, \mathbf{v}
243 is the horizontal wind vector, and $\dot{\sigma}$ is the vertical velocity in a sigma coordinate
244 (Janjić, 2005). An approximation was required because vertical velocity is not a
245 predictive, but rather a diagnostic variable in WRF-NMM. After an approximate
246 value of vertical velocity was obtained, the maximum vertical velocity was

247 calculated for horizontal points according to the following procedure: values of
248 Cloud Water Mass (CWM - total cloud condensate in WRF-NMM) $CWM \geq 10^{-5}$
249 (kg kg^{-1}) were searched for at each model grid point and surrounding neighboring
250 points along all vertical model levels. We defined a 5×10 grid point area
251 (approximately a square domain in Arakawa E-grid staggering used in WRF-
252 NMM) surrounding the central point in order to introduce a smooth transition for
253 the calculation of w_{max} . This procedure was applied to avoid taking into account
254 values of w_{max} in regions without clouds. If the CWM threshold was reached, the
255 value of w_{max} was calculated at a grid point and surrounding points at all vertical
256 levels, otherwise w_{max} was set to zero. Once the value of w_{max} was calculated, it
257 was possible to calculate values of lightning flash rate from Eq. (2). Since the
258 calculation of w (e.g., Eq. (3)) and w_{max} includes prognostic model variables, all
259 control variables can impact lightning flash rates.

260 Since both a new observation type (lightning flash rate) and an untested
261 observation operator (Eq. (2)) were introduced into the data assimilation system,
262 statistics of innovation vectors (observation minus guess) of lightning flash rates
263 needed to be examined first. Figure 1 shows the statistics of the normalized
264 innovation vectors $R^{-1/2}[y-h(x^f)]$ at several observation times. A skewed
265 histogram of the Probability Distribution Function (PDF) innovation vectors (left)
266 can be readily seen, implicitly indicating that the observed values of lightning
267 flash rate were considerably larger than the guess. Therefore, it was necessary
268 to perform a correction. An option could have been to increase the value of
269 parameter c in Eq. (2) to reduce the skewness. However, trial experiments
270 indicated a large uncertainty of the parameter c from one observation time to
271 another, in occasions ranging over two orders of magnitude. In order to deal with
272 this error of the observation operator (Eq. (2)), an adjustable multiplicative
273 correction parameter ($\alpha > 0$) was included so that h_2 would become αh_2 . At each
274 observation time an optimal parameter α_{opt} was estimated by minimizing the
275 following cost function:

276
277

278

$$J(\alpha) = \frac{1}{2} [\log(\alpha) - \log(\alpha_0)]^T \mathbf{W}^{-1} [\log(\alpha) - \log(\alpha_0)]$$

$$+ \frac{1}{2} [\log(y) - \log(\alpha h(x^f))]^T \mathbf{R}_L^{-1} [\log(y) - \log(\alpha h(x^f))] \quad (4)$$

279 where \mathbf{R}_L is the observation error covariance associated with a logarithmic
 280 transformation, α_0 is a guess value, and \mathbf{W} is the uncertainty matrix of the guess
 281 value. The choice of a logarithmic transformation was influenced by the fact that
 282 lightning flash rate is strictly positive definite and that such procedure could better
 283 deal with the large uncertainty of the parameter α . As shown in the appendix
 284 (Sect. 7), the solution of α_{opt} , which minimizes the cost function, i.e., Eq. (4), is
 285 given by:

286

$$\alpha_{opt} = \exp \left[\frac{\frac{1}{N_{obs}} \sum_{i=1}^{N_{obs}} \log \left(\frac{y}{h(x)} \right)_i}{1 + \frac{r_0}{w_0}} \right]. \quad (5)$$

287 where N_{obs} is the number of observations, $diag(\mathbf{W})=w_0$ and $diag(\mathbf{R}_L)=r_0$.
 288 Therefore the lightning observation transformation operator (Eq. (2)) was
 289 substituted by

290

$$h_2 = f = \alpha_{opt} c w_{max}^\beta. \quad (6)$$

291 The observation operator transformation (e.g., Eq. (6)) is defined over a 2-
 292 dimensional horizontal domain only since flash rate f is a horizontal field (e.g.,
 293 number of hits per area and time). This requires w_{max} to be 2-dimensional as well.
 294 Therefore, w_{max} is defined for each horizontal grid point, as the maximum value of
 295 vertical velocity (w) over all vertical levels. The flow diagram of the data
 296 assimilation system and the lightning observation operator are illustrated in Fig.
 297 2.

298

299 **2.4 Information content of lightning observations**

300

301 In general terms, the impact of observations can be quantified using an
 302 uncertainty reduction after data assimilation. Since entropy measures the

303 uncertainty, one can use the formalism of Shannon information theory (Shannon
 304 and Weaver 1949) to define information content of observations as an entropy
 305 difference before and after data assimilation. As shown in Rodgers (2000), the
 306 entropy is considerably simplified with a Gaussian probability assumption and
 307 information content can be conveniently expressed in terms of degrees of
 308 freedom for signal (d_s),

$$309 \quad d_s = \text{trace}[\mathbf{I} - \mathbf{P}_a \mathbf{P}_f^{-1}], \quad (7)$$

310 where *trace* is the trace function, \mathbf{I} is the identity matrix, and \mathbf{P}_a is the analysis
 311 error covariance. This can be further reduced in terms of the eigenvalues of the
 312 observation information matrix, given by:

$$313 \quad \mathbf{P}_f^{T/2} \mathbf{H}^T \mathbf{R}^{-1} \mathbf{H} \mathbf{P}_f^{1/2} = \mathbf{U} \mathbf{\Lambda} \mathbf{U}^T \quad (8)$$

314 where $\mathbf{\Lambda}$ and \mathbf{U} are the eigenvalues and eigenvectors matrices, respectively, and
 315 \mathbf{H} is the Jacobian of the observation operator. The degrees of freedom for signal
 316 are then

$$317 \quad d_s = \sum_i \frac{\lambda_i^2}{1 + \lambda_i^2} \quad (9)$$

318 where λ_i are the eigenvalues. Zupanski et al. (2007) showed that this formula
 319 could also be useful in reduced-rank, ensemble space calculations, in which the
 320 summation is performed over the number of ensemble members. Since an
 321 eigenvalue decomposition of the observation information matrix is a component
 322 of the MLEF algorithm, additional cost of calculating d_s is minimal. By calculating
 323 the degrees of freedom for signal we can quantify the impact of the lightning
 324 observations in terms of an uncertainty reduction. Note that Eq. (9) has non-
 325 negative values between 0 and N_{ens} , depending on the structure of the
 326 observation information matrix. If there is a negligible impact of lightning
 327 observations the number of degrees of freedom for signal will be close to zero,
 328 i.e. much smaller than the number of ensemble members.

329 **3 Experimental Design**

330

331 **3.1 General synoptic description of the case study**

332

333 As a proof of concept case for regional lightning data assimilation over a
334 continental area we selected the severe weather event that occurred on 27-28
335 April 2011, where an estimated 292 tornadoes hit the southeastern, mid-west
336 and northeast United States, according to the Storm Report Center (Fig. 3, top
337 panel). A figure of 500 hPa heights, with color contours of wind speed and
338 surface observations from the Forecast Systems Laboratory (Fig. 3, bottom
339 panel) shows that atmospheric conditions created a perfect scenario for severe
340 weather development. An upper-level low centered on Minnesota along with the
341 advance of a deep trough and its associated jet streak (wind speed exceeding
342 41.15 m sec^{-1}) aloft led to rapid atmospheric destabilization in the afternoon of 27
343 April. Surface moist-warm flow arrived from the Gulf of Mexico, with dew points
344 exceeding $21 \text{ }^\circ\text{C}$ and wind gusts over 7.72 m sec^{-1} at the Alabama coast. An
345 upper level disturbance sparked a broad area of showers and thunderstorms as it
346 moved across the frontal boundary on the previous evening. The eastern edge
347 of this line of showers and storms continued to move eastward, in concert with
348 the upper-level disturbance, reaching the northwest Alabama border around
349 0700 UTC on the 27th. Meanwhile, surface winds backed to the south-southeast
350 as the disturbance moved into the area, while winds at the 850 hPa level (around
351 1,500 m) increased to $26\text{-}28 \text{ m sec}^{-1}$ and became more southerly. The
352 combination of high low-level moisture and increasing shear provided the setup
353 for damaging winds, large hail and brief tornadoes. This line experienced further
354 intensification as it moved into northwest Alabama, especially after 0900 UTC.
355 This line of severe storms pushed into northwestern Alabama prompting a
356 tornado watch for all of northern Alabama and portions of southern middle
357 Tennessee until 1400 UTC. A deep layer shear and moisture increased
358 dramatically later in the afternoon and evening of the 27th ahead of the strong
359 cold front. This combination of strong instability and high shear continued through

360 the evening hours ahead of the cold front before it pushed east of the area into
361 Georgia. This produced the last and most violent round of severe weather, which
362 began around 2030 UTC for northern Alabama as supercells began to line up to
363 the southwest of the area. During the early afternoon hours, the potential for
364 destructive tornadoes was highlighted by the Storm Prediction Center's upgrade
365 to a rare High Risk for severe weather around 1300 UTC. This prompted a
366 Particularly Dangerous Situation (PDS) tornado watch, which was issued for
367 northern Alabama and portions of southern middle Tennessee at 1945 UTC. The
368 potential really ramped up from noon through 2100 UTC. During this period,
369 much of Alabama experienced numerous supercell thunderstorms producing
370 strong to violent tornadoes, including five EF-4 tornadoes and one EF.5 in the
371 Huntsville Forecast Area (NOAA Service Assessment, Hayes, 2011,
372 http://www.srh.noaa.gov/hun/?n=hunsur2011-04-27_setup).

373

374 **3.2 Model and domain configuration**

375

376 The WRF-NMM version 3 model from the Developmental Testbed Center
377 [<http://www.dtcenter.org>] was employed in this study. WRF-NMM was developed
378 by the NOAA/National Centers for Environmental Prediction (NCEP) (Janjić et
379 al., 2010). For simplicity, only some physics and dynamics choices are
380 mentioned. The microphysics option was Ferrier (Ferrier, 2005), which includes
381 prognostic mixed-phase processes. The longwave and shortwave radiation
382 options were the Geophysical Fluid Dynamics Laboratory (GFDL) schemes. The
383 GFDL longwave radiation scheme includes the transmission and absorption of
384 carbon dioxide, ozone, and water vapor in multiple spectral bands. Likewise, in
385 the GFDL shortwave scheme, ozone and water vapor are the main absorbers.
386 Both schemes include cloud microphysical effects (Falkovich et al., 2005). The
387 planetary boundary layer option was the Mellor-Yamada-Janjinc (Janjić, 1994).
388 The land surface option was the NOAH Land-Surface model (Ek et al., 2003)
389 with soil temperature and moisture in four layers, fractional snow cover and
390 frozen soil physics. For the cumulus parameterization, Betts-Miller-Janjić was

391 selected. This scheme adjusts deep shallow convection with a relaxation towards
392 variable humidity and temperature profiles (BMJ-Janjić 1994, 2000).

393 The WRF-NMM simulations in this study were configured with two
394 domains. Domain 1 (D01) had a horizontal grid spacing of 27 km and a size of
395 1350 by 2592 km² (50 x 96 grid points). This domain covered parts of the mid-
396 west, the Gulf of Mexico, the Atlantic Ocean, and the eastern United States.
397 Domain 2 (D02), centered on Alabama, had a horizontal grid spacing of 9 km and
398 a size of 540 by 1170 km² (60 x 130 grid points) (Fig. 4). Both domains had a
399 vertical extent of 27 vertical levels.

400

401 **3.3 Data sets and data assimilation system setup**

402

403 The ensemble boundary conditions are obtained from the NCEP Global Forecast
404 System (GFS) using the WRF preprocessing system (WPS). With the exception
405 of the initial ensemble preparation (i.e. cycle0 in our terminology), the initial
406 conditions for the ensemble members are obtained through the MLEF algorithm
407 by adding the analysis square root error covariance columns to the analysis.
408 Further information about the MLEF methodology can be found in Zupanski
409 (2005) and Zupanski et al. (2008). The localization setting for the ensemble-
410 based covariance includes a de-correlation length of 90 km. The data
411 assimilation period starts at 1800 UTC 26 April 2011, ending on 1200 UTC 28
412 April 2011. Note that there is no data assimilation at the initial time.

413 In the present study, WWLLN data were assimilated. The WWLLN is an
414 experimental lightning detection network that provides the location of cloud-to-
415 ground (CG) and some intra-cloud lightning (IC) strikes in real-time, it has a
416 global coverage with 10 km location accuracy and flash detection accuracy
417 greater than 50% (Lay, 2004). WWLLN is for the most part; a time average of
418 geo-located CG lightning flashes that cannot address the cloud-resolving
419 characteristics of lightning. Nonetheless, for the purposes of evaluating the
420 impact of lightning observations on the storm environment, making a distinction
421 between CG and IC lightning is beyond the scope of this study. The ensemble

422 size was set to 32 in order to match the number of processors per node, with a
423 data assimilation interval of 6 hours to match the frequency of the Global
424 Forecast System (GFS) input files. The 6-hourly averaged lightning flash rates
425 (± 3 hours) were assimilated at each central time t_n ($n > 0$). An initial 6-hour
426 forecast was obtained at cycle0 from WRF-NMM with the GFS files (from t_{n-3h} to
427 t_{n+3h}) and it was used as a first guess to obtain the analysis solution for the next
428 cycle. The background state x^f , or prior, is an estimate of the most likely
429 dynamical state; it is a deterministic forecast from the previous assimilation cycle.
430 The analysis solution was obtained as a maximum likelihood estimate from the
431 assimilation of observations at the central time t_n (Zupanski, 2005). These steps
432 were repeated during each cycling period. Figure 5 shows the data assimilation
433 timeline. The observational error was assumed to be $0.10 \text{ hits km}^{-2} \text{ h}^{-1}$.

434

435 **3.4 Description of the experiments**

436

437 Three simulations were performed to assess the impact of the assimilation of
438 lightning flash rates into a mesoscale NWP:

439 1. The first experiment was a single observation test (1-OBS), performed to
440 evaluate the impact of assimilating lightning flash rates at a single
441 WWLLN location (34.5°N , 89°W) on the analysis increment (analysis
442 minus background) of a subset of the control variables (q , T , U , and V)
443 mentioned in Sect. 2.1 and to implicitly illustrate the complex structure of
444 the flow-dependent forecast error covariance. The difference between the
445 initial observation and the guess was assumed to be one standard
446 deviation of the observation error covariance \mathbf{R} , i.e., $y = x^f + \sigma_R$ where
447 $\sigma_R = 1$.

448 2. The second experiment was a control run, without the assimilation of
449 lightning data, referred to as no-data-assimilation (NODA). Note, however,
450 that lightning observations were still present in the simulation in order to
451 define the optimal regression parameter α_{opt} .

452 3. In addition to the two simulations mentioned before, an experiment that
453 included the assimilation of WWLLN lightning data (LIGHT) was
454 performed. LIGHT had the same set-up as the NODA simulation; the only
455 difference was the assimilation of lightning flash rates.

456

457 **4 Results**

458

459 In the following sections, we present an evaluation of the impact of the
460 assimilation of lightning data for the 27-28 April 2011 severe weather event
461 focusing on domain D02 (9km resolution). First, results of the (1-OBS)
462 experiment are shown, followed by an evaluation of the time-flow-dependent
463 forecast error covariance through the use of degrees of freedom for signal to
464 quantify the information added to the system by the assimilation of the lightning
465 observations. Then an evaluation of several synoptic fields from the LIGHT
466 simulation and validation of the DA system through comparisons with some
467 observations are presented. Thereafter, an assessment between the LIGHT and
468 NODA simulations through the calculation of Root Mean Square (RMS) errors of
469 the lightning observations is shown

470

471 **4.1 1-OBS experiment**

472

473 The difference between the analysis and the 6-hour forecast (background) was
474 evaluated. Figure 6a shows the 700 hPa analysis increments of specific humidity
475 (q) at 1800 UTC 27April 2011, or cycle 3 in the data assimilation period. This
476 time was chosen since tornados started developing over northern Alabama just a
477 couple of hours before. The black dot indicates the location of the single
478 observation being assimilated (34.5°N, 89°W). A clear dipole of positive and
479 negative analysis increments in q , with a magnitude of $\pm 4 \times 10^{-5} \text{ kg kg}^{-1}$, is
480 observed at opposite sides of the location of the single observation. The analysis
481 increment of temperature (T) at 700 hPa (Fig. 6b) shows regions of positive and
482 negative analysis increments, with a magnitude of $\pm 4 \times 10^{-2} \text{ degrees K}$, over the

483 same regions as q , but with opposite sign. The plot of wind speed at 700 hPa
484 (Fig. 6c) shows a positive analysis increment of $2.7 \times 10^{-1} \text{ m sec}^{-1}$ with maximum
485 values coinciding with the region of positive potential temperature increment. The
486 spatial extension of the impact of assimilating a single lightning strike on some of
487 the dynamical variables of the model in D02 (9 km resolution) was: (i) on specific
488 humidity the impact extends to approximately 12 grid points ($\sim 110 \text{ km}$), (ii) for
489 temperature to 20 grid points ($\sim 180 \text{ km}$), and (iii) for wind approximately 30 grid
490 points ($\sim 270 \text{ km}$).

491 The former Fig. (6a, 6b, and 6c) indicates that the assimilation of lightning
492 at a single location impacted the atmospheric environment at surrounding grid
493 points. The magnitude of the analysis increments indicates non-negligible
494 adjustments on dynamical variables of the mesoscale model. Most importantly, it
495 can be noted that the hybrid DA system was able to spatially spread the
496 information of a single lightning observation and influence the initial conditions of
497 specific humidity, temperature, the U and V components of the wind and other
498 control variable elements. These results are a manifestation of the complex
499 structure of the ensemble forecast error covariance matrix. This is important
500 since it indicates that the information from lightning observations can impact the
501 initial conditions and eventually the forecast of coarse resolution models.

502

503 **4.2 Evaluation of information content of the lightning observations**

504

505 In these experiments, the degrees of freedom for signal were computed in
506 ensemble subspace following Zupanski et al. (2007). The top-three plots in Fig. 7
507 show degrees of freedom for signal during three assimilation cycles (1, 2 and 3,
508 as an example) and observed GOES-IR and lightning flash rates at matching
509 times (bottom-three plots). The areas of highest density of WWLLN lightning
510 observations are in agreement with information content, implying that the time-
511 flow-dependent forecast error covariance had a direct relationship to the
512 observations throughout the assimilation period. Maximum values of degrees of
513 freedom for signal of 12, 22, and 10 for cycles 1, 3, and 5, respectively can be

514 observed in Fig. 7. These values indicate that the benefit of the observations is
515 important, otherwise these values would be close to zero, i.e. much smaller than
516 the number of ensemble members, 32 in this case. On the other hand, if the
517 former values were to approach the number of ensemble members, this would be
518 an indicator of the introduction of noise to the DA system by the observations and
519 their possible benefit would be nullified. Note however, that the agreement in
520 cycle 3 was not very good. It is possible that ensemble perturbations were not
521 large enough over northeastern Alabama where another maximum was missing.
522 This lack of agreement can arise from the use of a reduced rank ensemble
523 approach and consequently not having enough spread in the ensembles.
524 However, the agreement improved in subsequent cycles (e.g., shown for cycle
525 5).

526

527 **4.3 Impacts on the environment during the severe weather event**

528

529 The following results correspond to 0000 UTC 28 April 2011, cycle 5 in the data
530 assimilation time line, at the time when an EF4 tornado affected Tuscaloosa and
531 Birmingham, Alabama. Fields of wind, absolute vorticity and Convective
532 Available Potential Energy (CAPE) from both experiments (LIGHT and NODA)
533 portray a distinctive scenario of an environment favorable for the strengthening of
534 deep convection, but with some differences. Figure 8a shows background
535 (forecast) winds at 850 hPa for the NODA experiment. Figure 8b shows
536 background winds at 850 hPa for the LIGHT experiment. A core of increased
537 wind speed over northern Alabama can be observed in both plots. However, the
538 core of maximum wind speed has a larger spatial coverage in the LIGHT
539 experiment and based on computed differences, stronger winds with a
540 magnitude, in the order of 4 to 6 m sec⁻¹ were found in the LIGHT experiment.
541 Note that this region is co-located with an area of high density of WWLLN
542 lightning observations (Fig. 8c). Figures 9a and 9b correspond to the analysis
543 increment of the 850 hPa winds and absolute vorticity, respectively. Regions of
544 positive increments are found near the left-hand side in both plots (4 to 6 m sec⁻¹

545 in wind speed and $4 \times 10^{-4} \text{ sec}^{-1}$ in vorticity). Almost no analysis increments can
546 be found in the region where the densest lightning observations are located
547 (Alabama). Among possible reasons, we can mention the following: (i) the largest
548 forecast uncertainty (i.e. ensemble perturbations) typically occurs in the areas of
549 strongest dynamical instability, in this case, in the region where a dry line was
550 present over the states of Louisiana, Mississippi, Arkansas, and Missouri. Even
551 though, the dry line may not be characterized by the strongest lightning activity,
552 there were still some isolated lightning observations present over the domain as
553 seen in Fig. 8c, (ii) alternatively, it may be a consequence of using an ensemble-
554 based forecast error covariance that was not able to produce sufficient
555 uncertainty in all relevant areas.

556 Similarly, by analyzing CAPE at the forecast step for both experiments
557 (Fig. 10a,b), a region of high CAPE gradient is observed on the left hand side of
558 the domain, indicating the presence of a well-defined dry line. However, no
559 significant differences were found between both experiments for this particular
560 assimilation cycle (cycle 5). One possible reason is that there were no lightning
561 observations present at the core where the strongest CAPE was observed.
562 Therefore, lightning was not able to significantly impact CAPE. Further
563 investigation is required to see if the same behavior occurs for other cycles and
564 case studies.

565 Forecast CAPE was validated by comparing the model output with
566 observations from the Storm Prediction Center's Surface Mesoanalysis at 40 km
567 resolution. Figure 10c shows observed CAPE. A well-defined dry-line can be
568 readily seen in the plot of background CAPE (Fig. 10a,b), which coincides with
569 the location of a strong CAPE gradient on the observations (Fig. 10c). The
570 formation of a dry line can often be a precursor for severe thunderstorm
571 formation with tornadogenesis potential (Grazulis, 2001). Note however, that the
572 model missed the location of the core of Maximum CAPE ($\sim 3500 \text{ J kg}^{-1}$) by one
573 degree, latitude and longitude. The observed maximum CAPE was located over
574 the ocean, just off the Mississippi coast, while in the model output; the same core
575 was placed at the southern Mississippi-Louisiana border. Nonetheless, by

576 assimilating lightning flash rates, the analysis increased, thus increasing the
577 magnitude of winds and absolute vorticity at 850 hPa. The analysis increment of
578 wind, suggests that absolute vorticity was advected into the region of strong
579 CAPE gradient (dry-line).

580

581 **4.4 Statistics: analysis and forecast Root Mean Square (RMS) errors with** 582 **respect to the lightning observations (LIGHT vs. NODA)**

583

584 A qualitative comparison of atmospheric fields between the data assimilation
585 (LIGHT) and the control (NODA) experiments with observations may lead to
586 subjective conclusions on determining which experiment outperformed the other.
587 Statistical evaluations on the other hand, can provide useful diagnostics when
588 morphological differences are not obvious.

589 Analysis and forecast RMS errors with respect to the lightning
590 observations were calculated from a domain containing the observed lightning
591 flash rates at 10 km resolution during the 6-hour assimilation time window, as
592 described in Sect. 2.2. From Fig. 11a, the LIGHT experiment achieves a better fit
593 in the analysis compared to the NODA experiment, but not for cycle 6. A possible
594 reason could be that the system was exiting the model domain at that time. Since
595 the strongest convection and cold front moved away from the domain, there was
596 no significant lightning activity over the region. Consequently, the number of
597 lightning observations available for data assimilation significantly decreased and
598 the impact of lightning data assimilation was reduced. The analysis result is not
599 well retained in the forecast (Fig. 11b). This issue definitely requires further
600 investigation. A possible reason may be that there are no other types of
601 observations being assimilated, such as conventional and satellite observations
602 that would additionally constrain the analysis and eventually create dynamical
603 balance, further improving the analysis and consequently the forecast. Note that
604 lightning is just an additional type of observation. All available observations have
605 to be in agreement with each other at the same location. Therefore, in regions
606 where lightning observations are not in agreement with other types of

607 observations, the data assimilation algorithm will create the optimal observation
608 impact based on uncertainty of all observations in the region. In areas where
609 lightning observations are not available other measurements should help.

610

611 **5 Summary and future work**

612

613 In this study, the preliminary development and assessment of a methodology for
614 the assimilation of lightning observations through hybrid variational-ensemble
615 methods is presented. The aim of the study was to evaluate if lightning data
616 assimilation can be useful in mesoscale, regional, and global applications at a
617 coarse resolution in which convection cannot be explicitly resolved. The MLEF
618 system interfaced with WRF-NMM was utilized to investigate the impacts of
619 lightning data assimilation on a mesoscale NWP model. As a proof of concept,
620 this methodology was tested for the 27-28 April 2011 severe weather event in the
621 southeastern United States. Results indicate that lightning was capable of
622 spreading new information into the WRF-NMM model. Analysis increments of
623 750 hPa specific humidity, temperature, and winds indicate that the assimilation
624 of lightning flash rates could impact the initial conditions of a subset of model
625 variables (q , T , U and V) leading to dynamical balance as shown by the output
626 from the 1-OBS test. The information content of lightning data was quantified
627 through the calculation of degrees of freedom for signal. Regions of high density
628 of observed lightning flash rates were in agreement with information content
629 theory indicating that the time-flow-dependent forecast error covariance was
630 directly related to observations during the assimilation period.

631 Evaluation of some atmospheric fields from the LIGHT experiment
632 indicated that the assimilation of lightning data influenced winds, absolute
633 vorticity and CAPE. A core of increased background wind speed at 850 hPa
634 coincides with the location of the region of high density in lightning observations
635 for the same assimilation cycle, indicating that the assimilation of lightning data
636 had an impact on the increase of wind speed. Analysis increments of the 850
637 hPa wind, absolute vorticity and background CAPE indicated that vorticity was

638 advected into the region of strong CAPE gradient where a dry-line formed. All
639 these changes suggest the development of an environment favorable for
640 strengthening of deep convection.

641 Analyses and forecast RMS errors with respect to the lightning
642 observations from the LIGHT and NODA experiments indicated that LIGHT
643 achieved a better fit at the analysis step compared to the NODA experiment.
644 However, the 6-hour forecast after assimilation did not show any clear
645 improvements in terms of the RMS errors. This requires further investigation.

646 The methodology presented in this study represents an initial step towards
647 developing a comprehensive multivariate, multi-scale, multi-sensor operational
648 data assimilation system that prepares for the assimilation of lightning along with
649 different types of operational observations and for multiple applications. As a first
650 step, we intended to verify if the data assimilation techniques described here
651 could be accomplished and that lightning data could add information content to a
652 modeling system with a coarse resolution similar to the ones used in operations.
653 Further studies are planned where this methodology will be tested for different
654 applications (e.g. different case studies, different models, and choice of
655 observation operators). Operational conventional and satellite observations will
656 be assimilated alongside lightning flash rates to further constrain the fit in the
657 analysis.

658

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660

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672 the authors and should not be construed as an official National Oceanic and
673 Atmospheric Administration (NOAA) or U.S. Government position, policy, or
674 decision.

675 **7 Appendix. Lightning flash rate observation operator correction: Weak**
 676 **Constrain**

677 Assume a multiplicative correction to the observation operator (i.e. correction in
 678 magnitude, not in the direction of the vector)

679

$$680 \quad h(x) \rightarrow \alpha h(x), \quad (\text{A1})$$

681

682 where $\alpha > 0$ is the unknown multiplication parameter.

683

684 Consider a logarithmic function of vectors since all vectors (i.e. y and $h(x)$) are
 685 positive definite and define a cost function with the adjustable parameter α :

$$686 \quad J(\alpha) = \frac{1}{2} [\log(\alpha) - \log(\alpha_0)]^T \mathbf{W}^{-1} [\log(\alpha) - \log(\alpha_0)] \quad (\text{A2})$$

$$+ \frac{1}{2} [\log(y) - \log(\alpha h(x^f))]^T \mathbf{R}_L^{-1} [\log(y) - \log(\alpha h(x^f))]$$

687 where \mathbf{R}_L is the observation error covariance associated with a logarithmic
 688 transformation, α_0 is a guess value, and \mathbf{W} is the uncertainty matrix of the guess
 689 value. The optimal parameter $\alpha_{opt} > 0$ that minimizes the cost function (A2) is
 690 searched for. Following a standard procedure of function minimization to solve:

691

$$692 \quad \left(\frac{\partial J(\alpha)}{\partial \alpha} \right)_{\alpha_{opt}} = 0 \quad . \quad (\text{A3})$$

693

694 Note that in order to differentiate with respect to α it may be more convenient to
 695 redefine the cost function (A2) in the following manner:

696

$$697 \quad J(\alpha) = \frac{1}{2} [\log(\alpha) - \log(\alpha_0)]^T \mathbf{W}^{-1} [\log(\alpha) - \log(\alpha_0)] \quad (\text{A4})$$

$$+ \frac{1}{2} \left[\log\left(\frac{y}{h(x)}\right) - \log(\alpha) \right]^T \mathbf{R}_L^{-1} \left[\log\left(\frac{y}{h(x)}\right) - \log(\alpha) \right]$$

698

699

700 The Jacobian of (A4) is

701

$$702 \quad \frac{\partial J(\alpha)}{\partial \alpha} = \frac{1}{\alpha} [1]^T W^{-1} [\log(\alpha) - \log(\alpha_0)] - \frac{1}{\alpha} [1]^T R_L^{-1} \left[\log\left(\frac{y}{h(x)}\right) - \log(\alpha) \right], \quad (\text{A5})$$

703

704 where $[1]$ is a vector with all components equal to one. After employing (A3)

705

$$706 \quad \frac{1}{\alpha} \left\{ (\log(\alpha)) [1]^T W^{-1} [1] + \log \alpha [1]^T R_L^{-1} [1] - [1]^T R_L^{-1} \left[\log\left(\frac{y}{h(x)}\right) \right] - \log(\alpha_0) [1]^T W^{-1} [1] \right\} = 0. \quad (\text{A6})$$

707

708 After multiplying (A6) by α (where $\alpha > 0$) (A6) can be rewritten as

709

$$710 \quad (\log(\alpha)) [1]^T [R_L^{-1} + W^{-1}] [1] - [1]^T R_L^{-1} \left[\log\left(\frac{y}{h(x)}\right) \right] - \log(\alpha_0) [1]^T W^{-1} [1] = 0. \quad (\text{A7})$$

711

712 From (A7):

713

$$714 \quad \log(\alpha) = \frac{\left([1]^T R_L^{-1} \left[\log\left(\frac{y}{h(x)}\right) \right] + \log(\alpha_0) [1]^T W^{-1} [1] \right)}{[1]^T [R_L^{-1} + W^{-1}] [1]}. \quad (\text{A8})$$

715

716 Finally, the optimal multiplicative parameter is given by:

717

$$718 \quad \alpha_{opt} = \exp \left[\frac{\left[[1]^T R_L^{-1} \left[\log\left(\frac{y}{h(x)}\right) \right] + \log(\alpha_0) [1]^T W^{-1} [1] \right]}{[1]^T [R_L^{-1} + W^{-1}] [1]} \right] \quad (\text{A9})$$

719

720 After employing a common assumption that the uncertainty matrix \mathbf{W} and the
 721 observation error matrix \mathbf{R}_L are diagonal, with $diag(\mathbf{W}) = w_0$ and $diag(\mathbf{R}_L) = r_0$,
 722 respectively,

723

$$724 \quad [1]^T W^{-1} [1] = N_{obs} w_0^{-1} \quad (A10)$$

725

$$726 \quad [1]^T [R_L^{-1} + W^{-1}] [1] = N_{obs} (r_0^{-1} + w_0^{-1}) \quad (A11)$$

727

$$728 \quad [1]^T R_L^{-1} \left[\log \left(\frac{y}{h(x)} \right) \right] = r_0^{-1} \sum_{i=1}^{N_{obs}} \left[\log \left(\frac{y}{h(x)} \right) \right]_i. \quad (A12)$$

729

730 where N_{obs} is the number of observations. By substituting (A10), (A11), and (A12)
 731 in (A9) gives:

732

$$733 \quad \alpha_{opt} = \exp \left\{ \frac{\frac{1}{N_{obs}} \sum_{i=1}^{N_{obs}} \log \left(\frac{y}{h(x)} \right)_i + \left(\frac{w_0}{r_0} \right)^{-1} \log(\alpha_0)}{1 + \left(\frac{w_0}{r_0} \right)^{-1}} \right\}. \quad (A13)$$

734

735 Without additional knowledge, a typical guess value is $\alpha_0 = 1$, which further
 736 simplifies the solution (A13) to

737

$$738 \quad \alpha_{opt} = \exp \left[\frac{\frac{1}{N_{obs}} \sum_{i=1}^{N_{obs}} \log \left(\frac{y}{h(x)} \right)_i}{1 + \frac{r_0}{w_0}} \right]. \quad (A14)$$

739

740

741 The above expression can be easily calculated in the observation operator and
 742 provide an adjustable correction factor.

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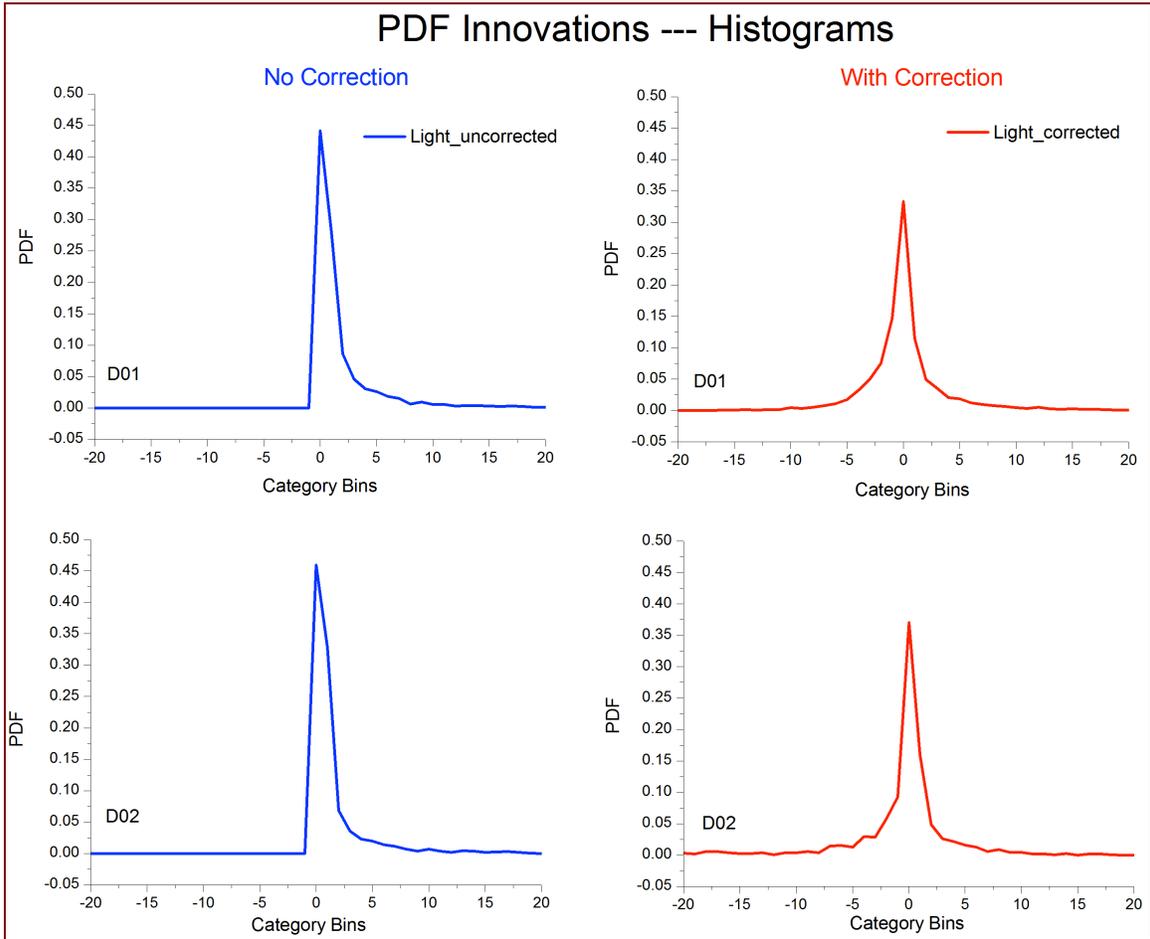
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858 **Figures**

859



860

861

862 **Fig. 1.** Statistics of normalized innovation vectors $R^{-1/2} [y - h(x^f)]$, or PDF
863 innovations for cycles 1-5 for both domains (D01 and D02) before (left-blue) and
864 after (right-red) correction. The skewed histograms on the left implicitly indicate
865 that the values of observed lightning flash rate are considerably larger than the
866 guess, a situation that required a correction.

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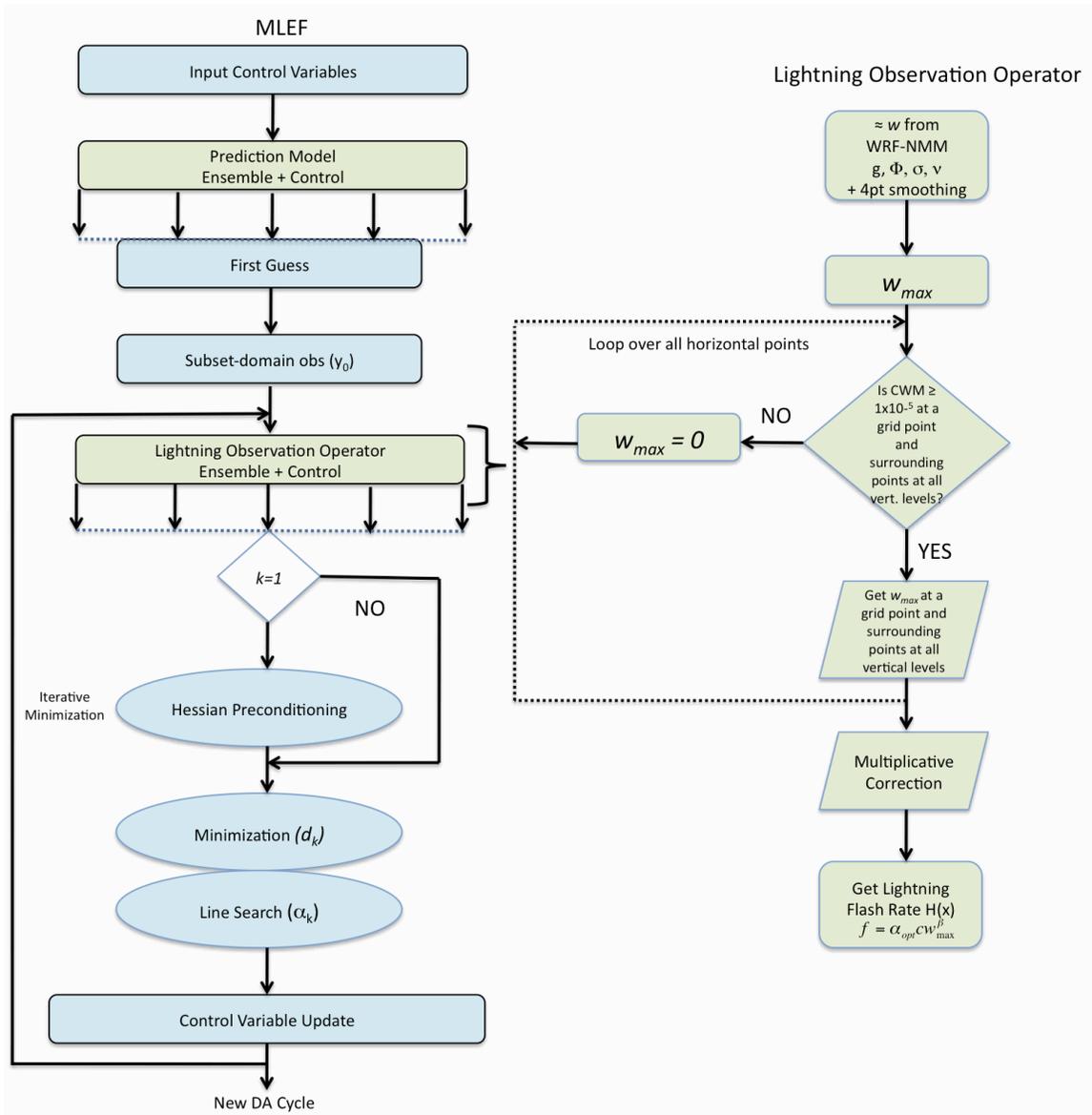
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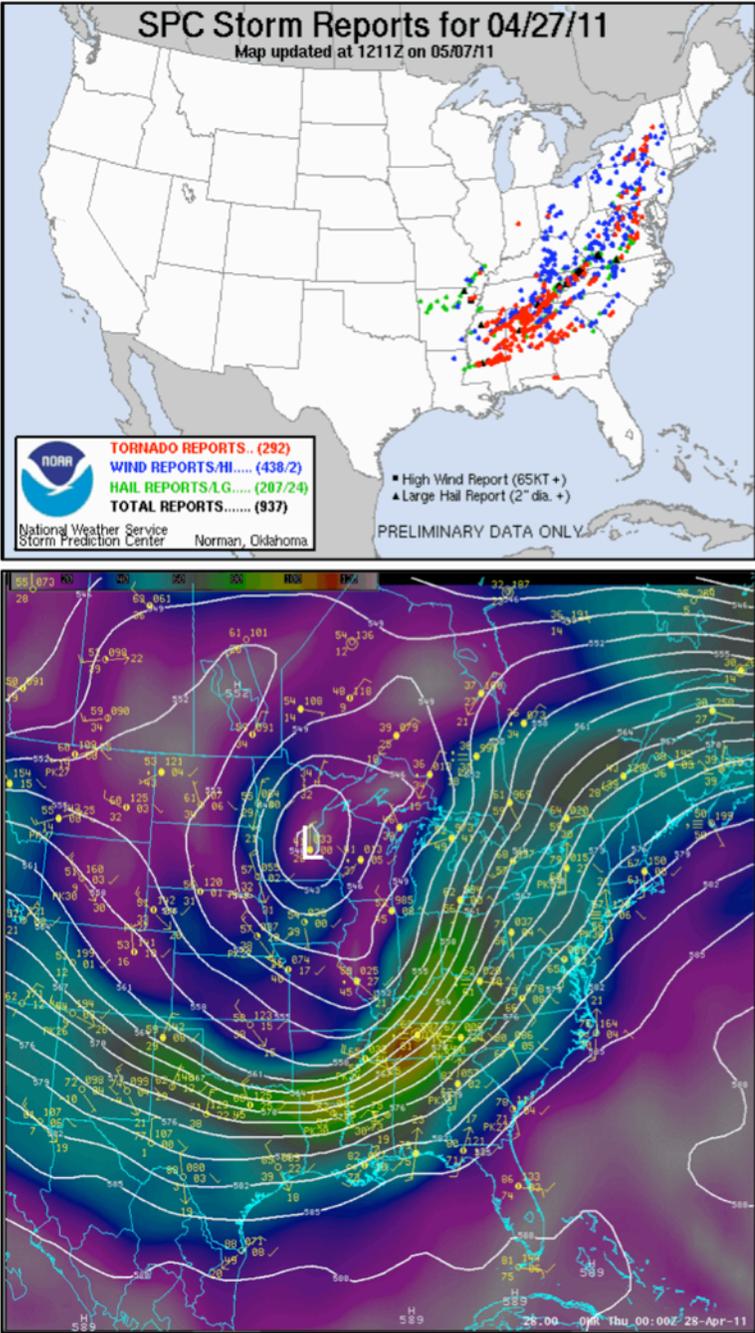
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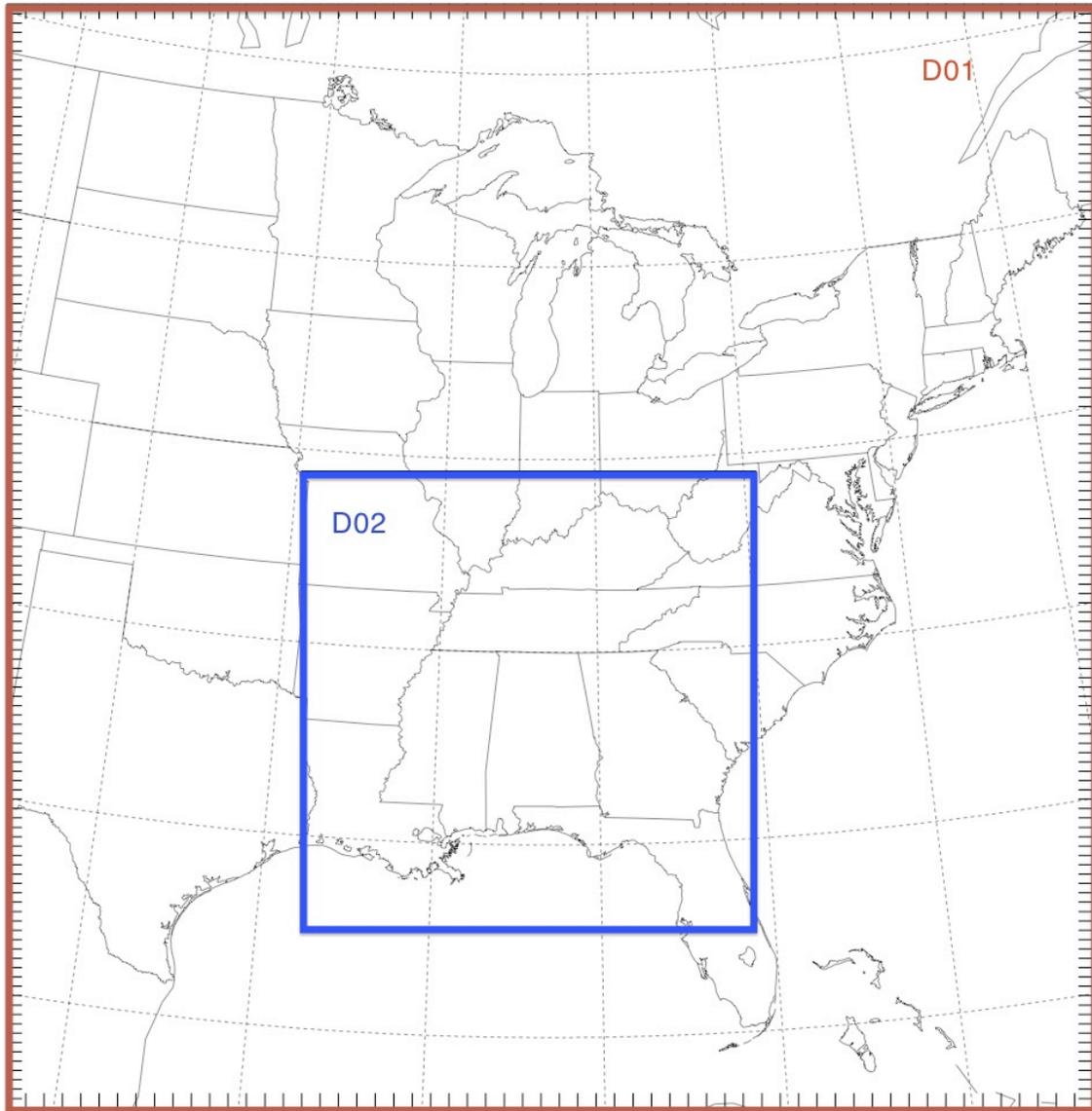
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Fig. 2. Flow chart of the data assimilation system, the left section is the MLEF system with all its components. The lightning observation operator algorithm is shown on the right-hand side of the flow chart.

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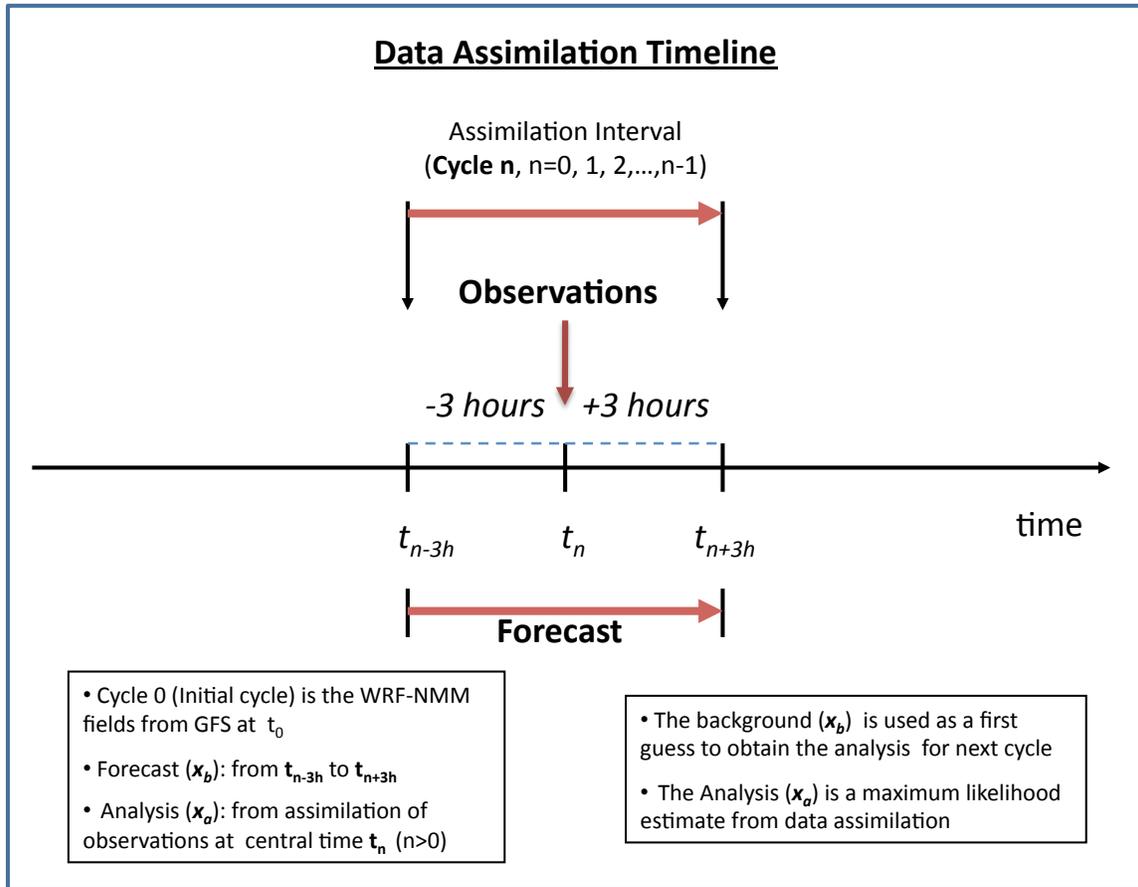


E-GRID E WE = 60. E SN = 120. DX = 0.2428. DY = 0.2428. REF LAT = 37.500. REF LON = -86.500

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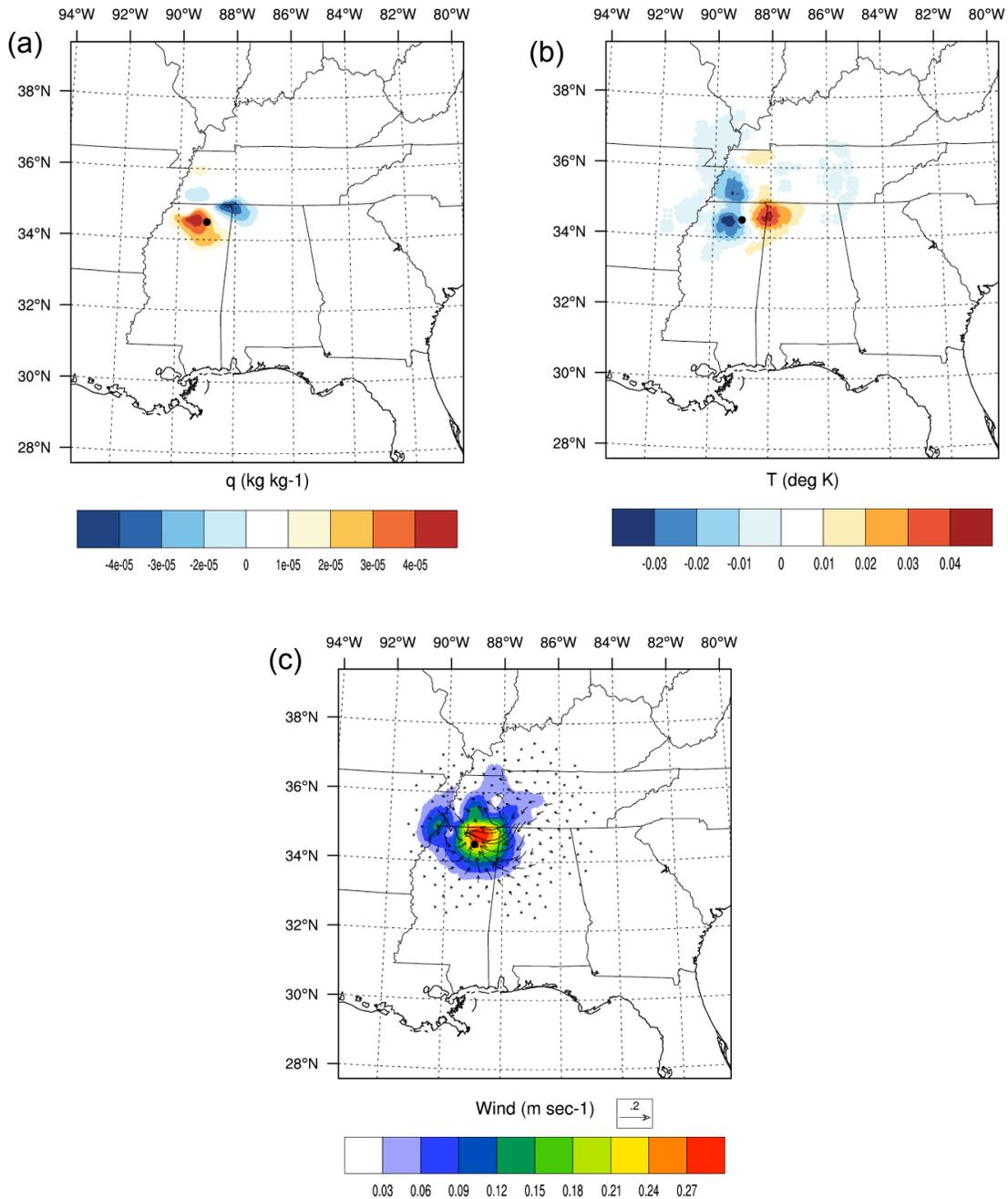
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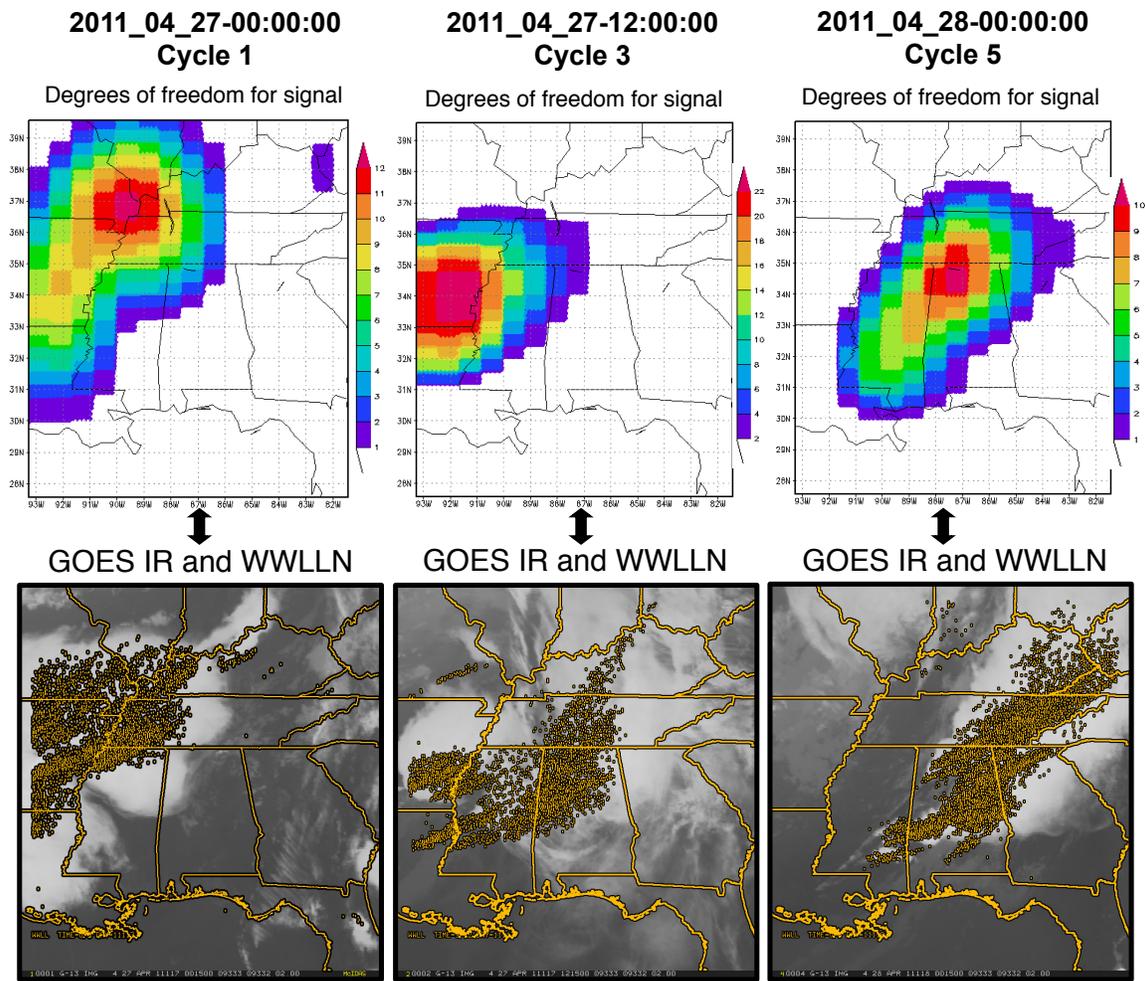


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Fig. 5. Data assimilation timeline, the data assimilation frequency for the lightning observation is 6 hours (± 3 hours) from a central time $t_n > 0$. The initial cycle (Cycle 0) is just the model (WRF-NMM) output fields from the GFS files, at t_n , the forecast, or background state (x_b) is obtained from t_{n-3h} to t_{n+3h} . The forecast is used as a guess to obtain the analysis solution for the next cycle.



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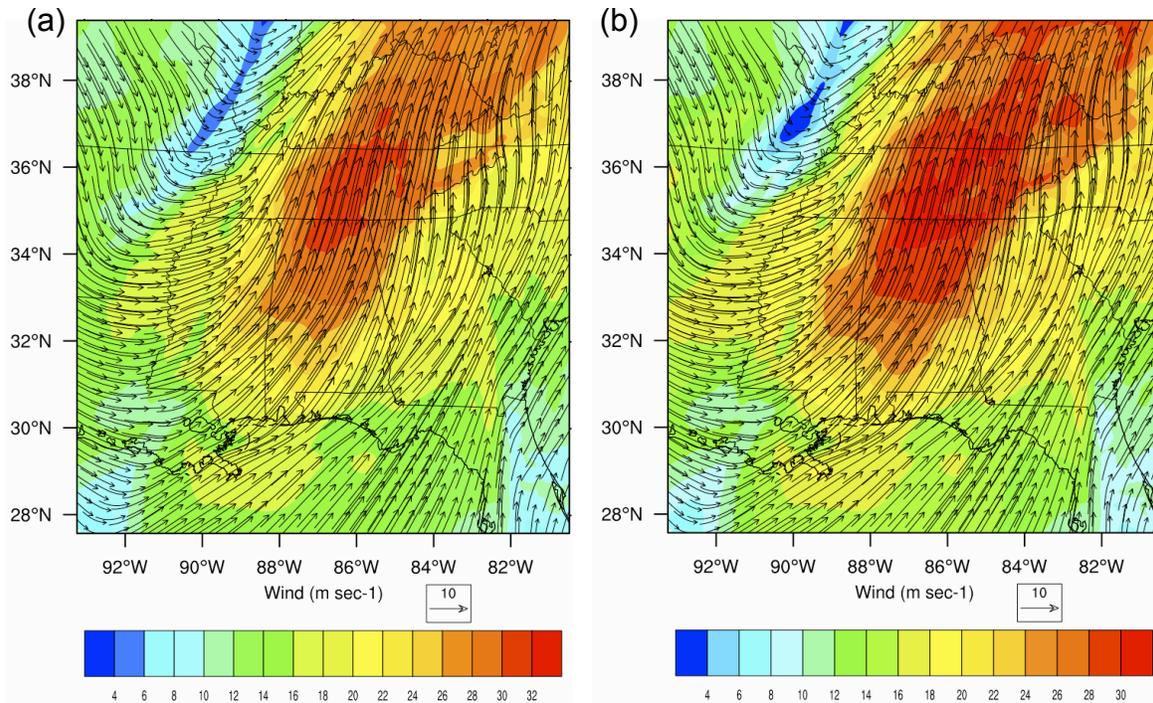
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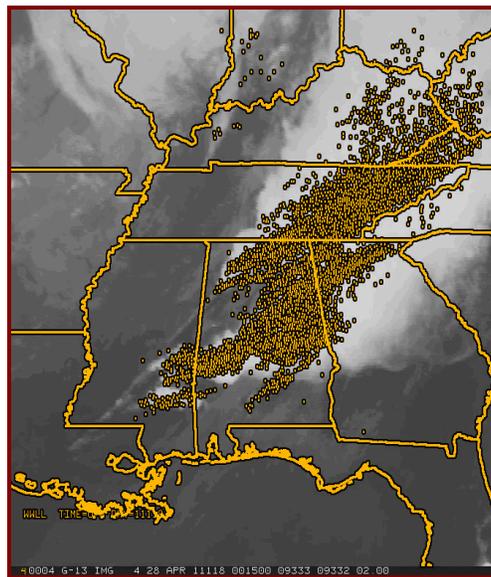
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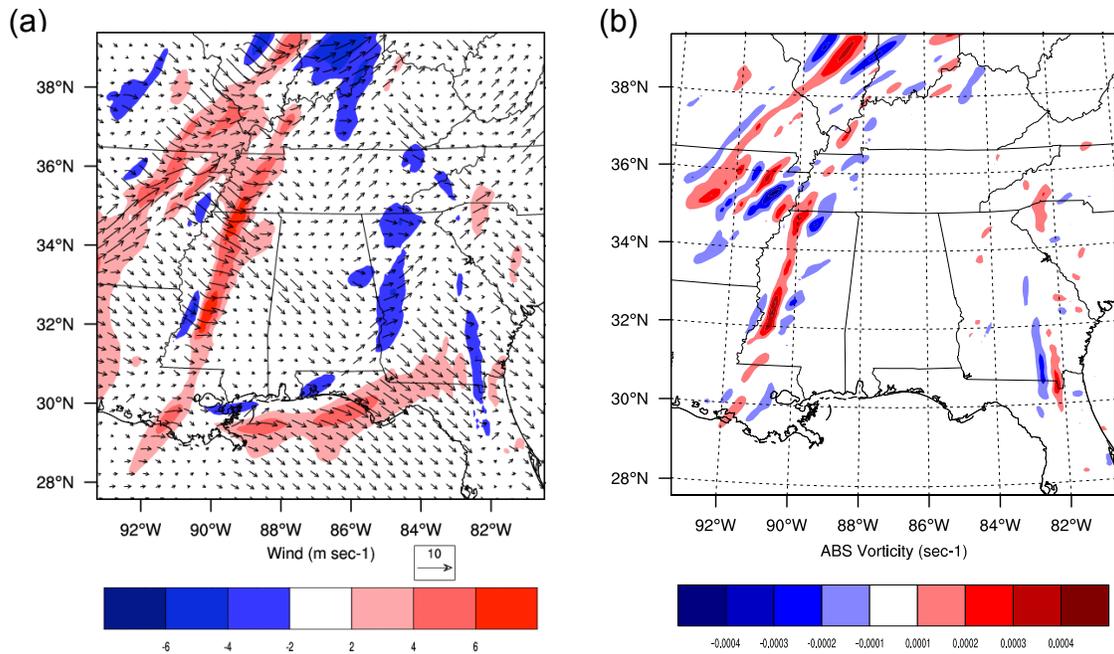
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(c)



973 **Fig. 8.** (a) Background (forecast) winds at 850 hPa at 0000 UTC 28April 2011 (cycle 5)
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 975 hPa at 0000 UTC 28April 2011 (cycle 5) from the lightning data assimilation experiment
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984 **Fig. 9.** Analysis increments at 850 hPa of (a) wind and (b) absolute vorticity at
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 986 left-hand side in both plots indicated by the ellipses. Winds are being advected
 987 into the region of strong CAPE seen in Fig. 10a.
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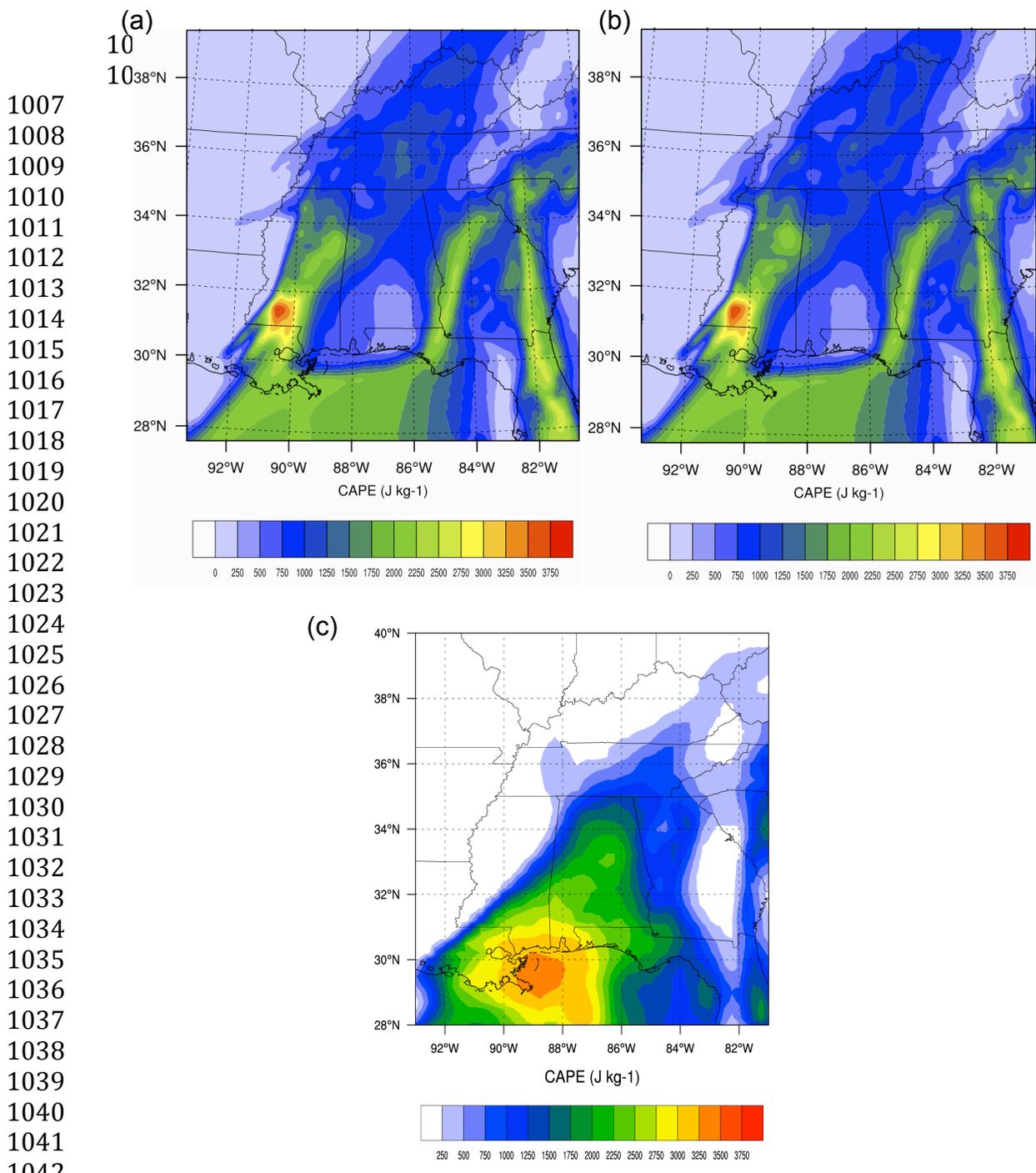
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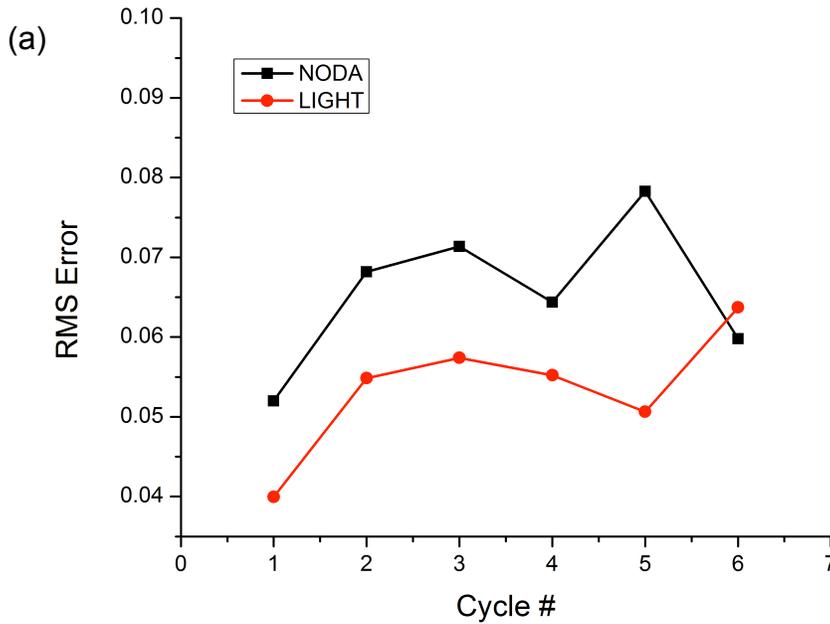
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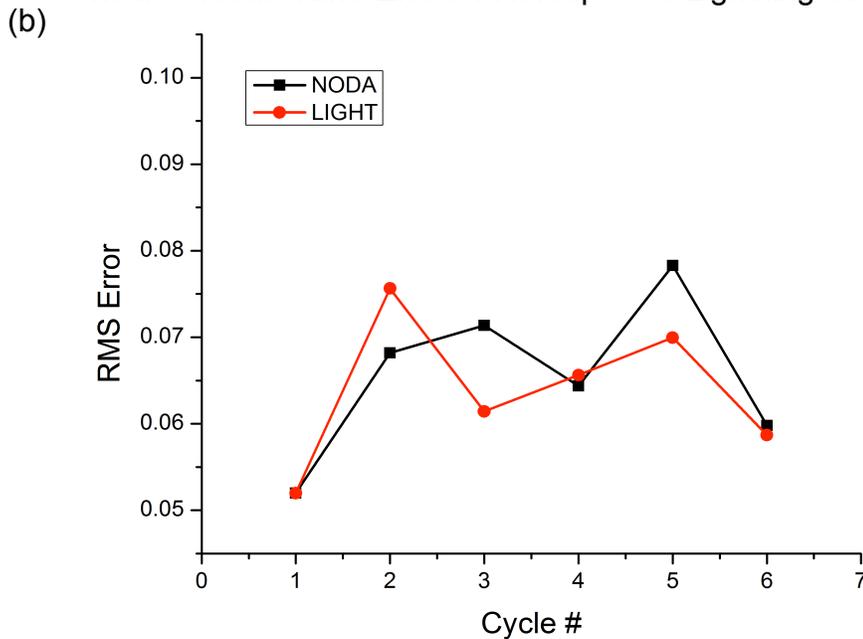
1043 **Fig. 10.** Background CAPE for (a) NODA and (b) LIGHT experiments, and (c)
1044 observed CAPE from the Storm Prediction Center's Surface Mesoanalysis at
1045 0000 UTC 28 April 2011 (cycle5). A region of high CAPE gradient is observed in
1046 the upper-left hand side of the domain, indicating the presence of a well-defined
1047 dry line, in agreement with observations, but there are no significant differences
1048 between both experiments. One reason is that there are no lightning
1049 observations in the region where the strongest CAPE was observed. Lightning
1050 data was not able to impact CAPE.

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Analysis RMS Errors with respect to Lightning Observations



6-hour Forecast RMS Errors with respect to Lightning Observations



1077 **Fig. 11.** Root mean square (RMS) errors with respect to lightning flash rate
1078 observations during six assimilation cycles at 6 h intervals: **(a)** Analysis RMS
1079 error. The RMS error reduction was achieved during the first 5 cycles of the
1080 assimilation period, while there is deterioration in the last cycle, possibly due to
1081 the fact that the system was exiting the model domain. **(b)** 6 h forecast RMS
1082 error. There is no clear improvement in the forecast, suggesting that additional
1083 development of the assimilation system might be required, such as an

1084 improvement of the observation operator, adding new observations, and possibly
1085 improving the forecast uncertainty estimation.
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1087 **Figure Captions**

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1089 **Fig. 1.** Statistics of normalized innovation vectors $R^{-1/2} [y - h(x^f)]$, or PDF
1090 innovations for cycles 1-5 for both domains (D01 and D02) before (left-blue) and
1091 after (right-red) correction. The skewed histograms on the left implicitly indicate
1092 that the values of observed lightning flash rate are considerably larger than the
1093 guess, a situation that required a correction.
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1095 **Fig. 2.** Flow chart of the data assimilation system, the left section is the MLEF
1096 system with all its components. The lightning observation operator algorithm is
1097 shown on the right-hand side of the flow chart.
1098

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